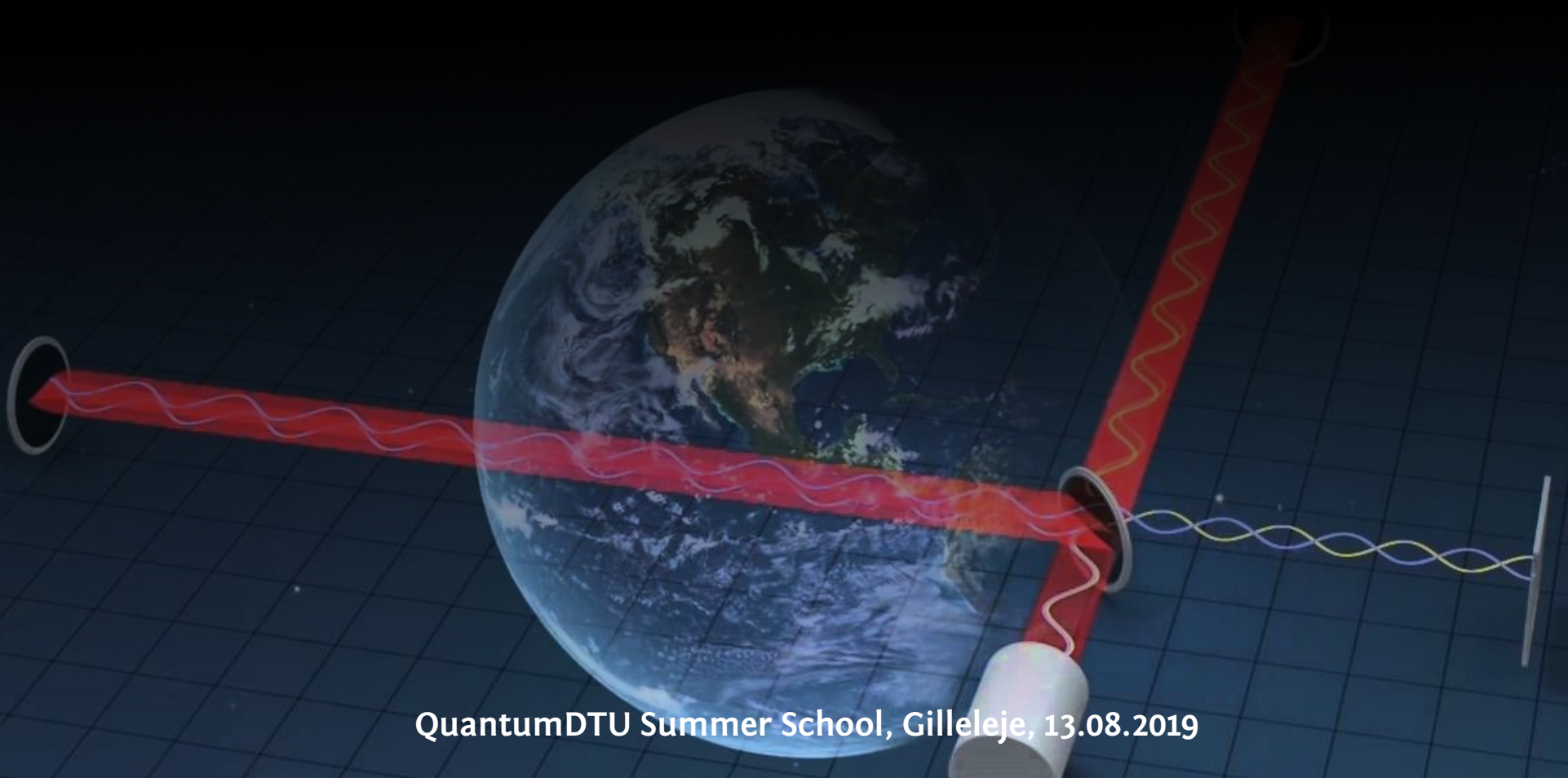


Noisy Quantum Metrology

Jonatan Bohr Brask
DTU Physics

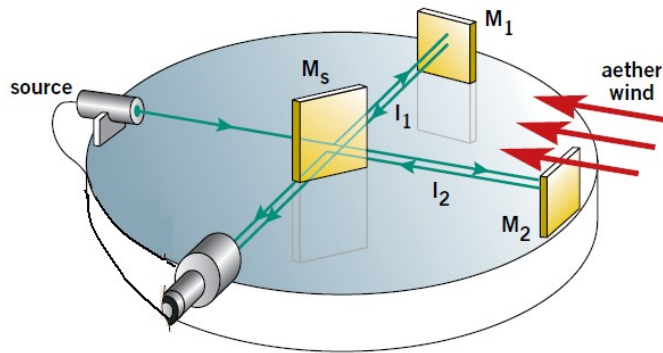


Why precision measurements?

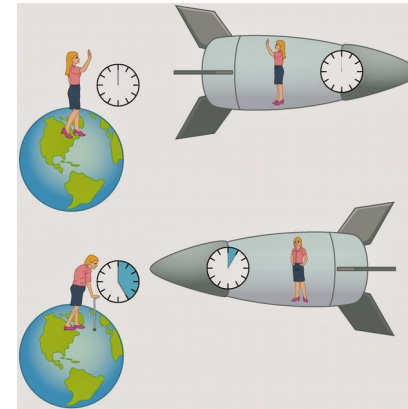
- Triggers science breakthroughs.

Why precision measurements?

- Triggers science breakthroughs.



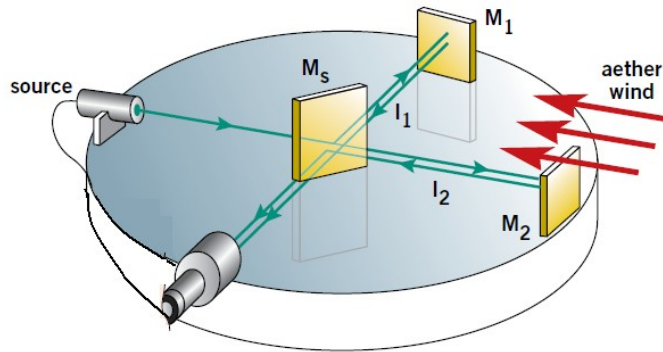
Michelson-Morley experiment, 1887



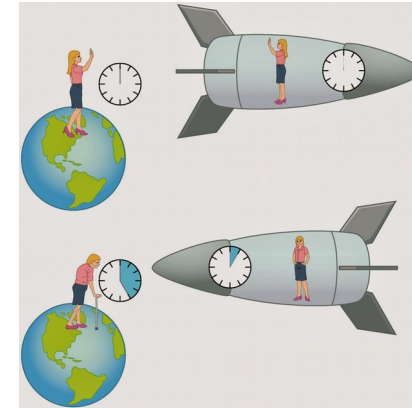
Special relativity

Why precision measurements?

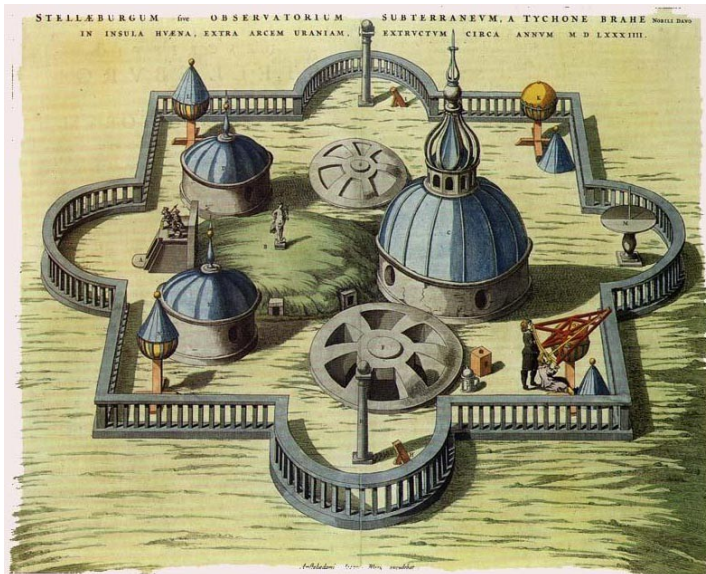
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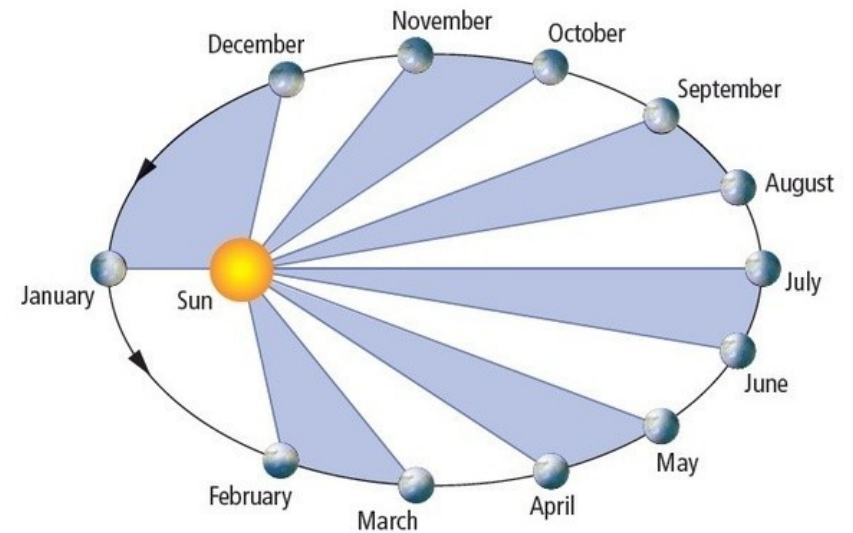
Michelson-Morley experiment, 1887



Special relativity



Tycho Brahe astronomical observations, late 1500s



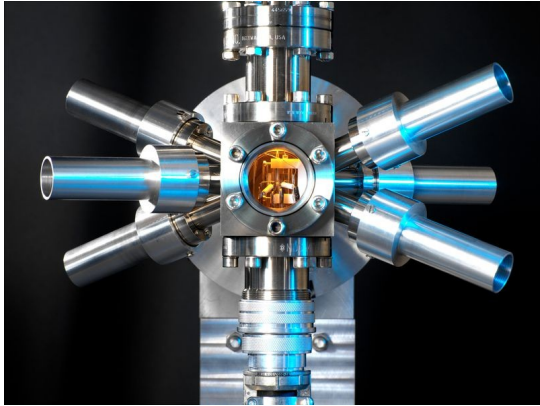
Kepler's laws, changing heavens

Why precision measurements?

- Enables technology

Why precision measurements?

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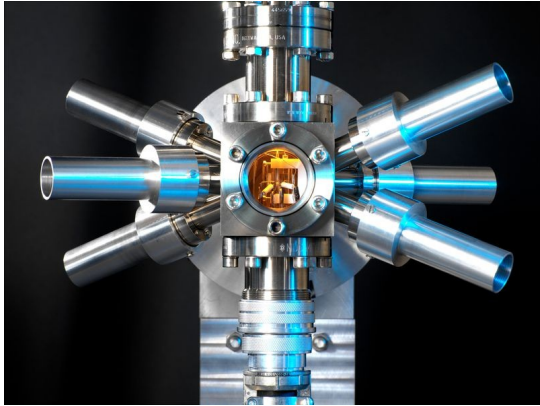
Precise time keeping (atomic clocks)



GPS

Why precision measurements?

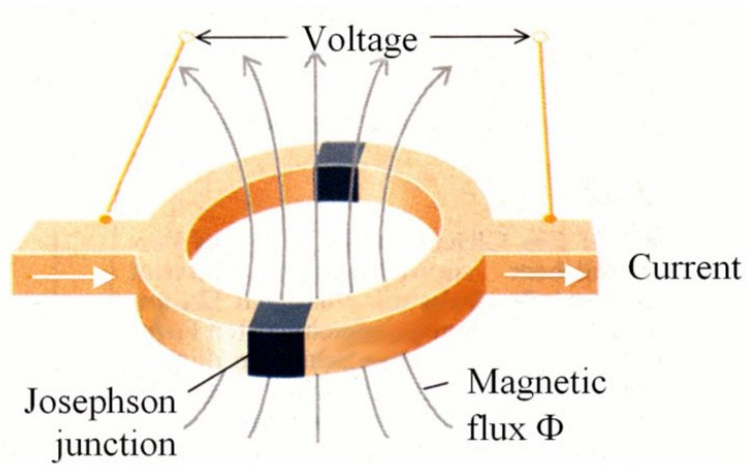
- Enables technology



Precise time keeping (atomic clocks)



GPS



Precise magnetometry (SQUIDs)



Magnetoencephalography

Why quantum?

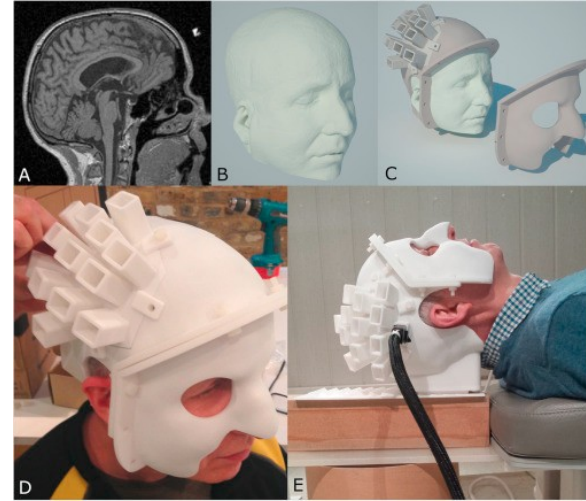
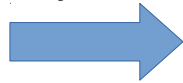
- Better precision without increasing probe size / measurement time.

Why quantum?

- Better precision without increasing probe size / measurement time.



Atomic vapor magnetometry /
Nitrogen-vacancy centers in diamond



Boto et al., NeuroImage, 149, 404 (2017)

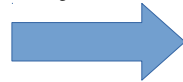
- Room temperature.
- Reduce sensor size w. quantum effects.

Why quantum?

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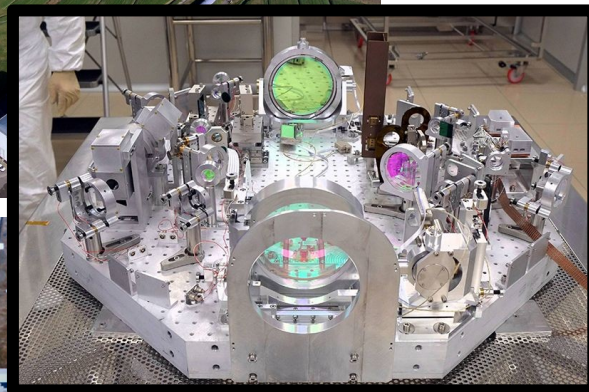


Atomic vapor magnetometry /
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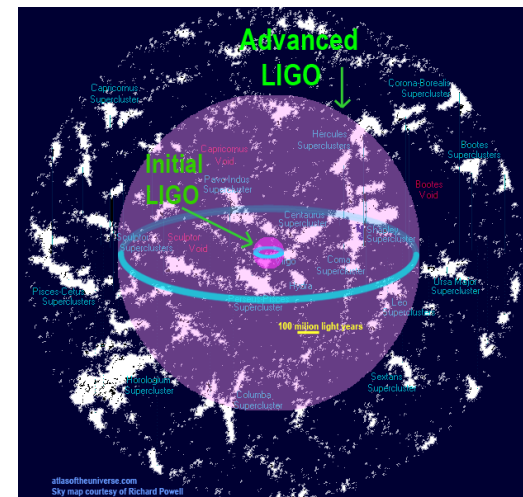


Boto *et al.*, *NeuroImage*, 149, 404 (2017)

- Room temperature.
- Reduce sensor size w. quantum effects.



Squeezed light



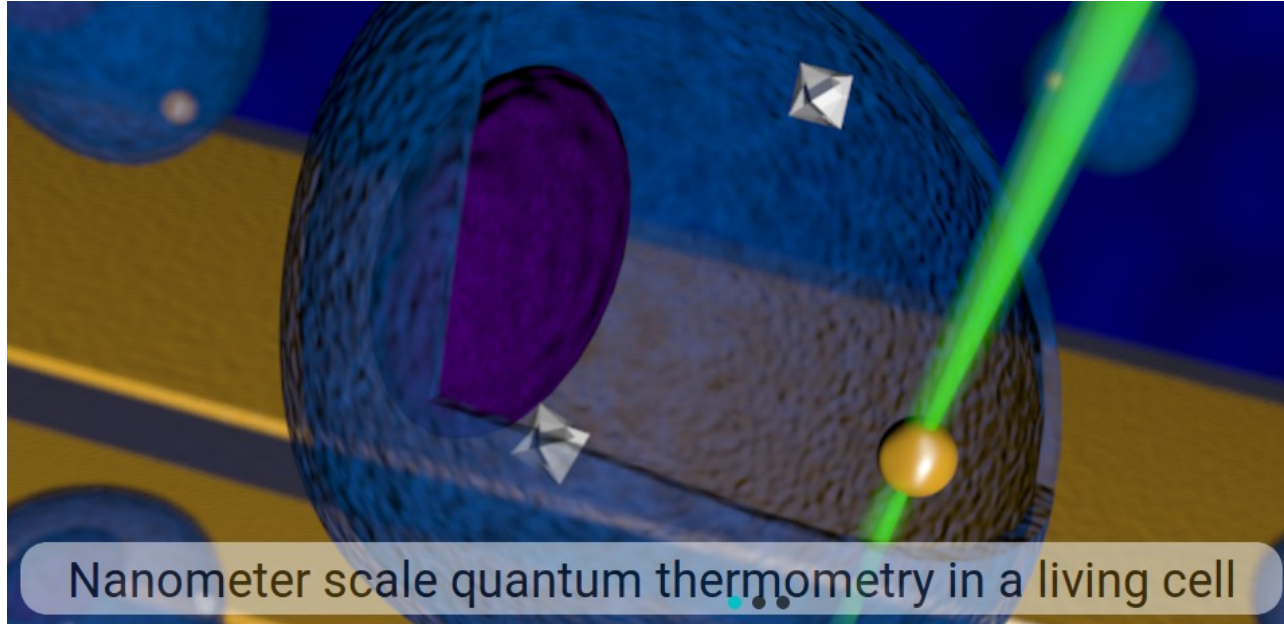
- Better sensitivity w/o increasing light power.

Why noisy?

Because any real system is...

Need to demonstrate that quantum techniques give an advantage under realistic conditions.

We would like to apply quantum metrology not just for well controlled lab physics but also for applications in more noisy environments.



Kucsko *et al.*, Nature 500, 54 (2013).

Outline

Noiseless quantum metrology

- Parameter estimation
- Precision scaling
- Fisher information
- Standard quantum limit
- Heisenberg limit

Noisy quantum metrology

- No-go results
- Directional noise
- Quantum error correction

Parameter estimation - examples

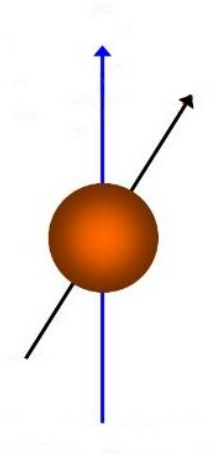
Magnetometry

B

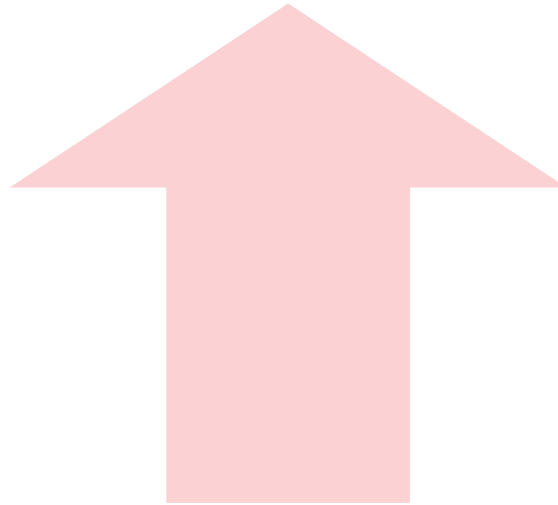


Parameter estimation - examples

Magnetometry

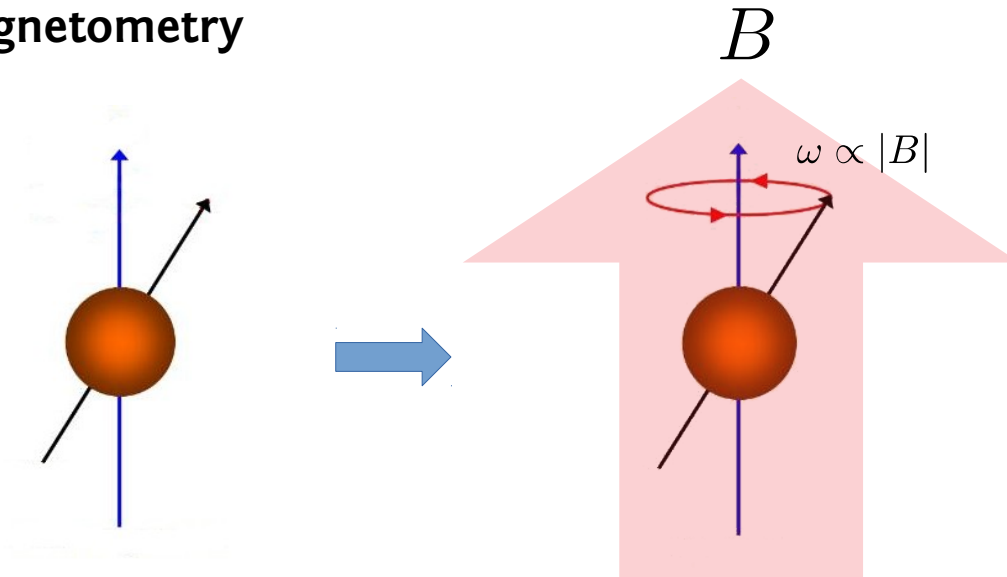


B



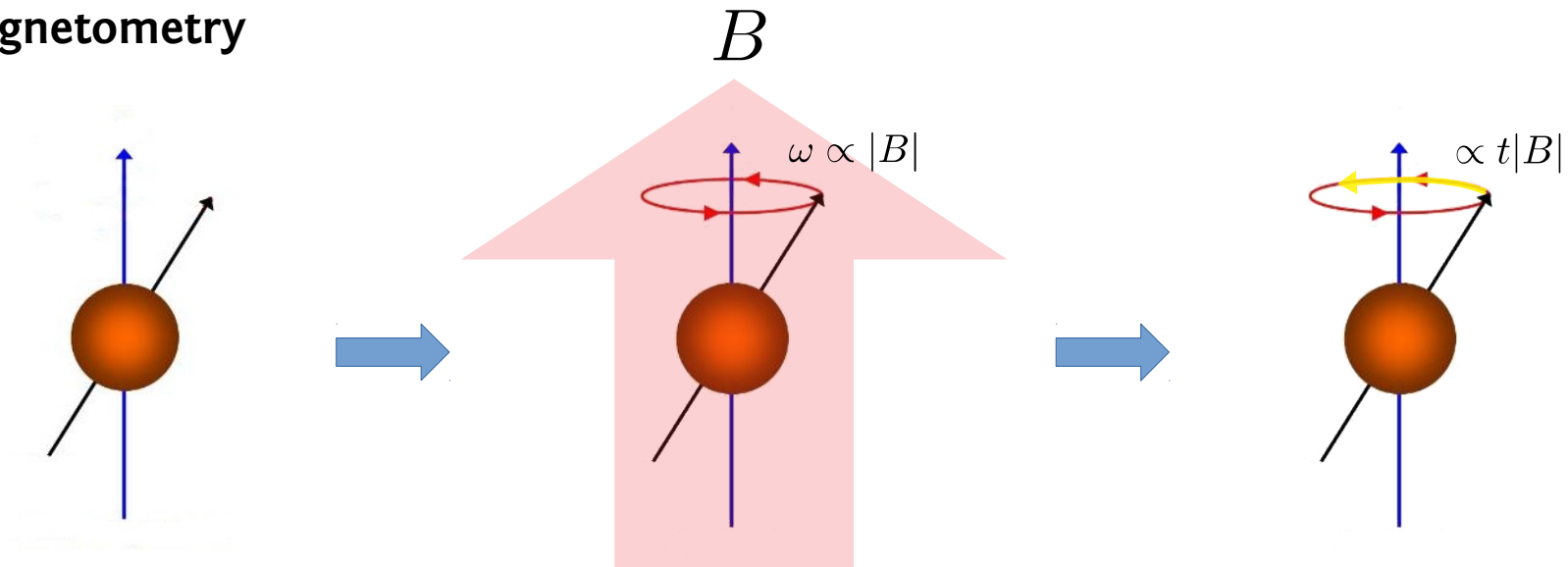
Parameter estimation - examples

Magnetometry



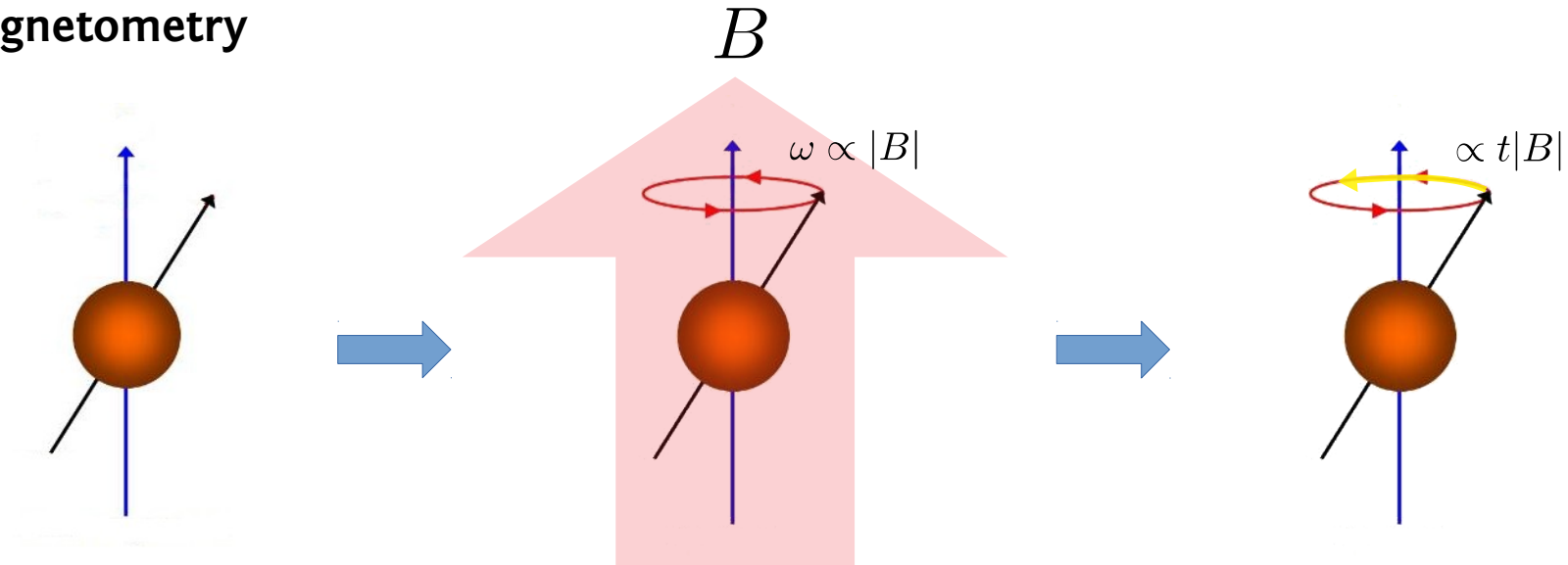
Parameter estimation - examples

Magnetometry

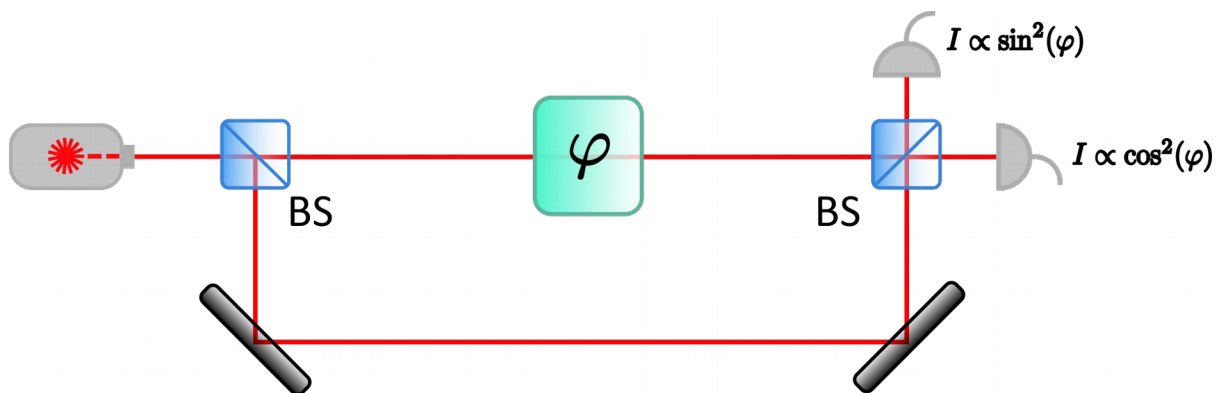


Parameter estimation - examples

Magnetometry

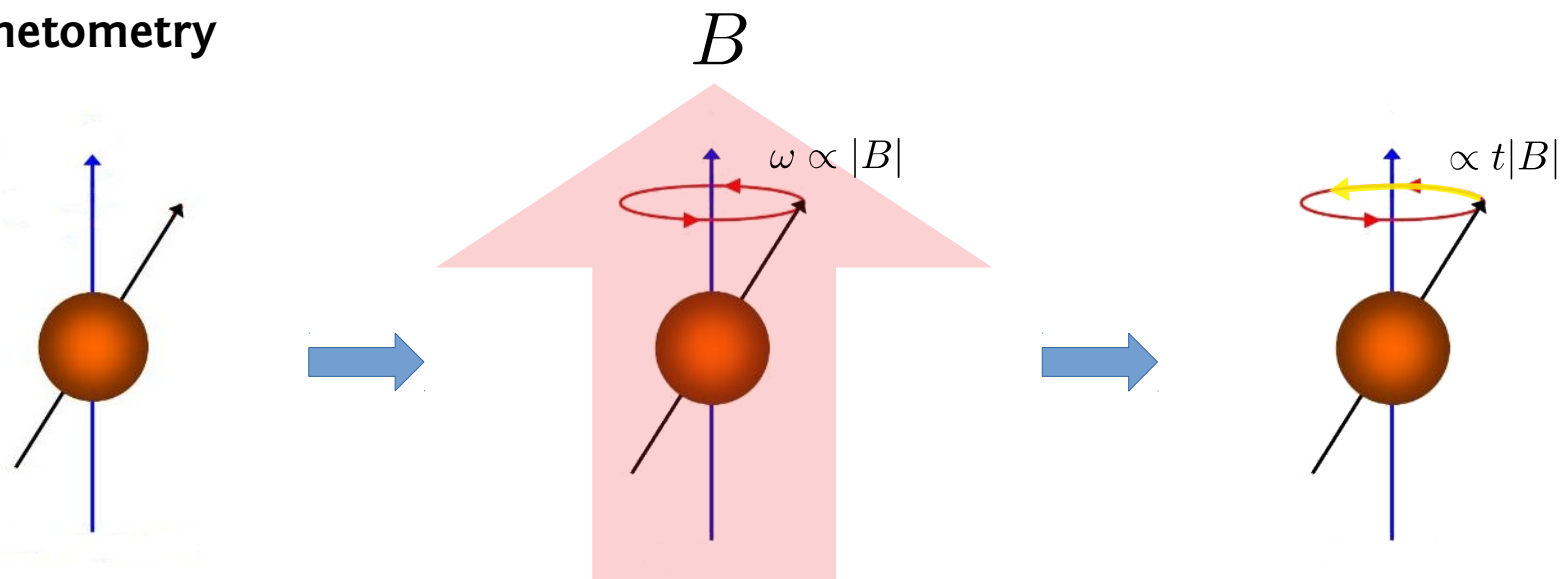


Optical phase estimation

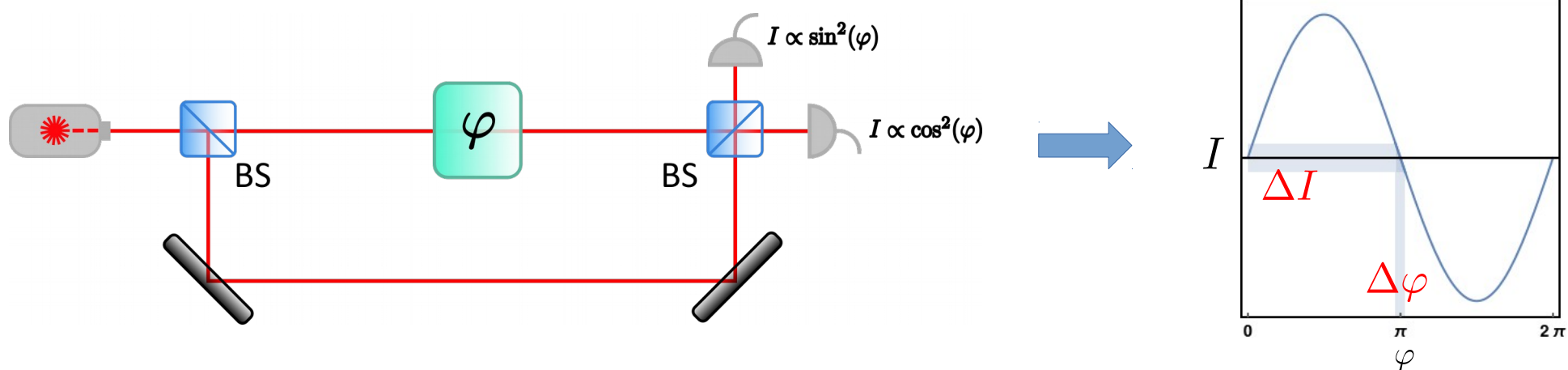


Parameter estimation - examples

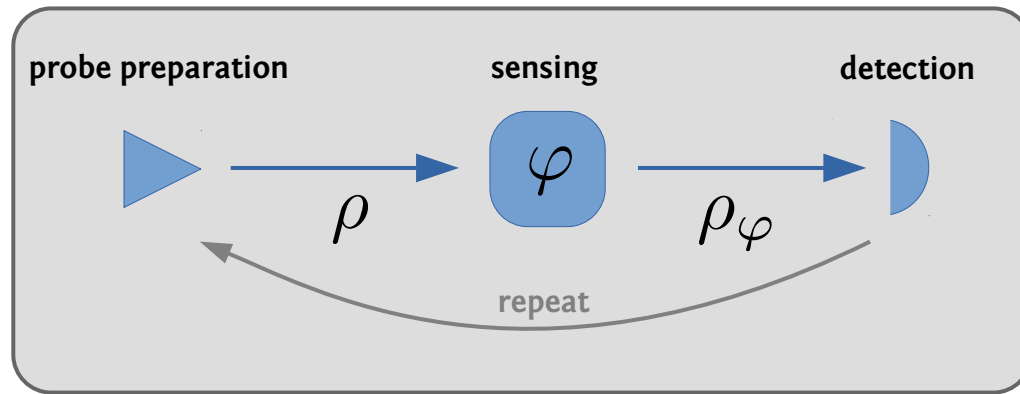
Magnetometry



Optical phase estimation

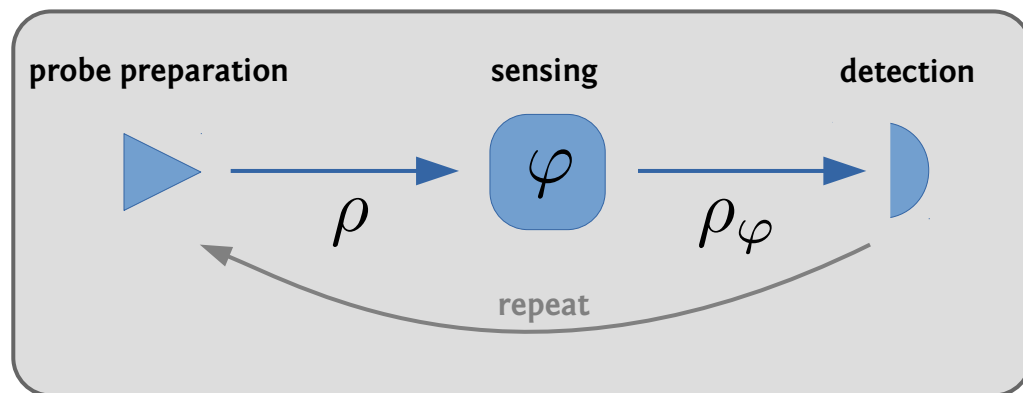


Single-parameter estimation

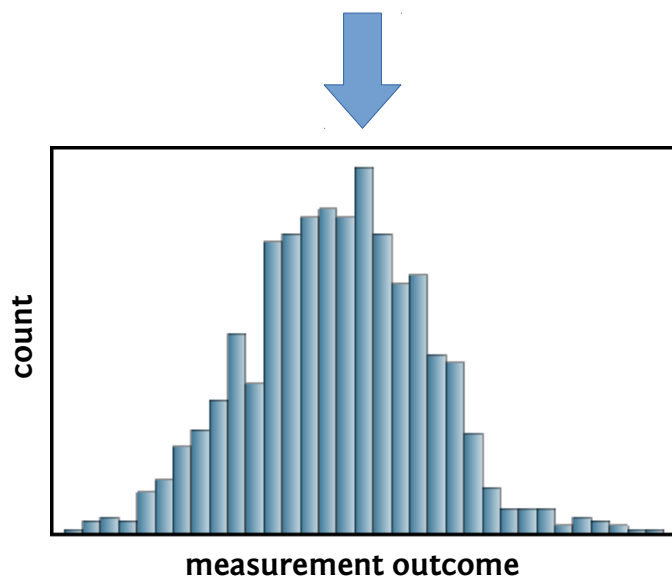


Measurements.

Single-parameter estimation

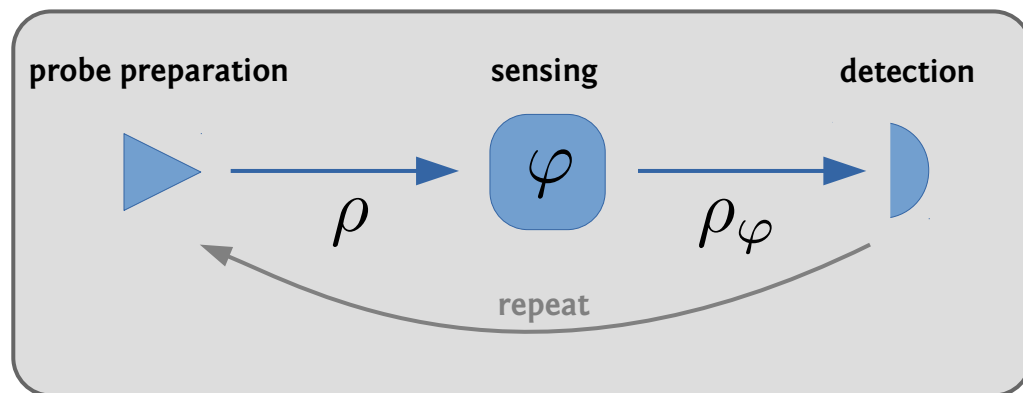


Measurements.

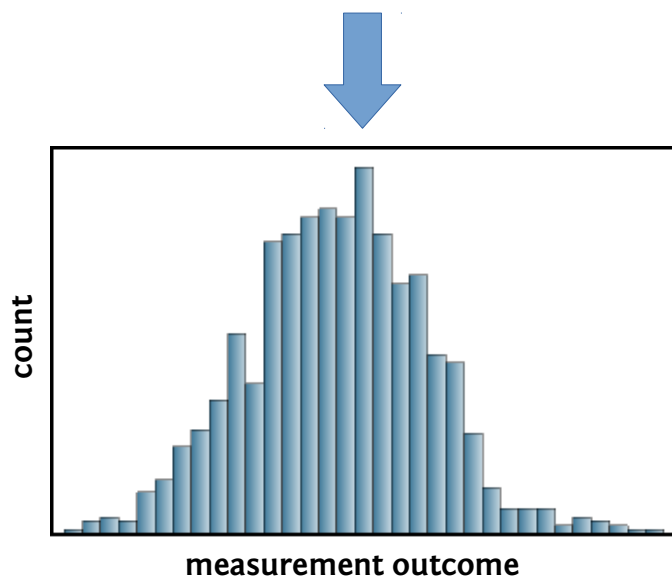


Raw data.

Single-parameter estimation



Measurements.



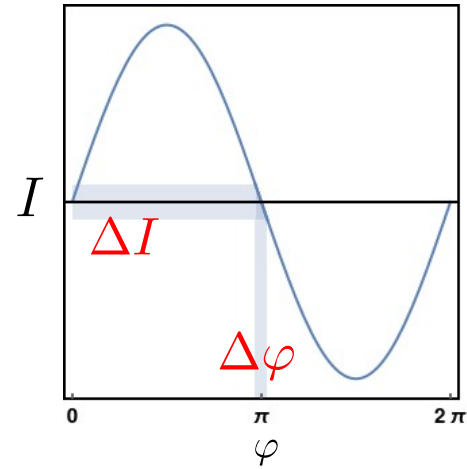
Raw data.

$$\hat{\varphi} = f(x_0, x_1 \dots)$$

Estimation (post processing).

Focus on local estimation

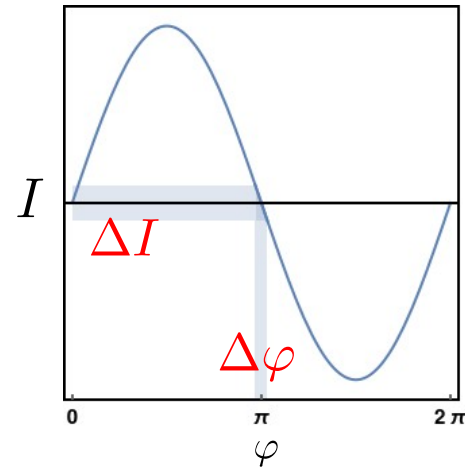
- Sensitivity to small changes.
- Asymptotic scaling of the achievable precision.



ν

Focus on local estimation

- Sensitivity to small changes.
- Asymptotic scaling of the achievable precision.



Classical vs. Quantum

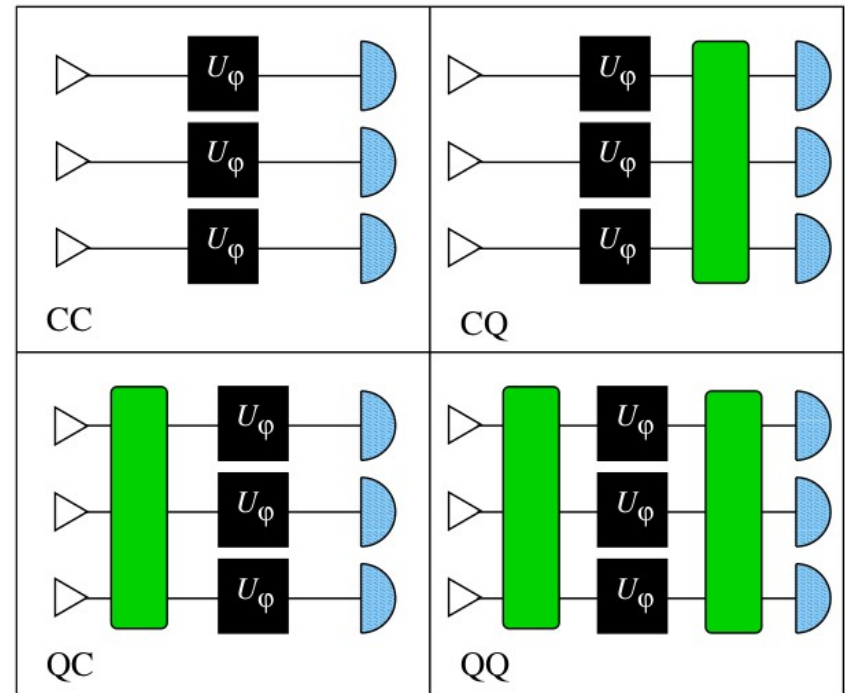
- Compare precision for fixed resources:

N size of probe used.

T total measurement time.

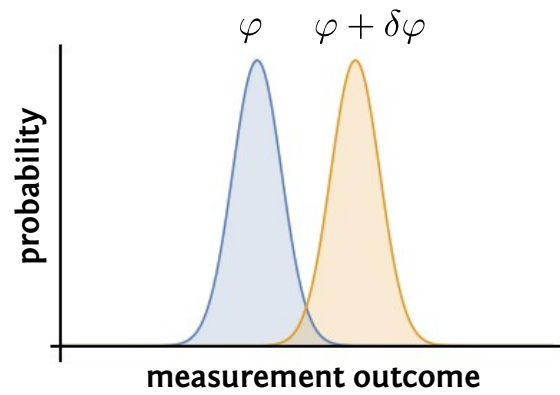
ν number of repetitions.

$$\Delta^2 \varphi \sim ?(N, T, \nu)$$

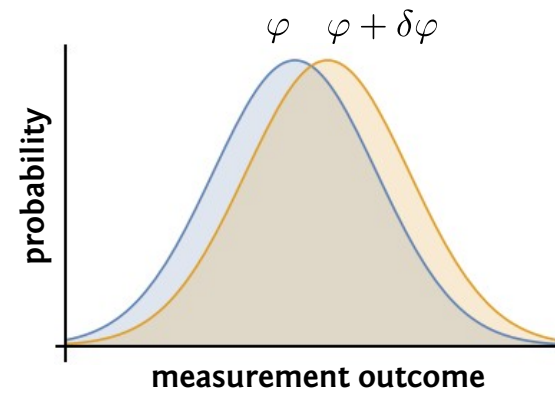


Quantifying sensitivity - Fisher information

very sensitive...

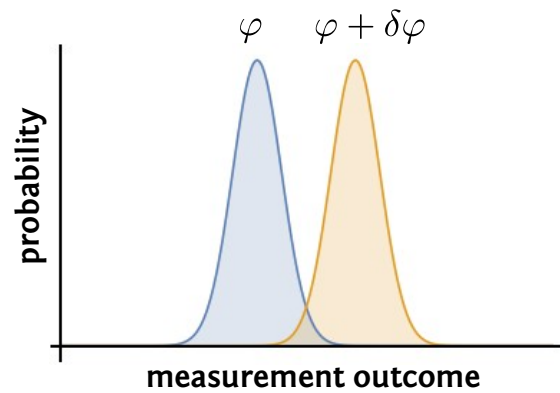


not very sensitive...

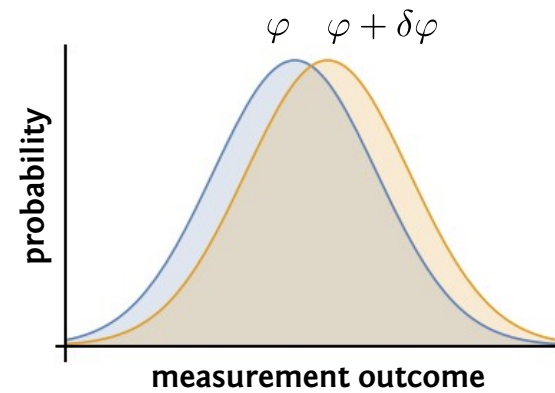


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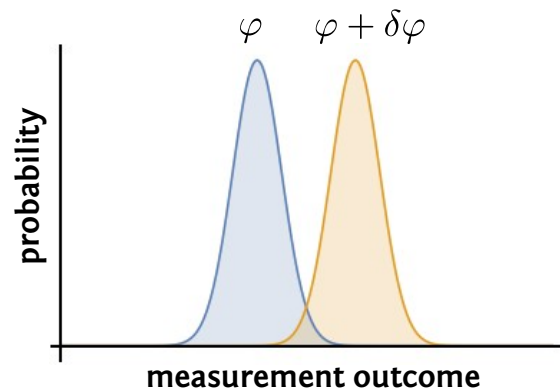


Fisher information - measures information about the parameter in the distribution

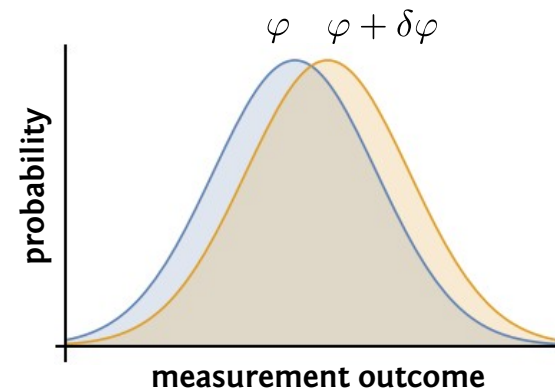
$$\mathcal{F}_\varphi = \mathcal{F}(p_\varphi) = \sum_x \left(\frac{\partial}{\partial \varphi} \log p_\varphi(x) \right)^2 p_\varphi(x)$$

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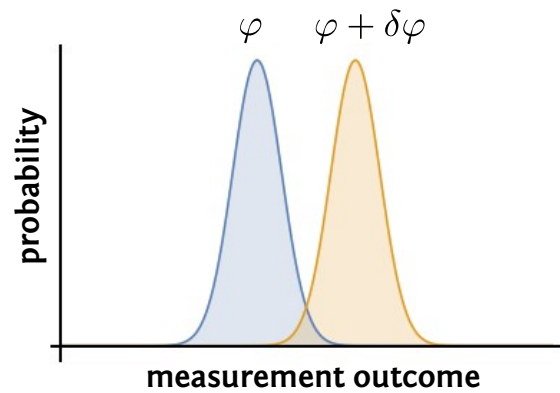


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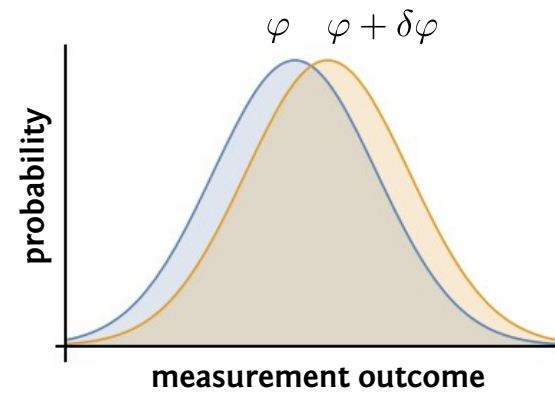
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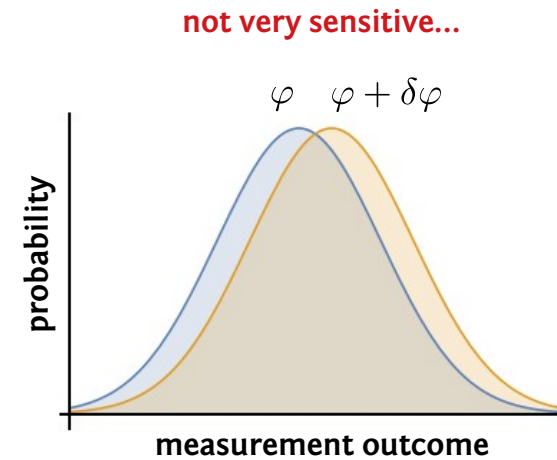
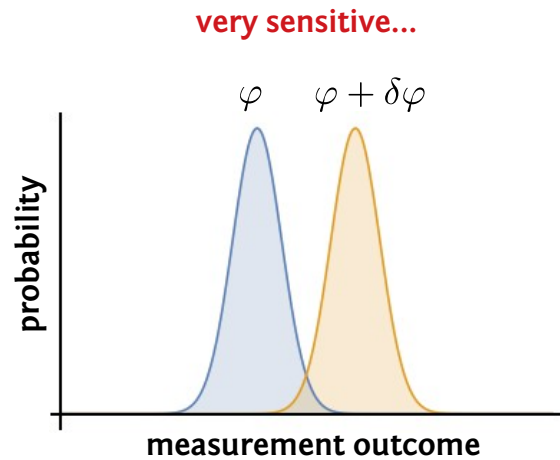
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Bounding estimation error - Cramér-Rao bound

$$\Delta^2 \hat{\varphi} \geq \frac{1}{\mathcal{F}_\varphi}$$

- Valid for any unbiased estimator.
- Can be saturated in the limit of many repetitions (and often before).



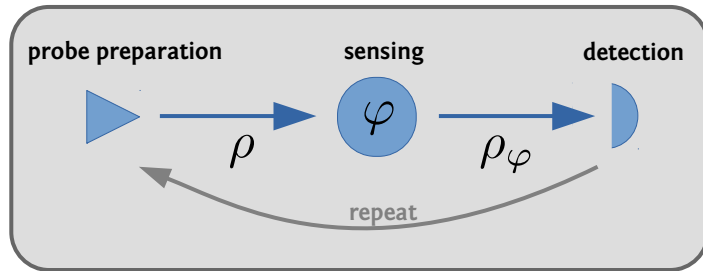
Harald Cramér



C. R. Rao

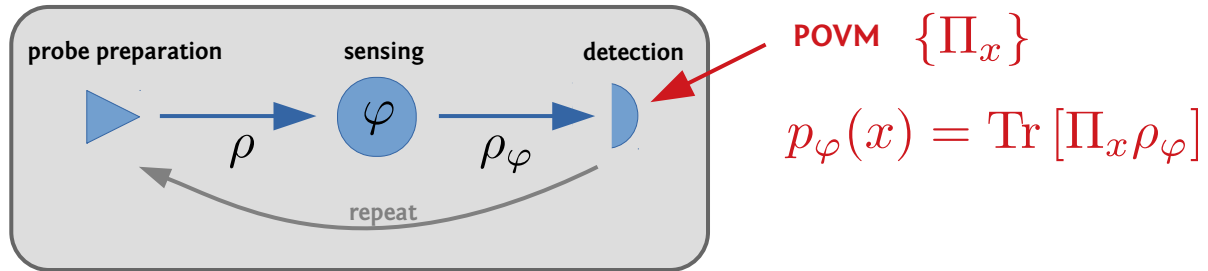
Quantum Fisher information

The Fisher information depends on the state and measurement.



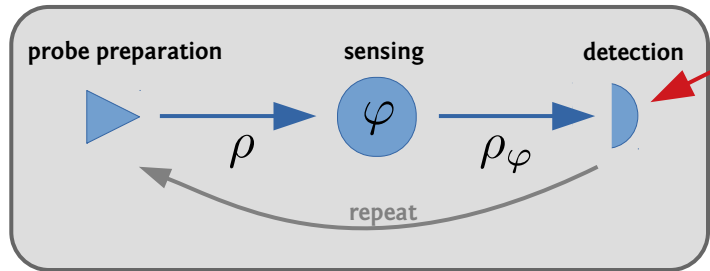
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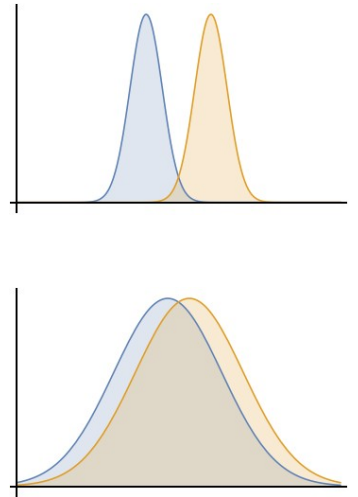


POVM $\{\Pi_x\}$

$$p_\varphi(x) = \text{Tr} [\Pi_x \rho_\varphi]$$

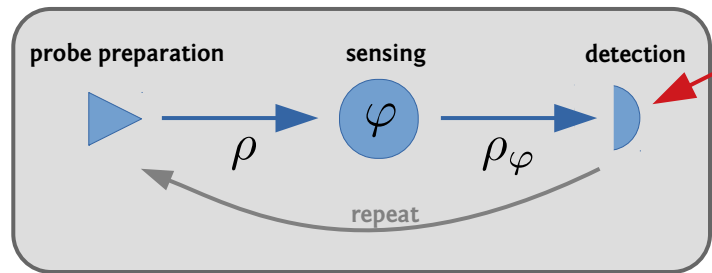
$\{\Pi_x\}$

$\{\Pi'_x\}$



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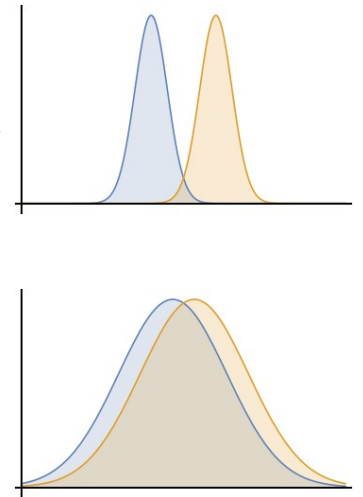


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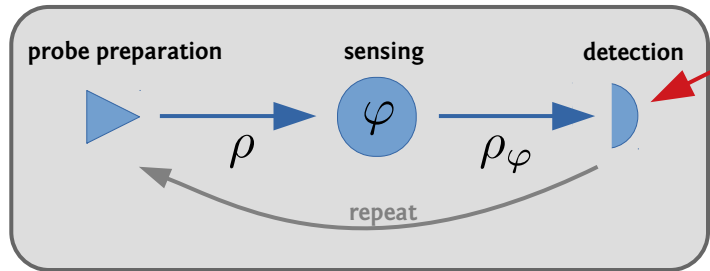


The *Quantum Fisher Information (QFI)* corresponds to the best possible measurement

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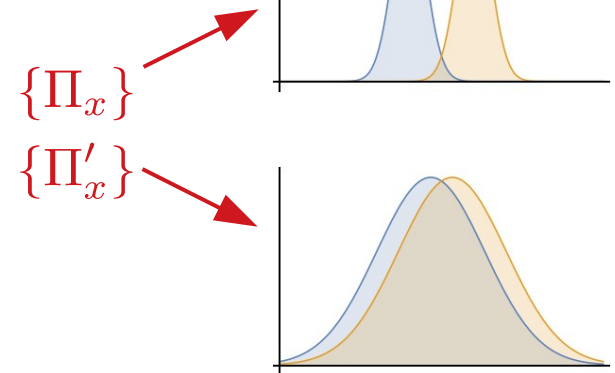
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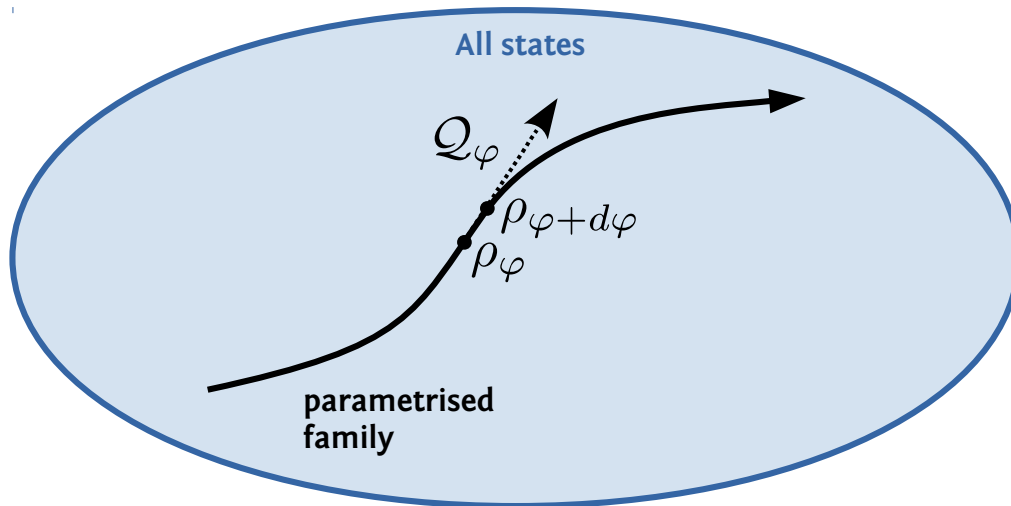
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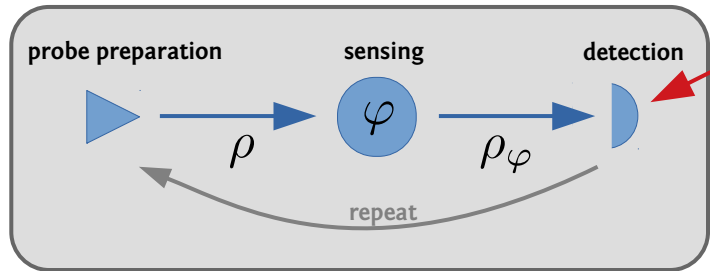
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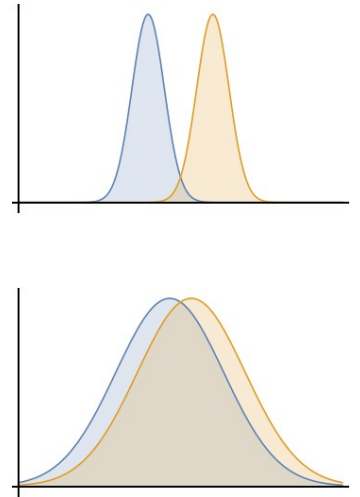


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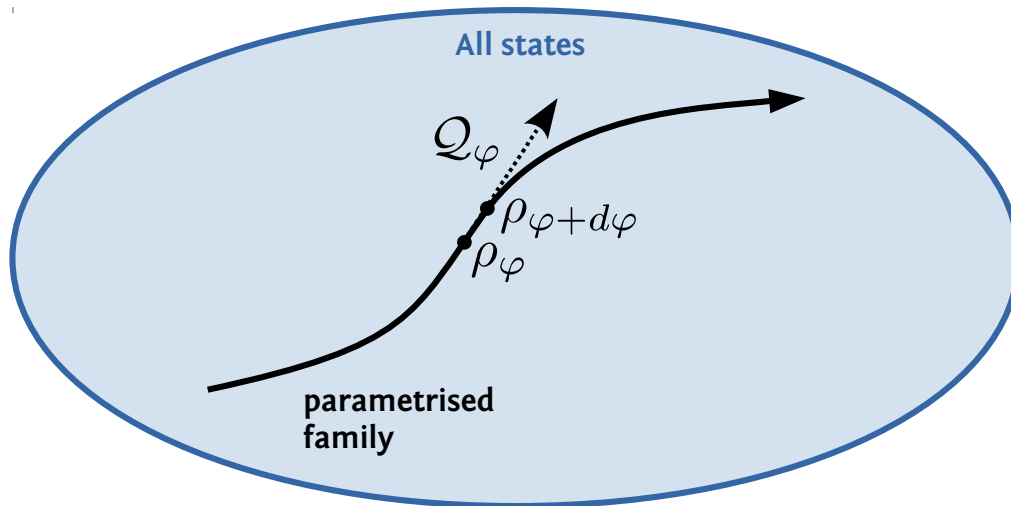
$\{\Pi_x\}$

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The *Quantum Fisher Information (QFI)* corresponds to the best possible measurement

$$\mathcal{Q}_\varphi = \mathcal{Q}(\rho_\varphi) = \max_{\{\Pi_x\}} \mathcal{F}(p_\varphi)$$



Quantum Cramér-Rao bound

$$\Delta^2 \hat{\varphi} \geq \frac{1}{\mathcal{Q}_\varphi}$$

can be saturated if the optimal measurement can be implemented.

Standard Quantum Limit (SQL)

The QFI is additive...

$$\mathcal{Q}(\rho_{\varphi}^{\otimes N}) = N \mathcal{Q}(\rho_{\varphi})$$

information from independent trials simply adds

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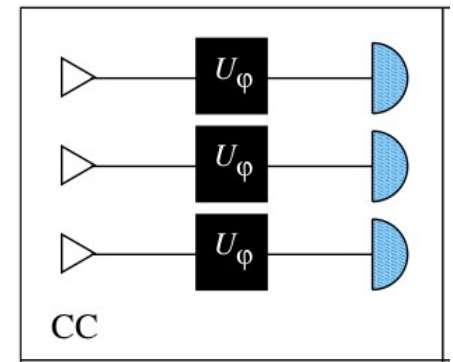
parameter-independent mixing can only make things worse



Classically, the error scales linearly with probe size

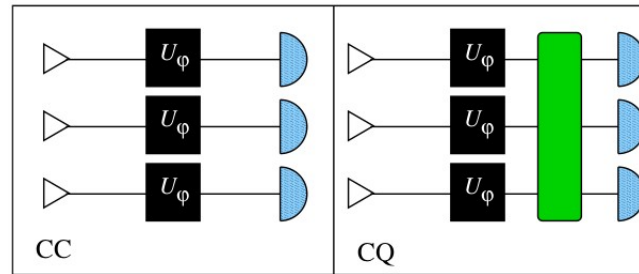
$$\Delta^2_{\varphi} \geq \frac{1}{N \mathcal{Q}_{\varphi}} \quad (\text{w/o entanglement})$$

This linear scaling with N is the SQL.



Can we beat the SQL?

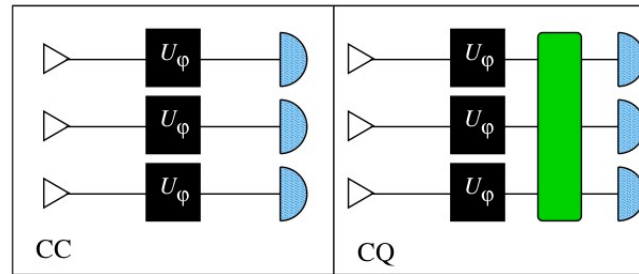
- For classical input, collective measurements do not help.



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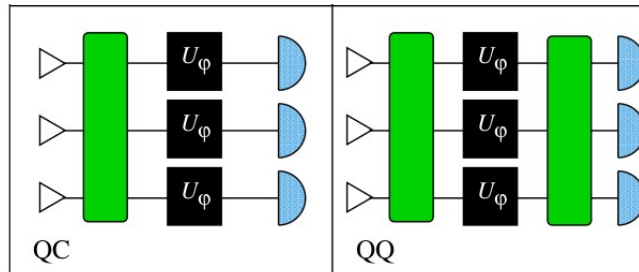
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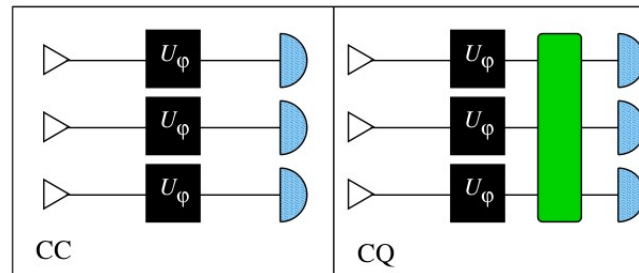
- Entangled states *do improve precision*. collective measurements are not required. xxx



$$\Delta^2 \varphi < \frac{1}{N Q_\varphi}$$

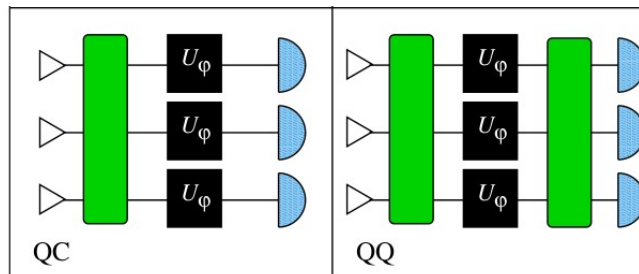
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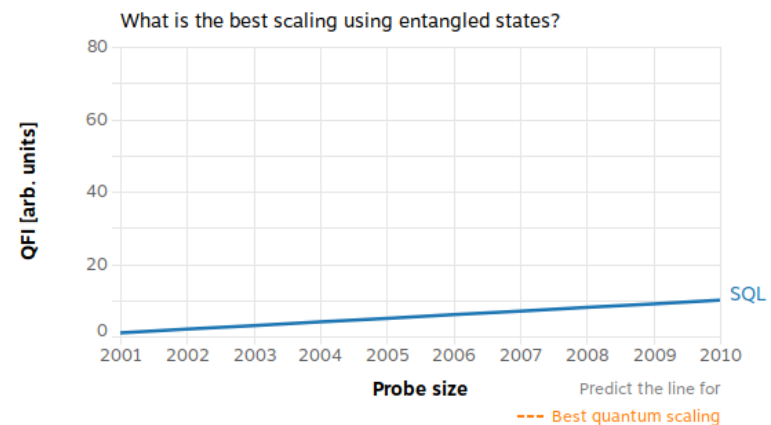
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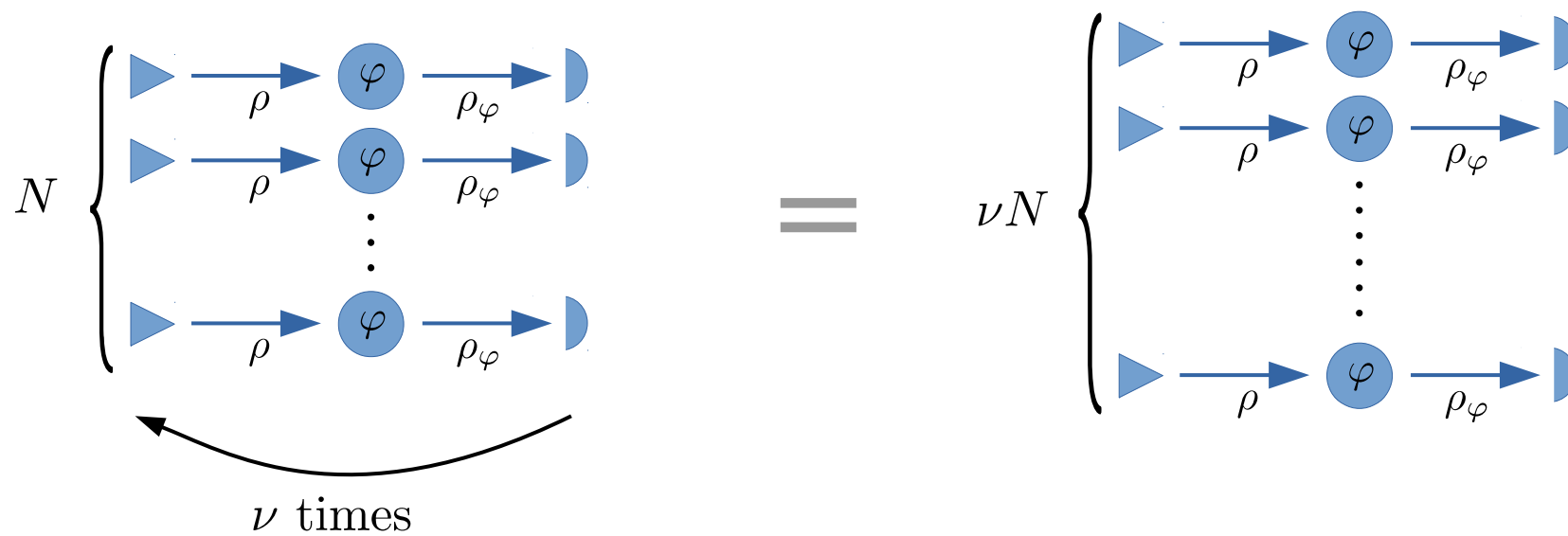
$$\Delta^2 \varphi < \frac{1}{N Q_\varphi}$$

What do you think the best scaling is?

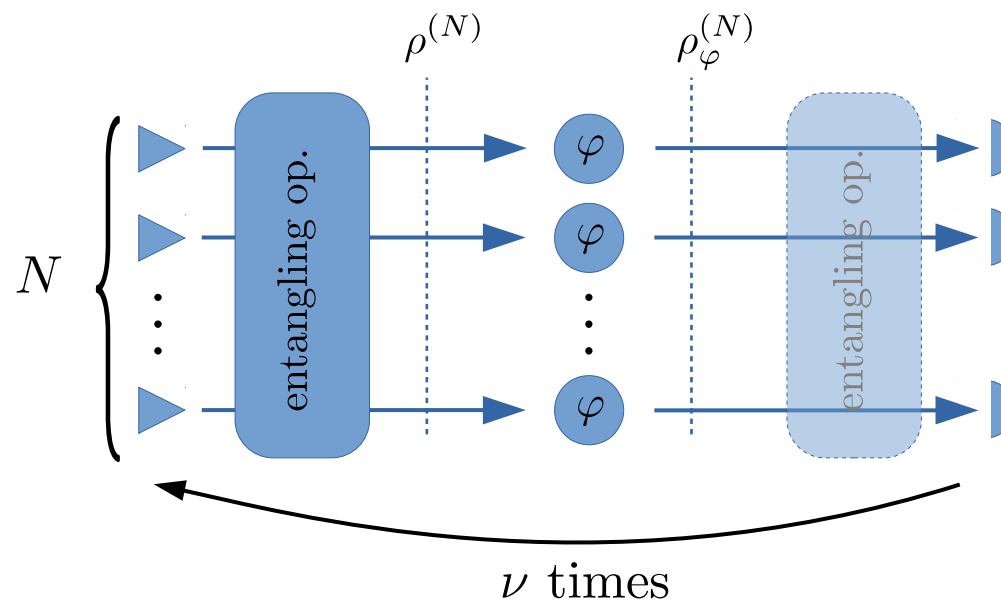
jonatanbohrbrask.dk/scaling



For independent probe particles



For entangled probes



The Heisenberg Limit

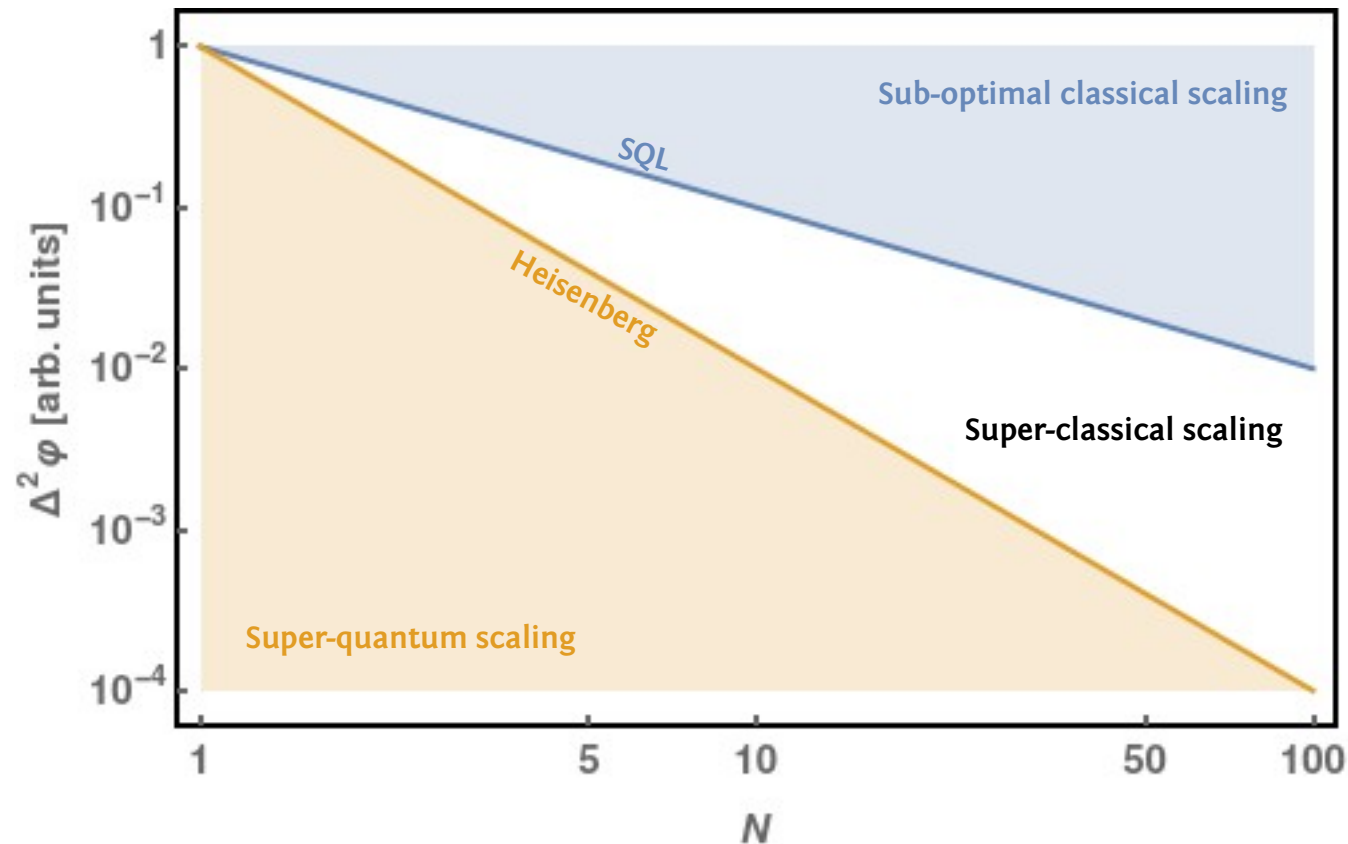
The best classical scaling

$$\Delta^2 \varphi \propto \frac{1}{\nu N}$$

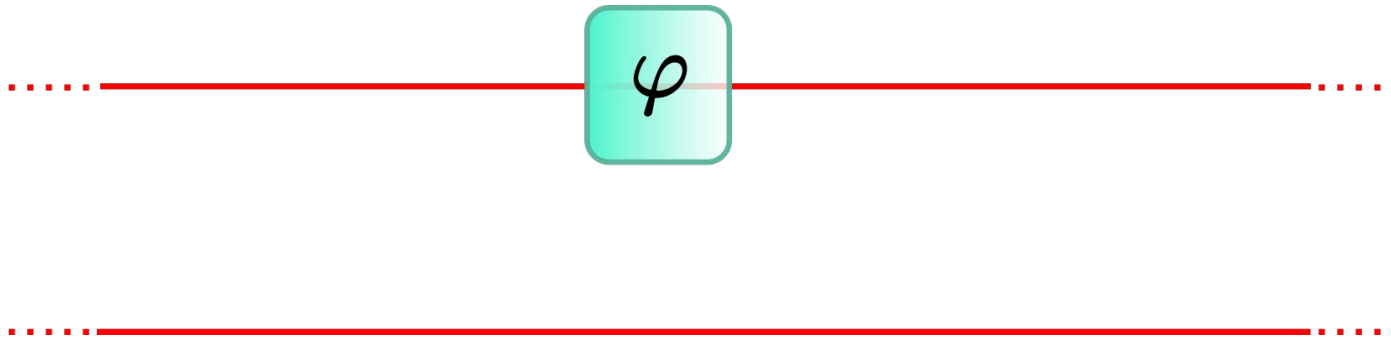
The best quantum scaling

$$\Delta^2 \varphi \propto \frac{1}{\nu N^2}$$

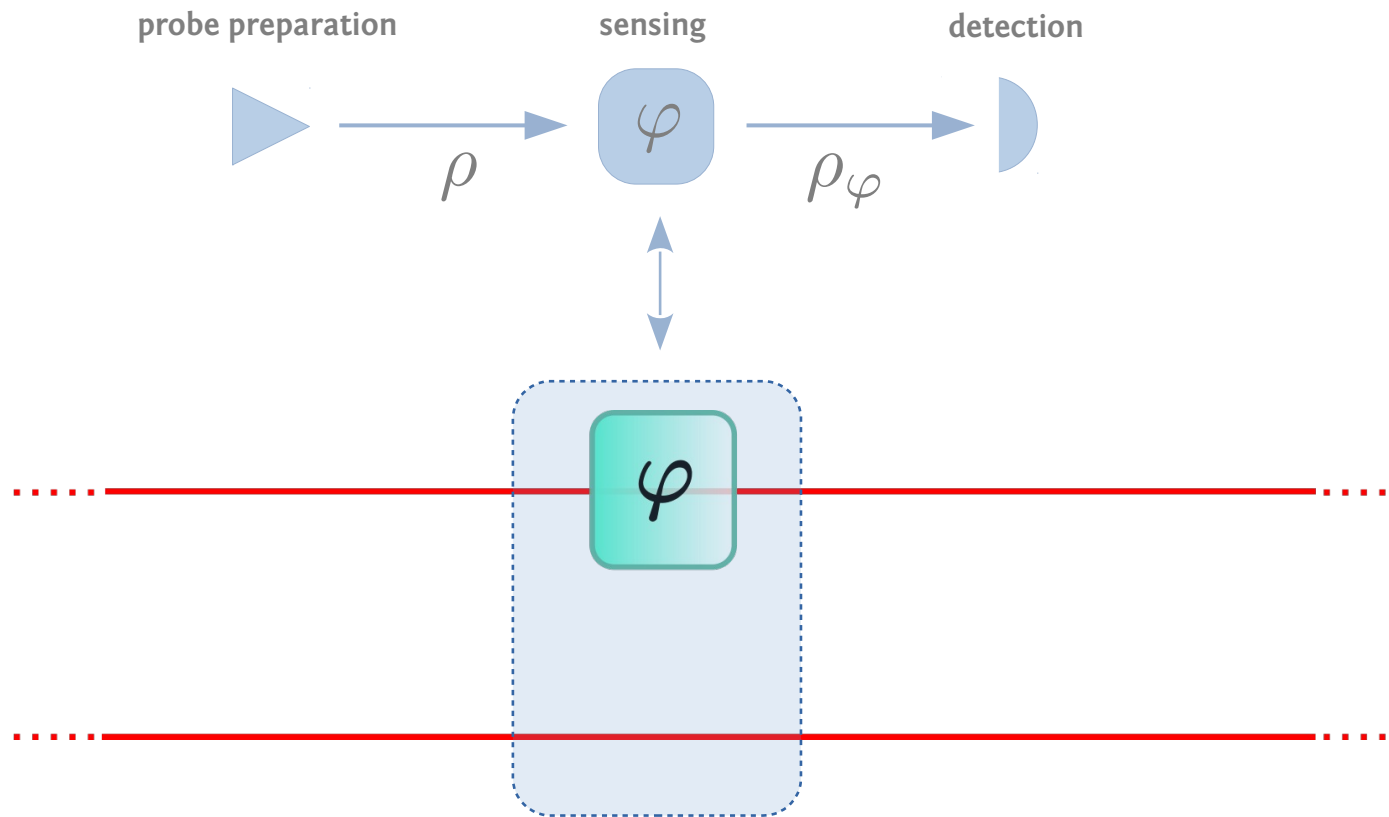
On a log-log scale...



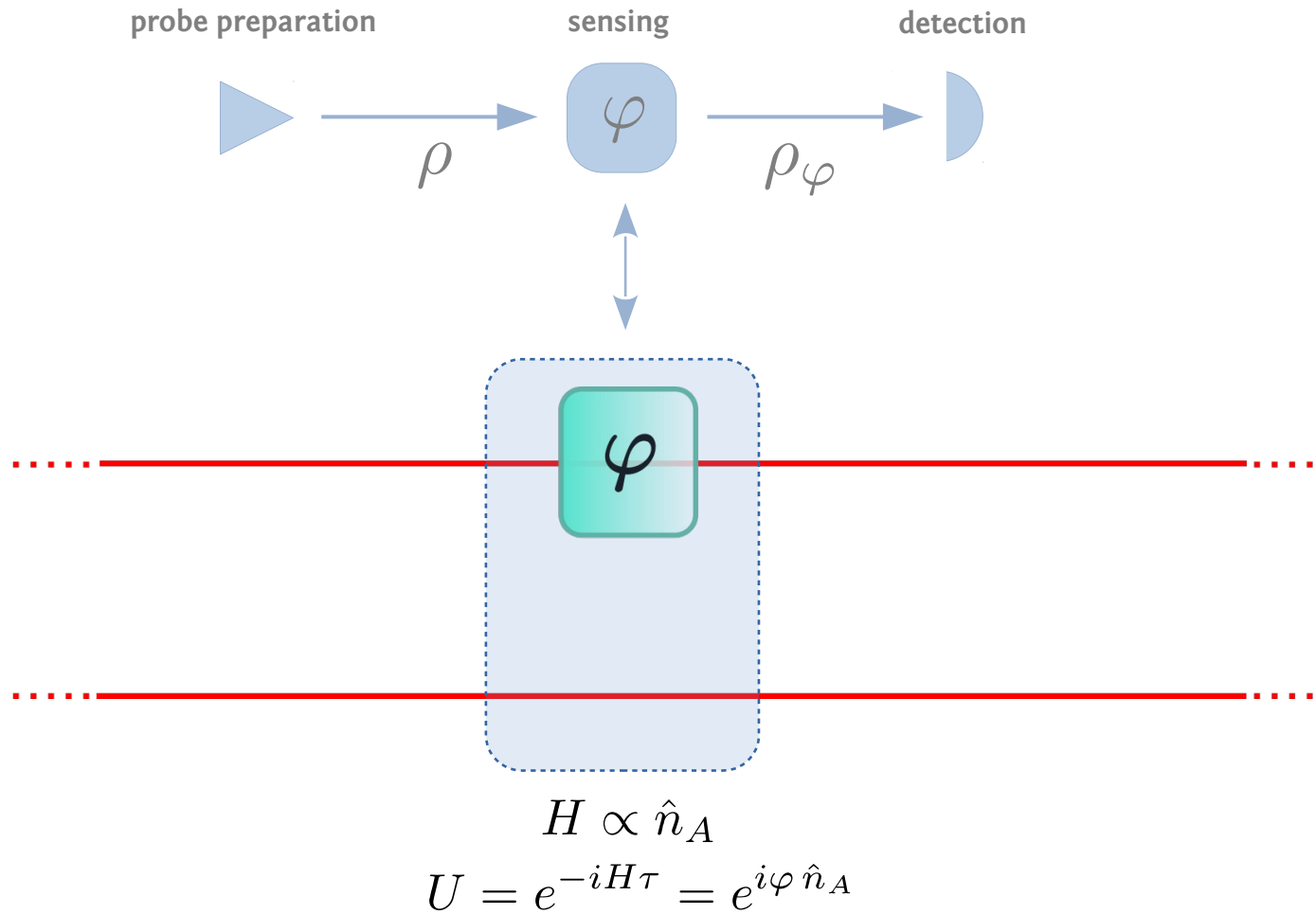
Achieving the Heisenberg limit – optical phase estimation



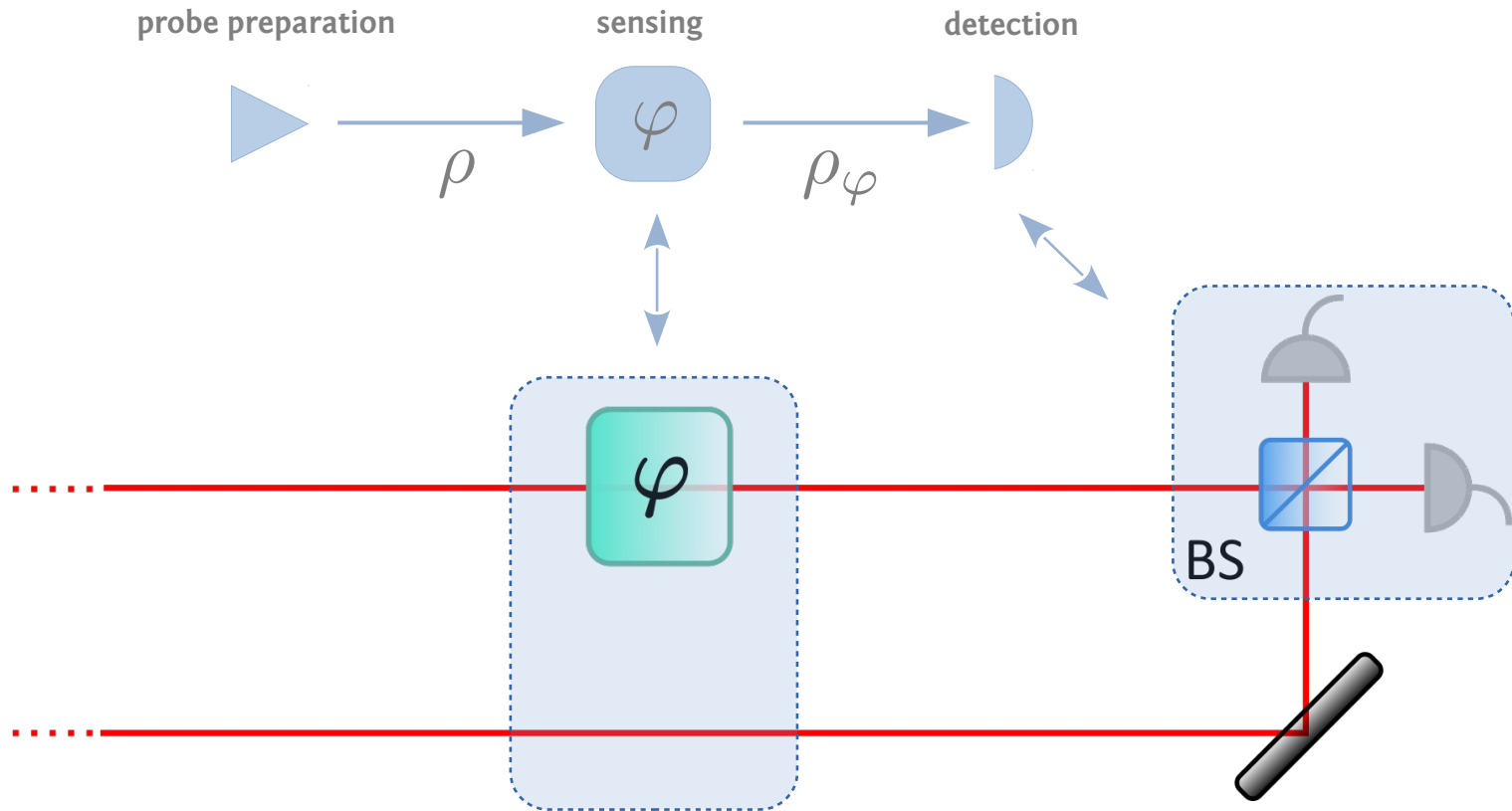
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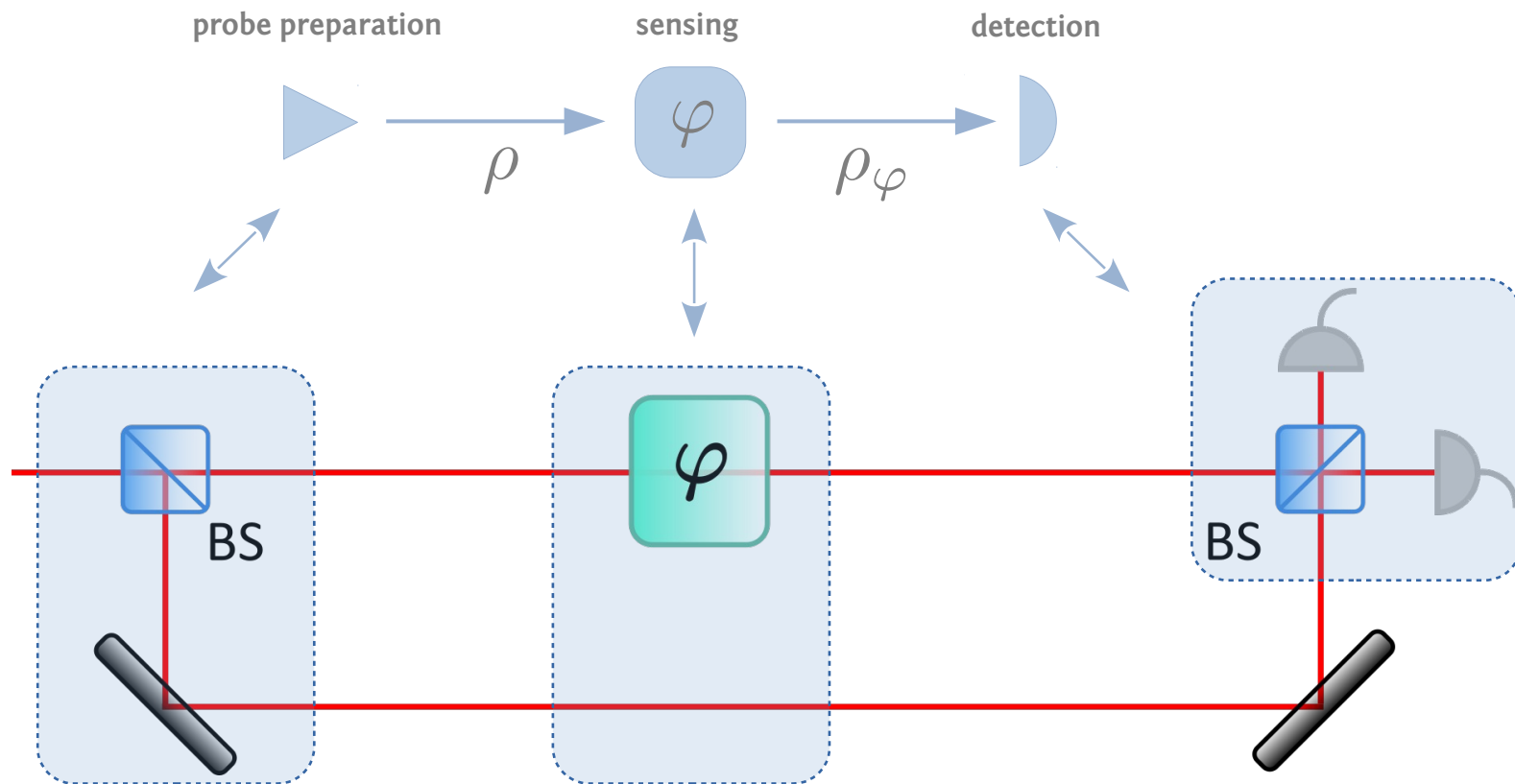


Achieving the Heisenberg limit – optical phase estimation



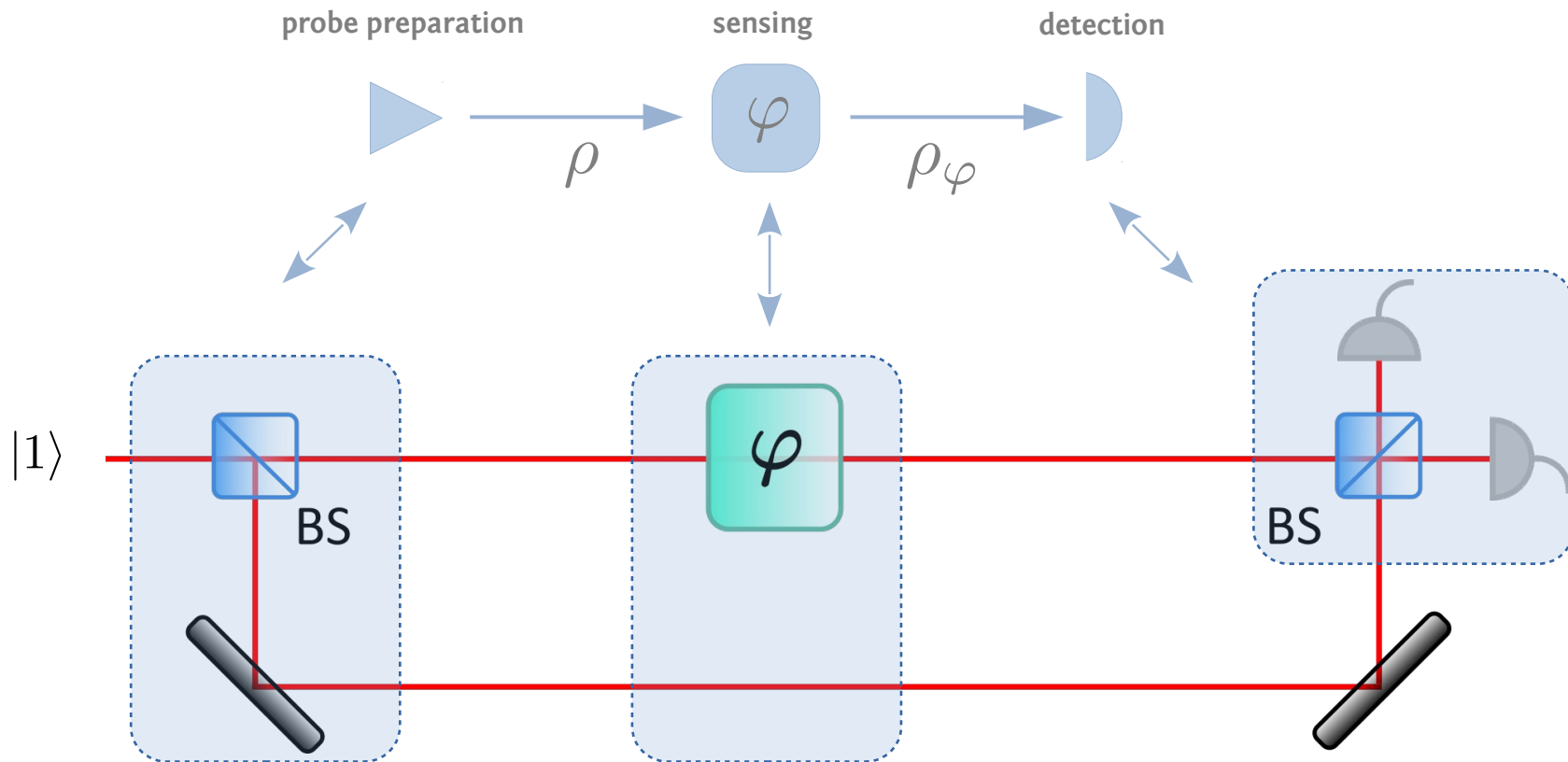
$$H \propto \hat{n}_A$$
$$U = e^{-iH\tau} = e^{i\varphi \hat{n}_A}$$

Achieving the Heisenberg limit – optical phase estimation



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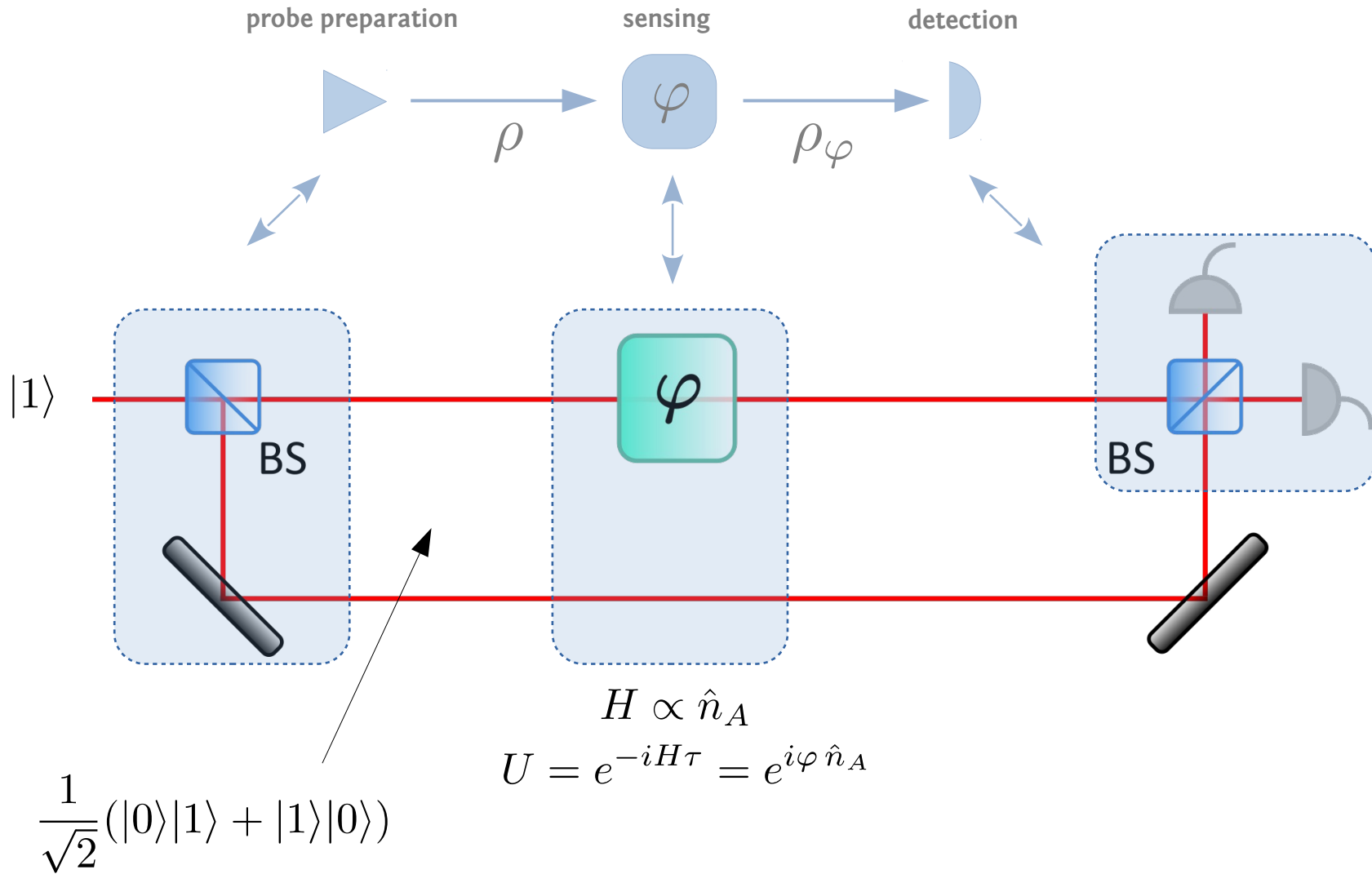
Achieving the Heisenberg limit – optical phase estimation



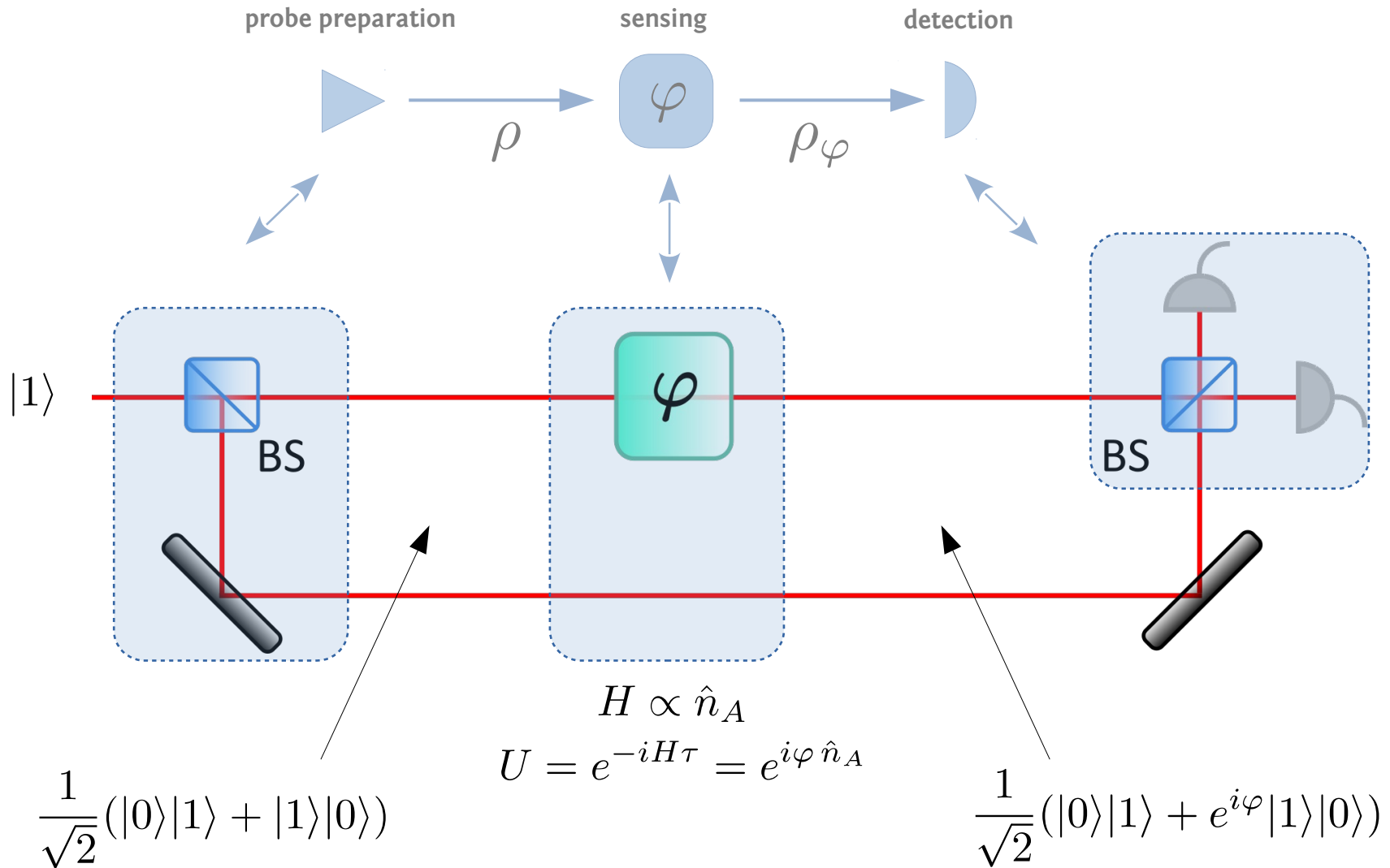
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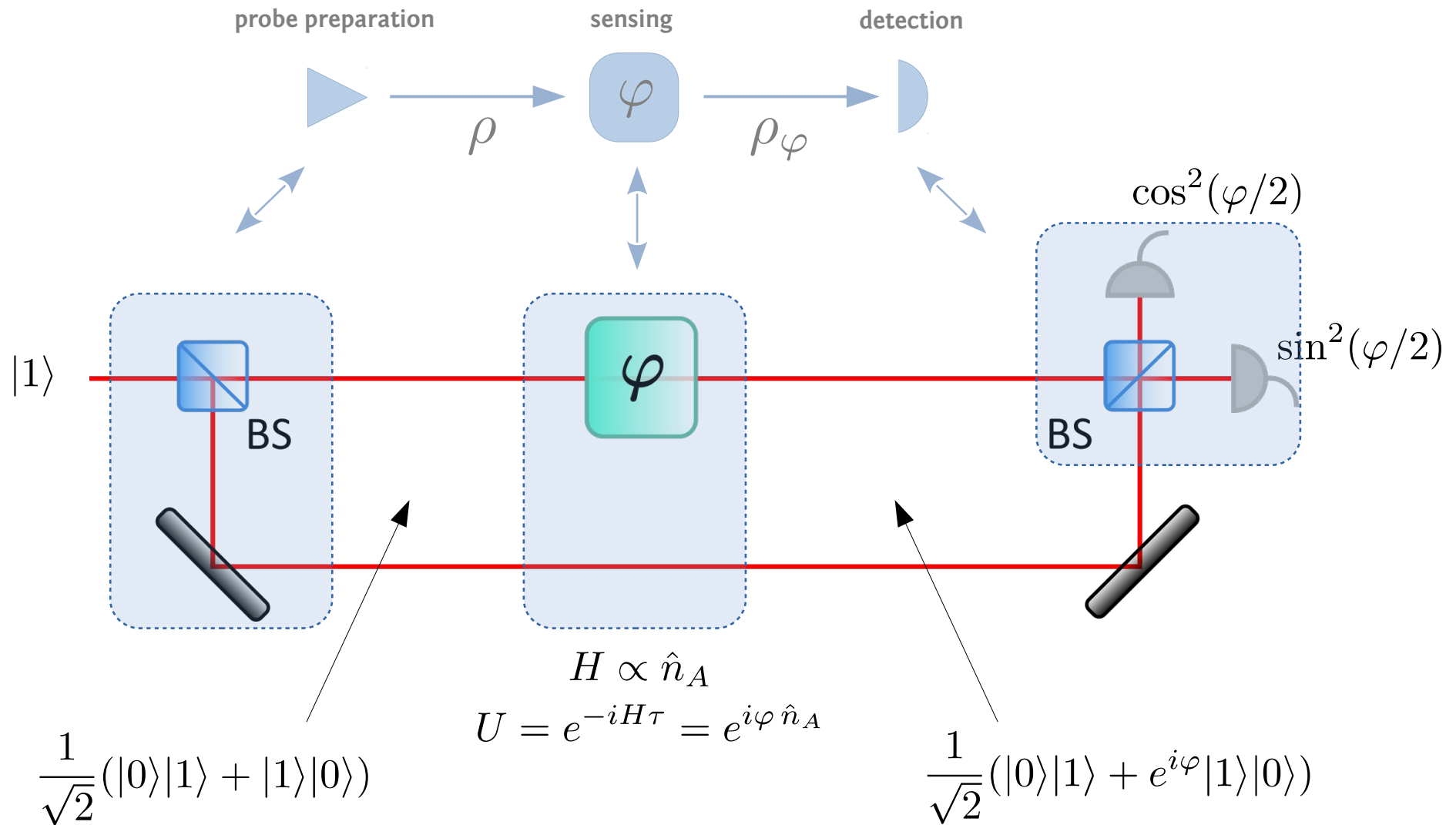
Achieving the Heisenberg limit – optical phase estimation



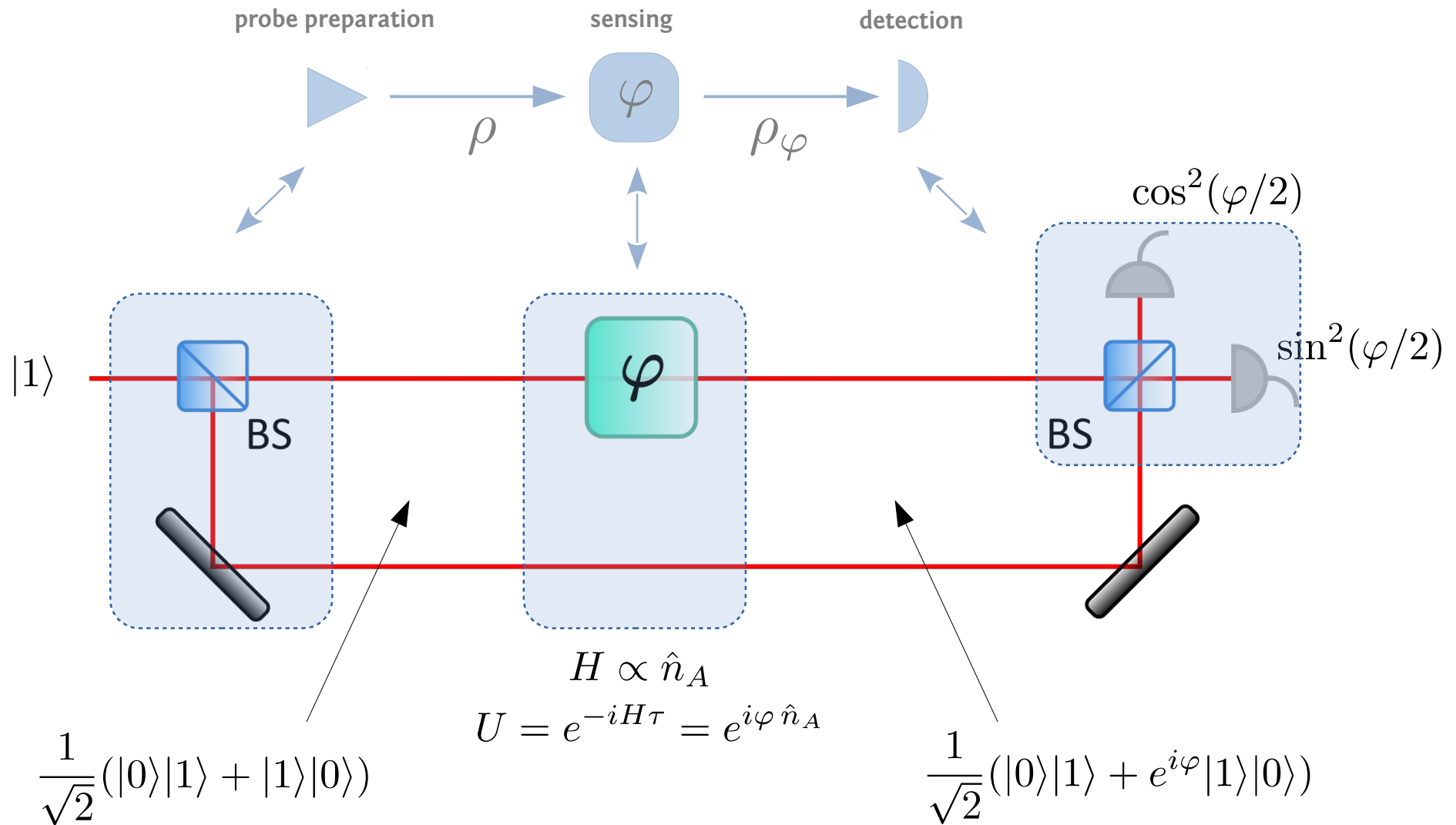
Achieving the Heisenberg limit – optical phase estimation



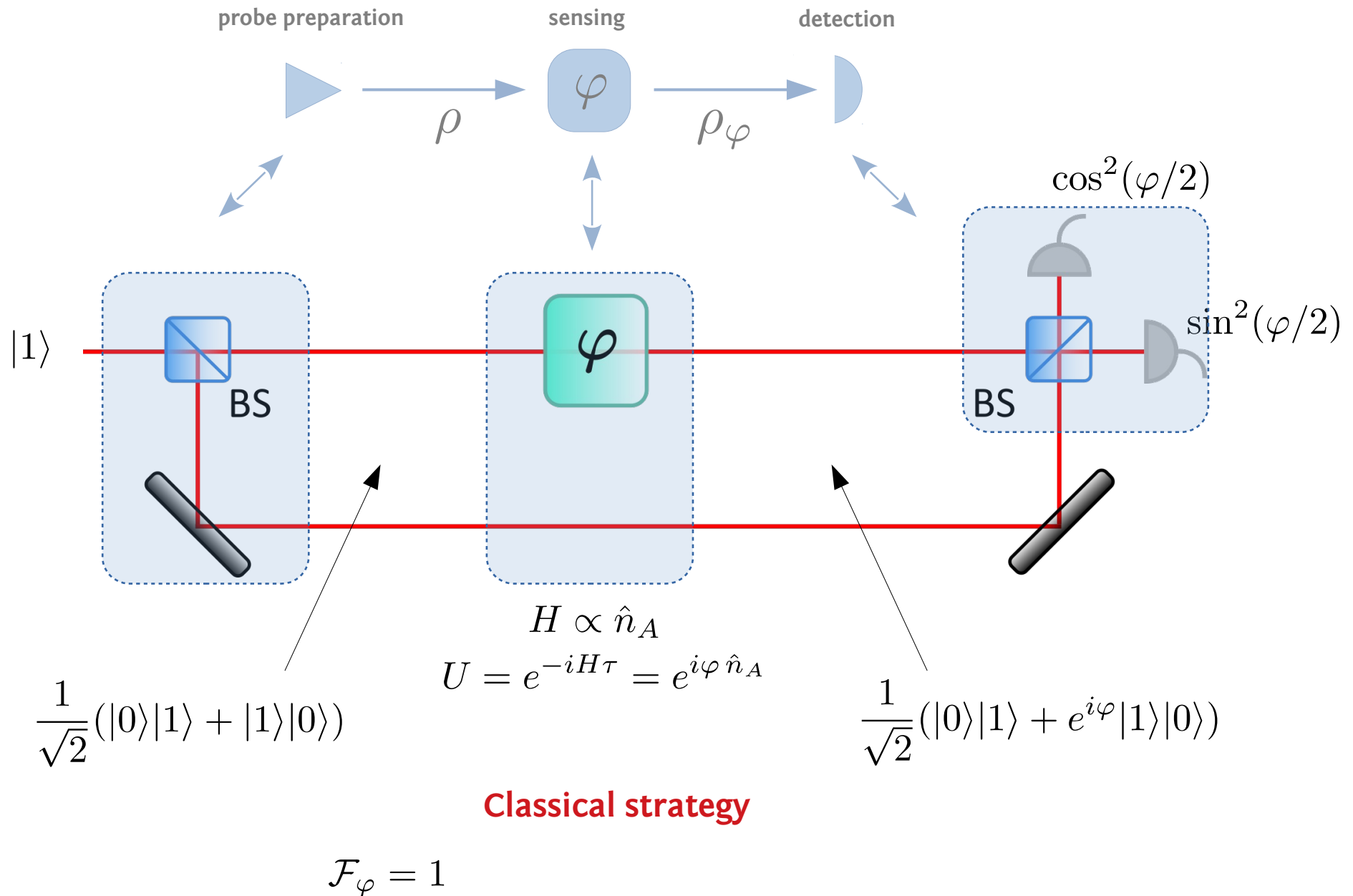
Achieving the Heisenberg limit – optical phase estimation



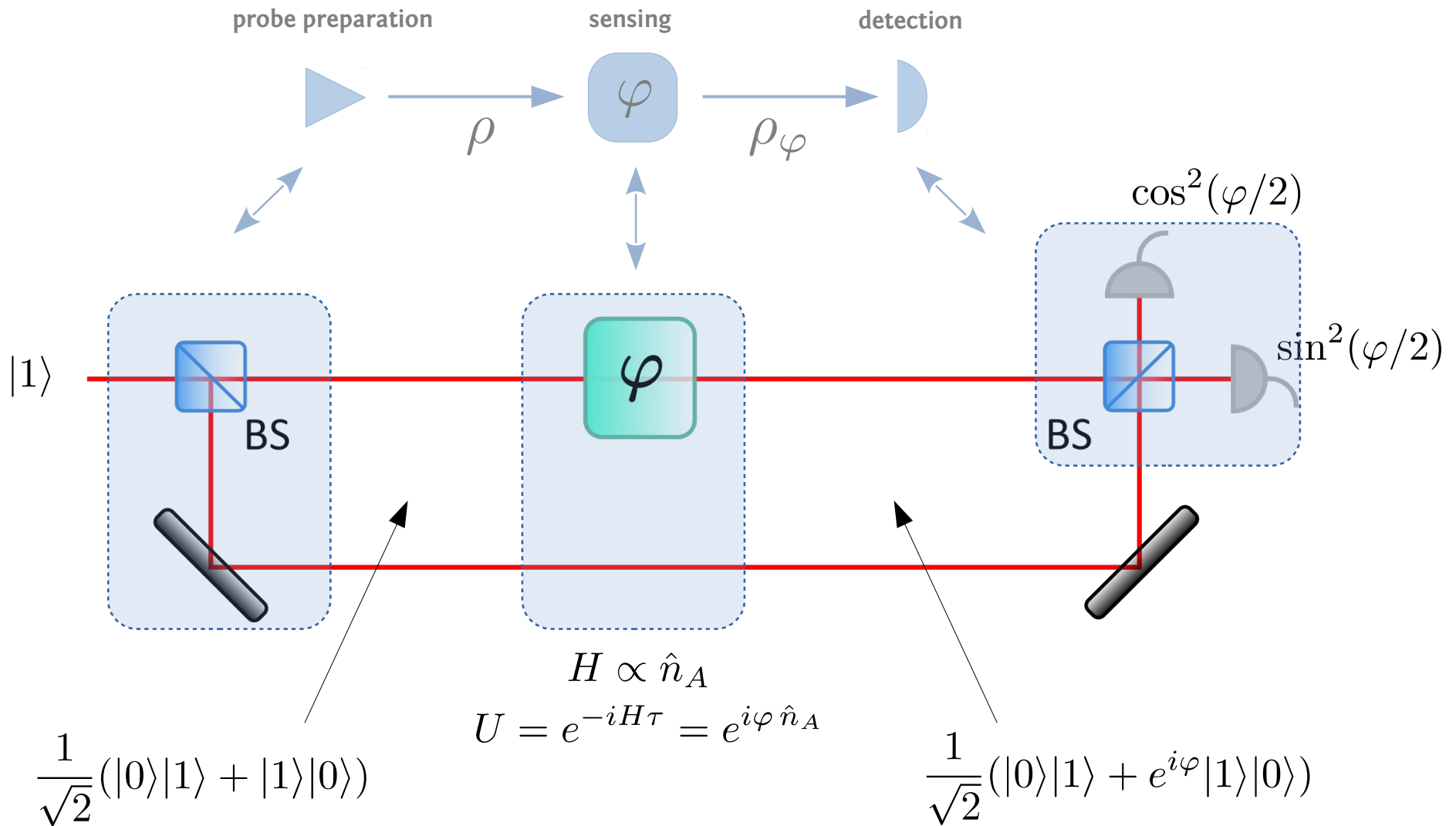
Achieving the Heisenberg limit – optical phase estimation



Achieving the Heisenberg limit – optical phase estimation



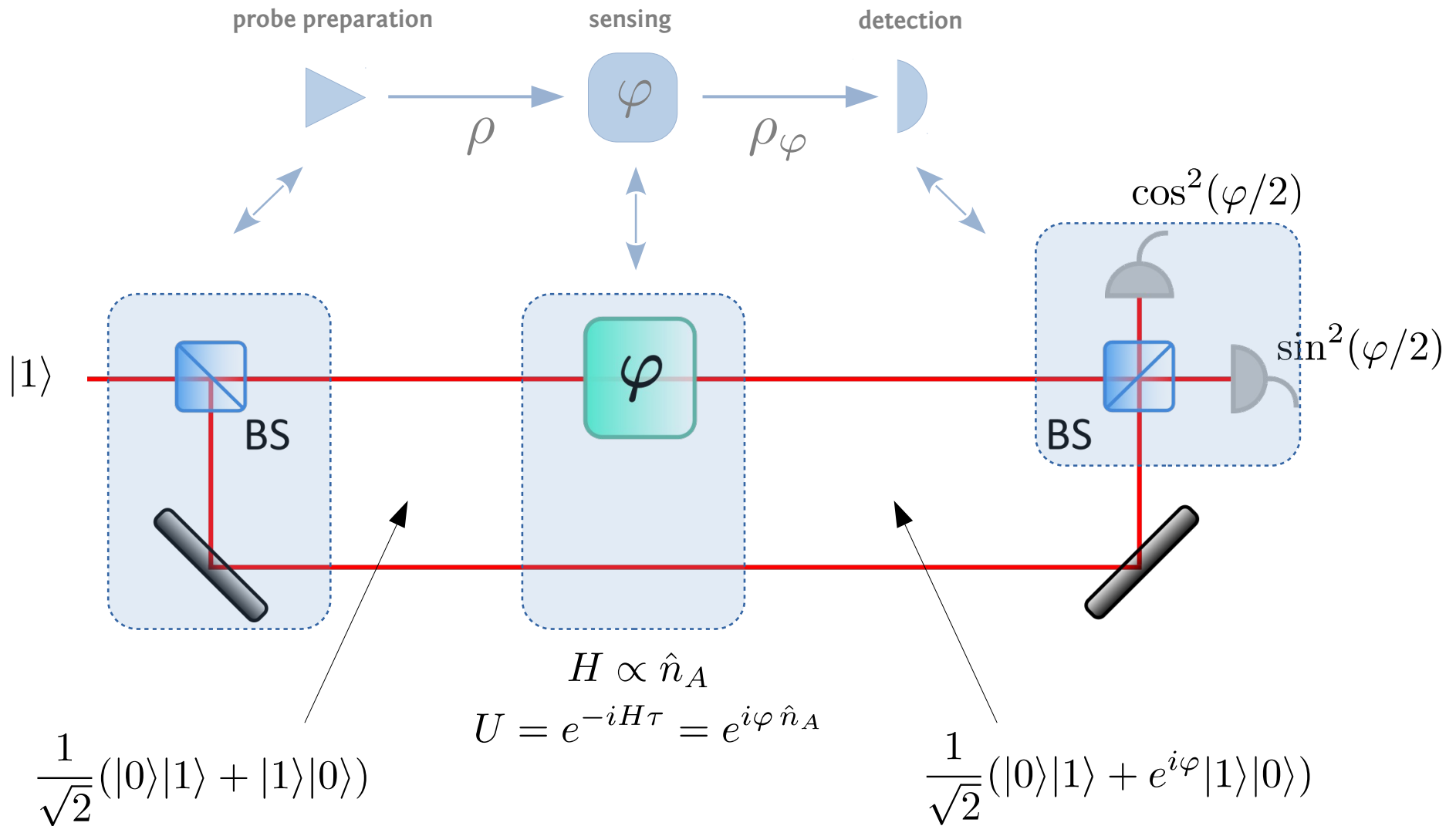
Achieving the Heisenberg limit – optical phase estimation



Classical strategy

$\mathcal{F}_\varphi = 1$
repeat νN times

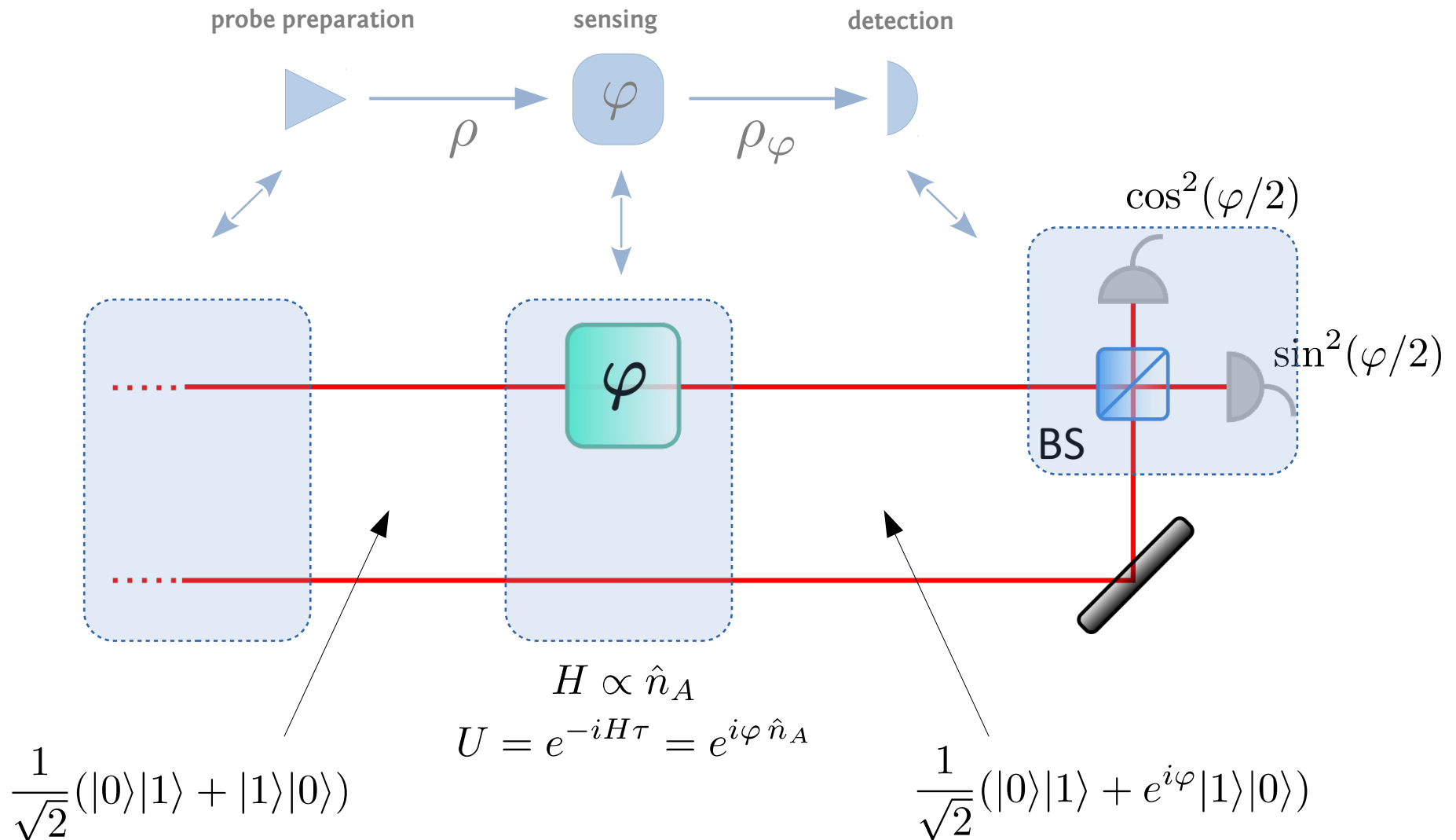
Achieving the Heisenberg limit – optical phase estimation



Classical strategy

$$\mathcal{F}_\phi = 1 \quad \text{repeat } \nu N \text{ times} \quad \longrightarrow \quad \Delta^2 \phi = \frac{1}{\nu N}$$

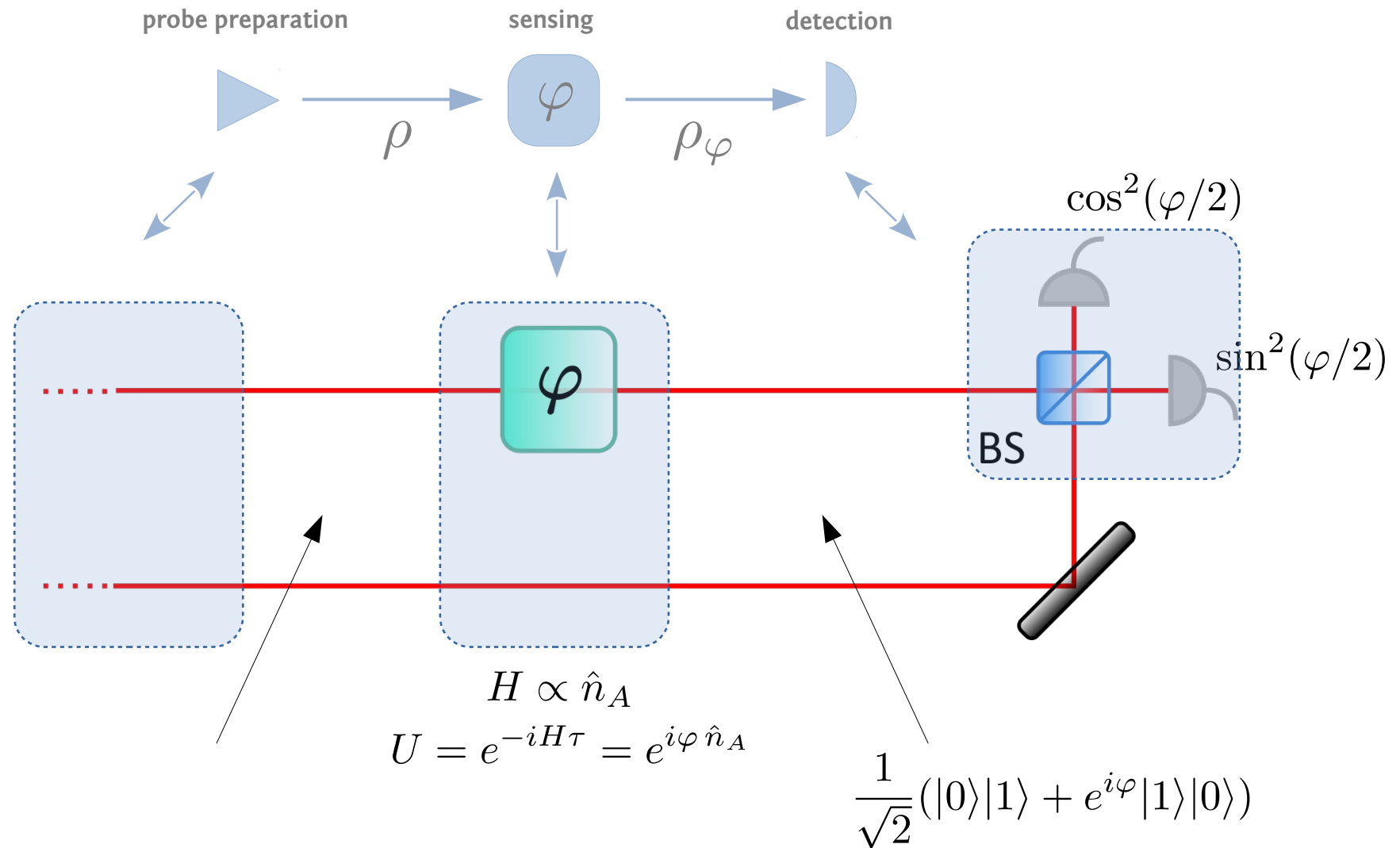
Achieving the Heisenberg limit – optical phase estimation



Classical strategy

$$\mathcal{F}_\varphi = 1 \quad \text{repeat } \nu N \text{ times} \quad \longrightarrow \quad \Delta^2 \varphi = \frac{1}{\nu N}$$

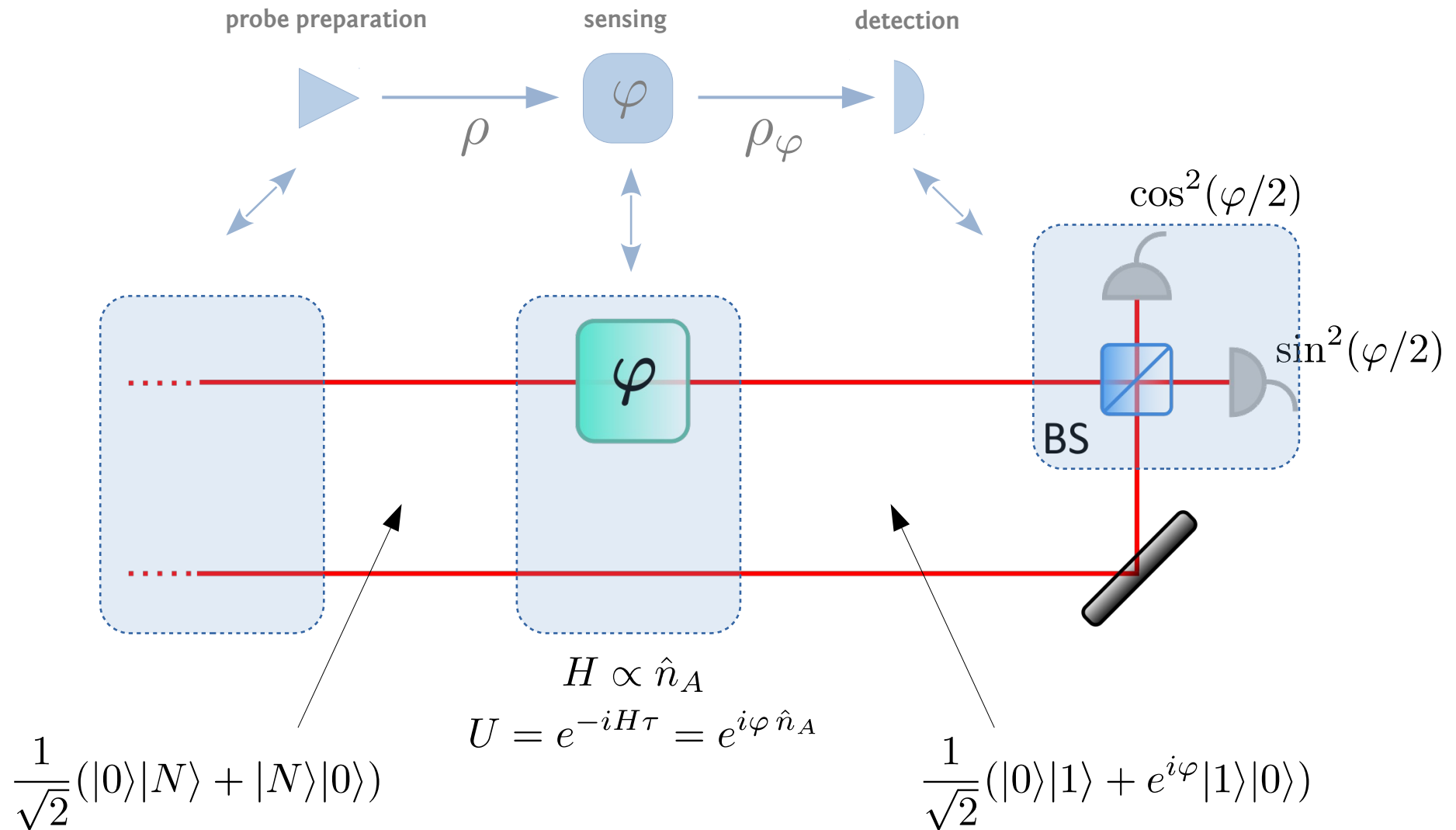
Achieving the Heisenberg limit – optical phase estimation



Classical strategy

$$\begin{array}{ccc} \mathcal{F}_\varphi = 1 & \xrightarrow{\text{repeat } \nu N \text{ times}} & \Delta^2 \varphi = \frac{1}{\nu N} \end{array}$$

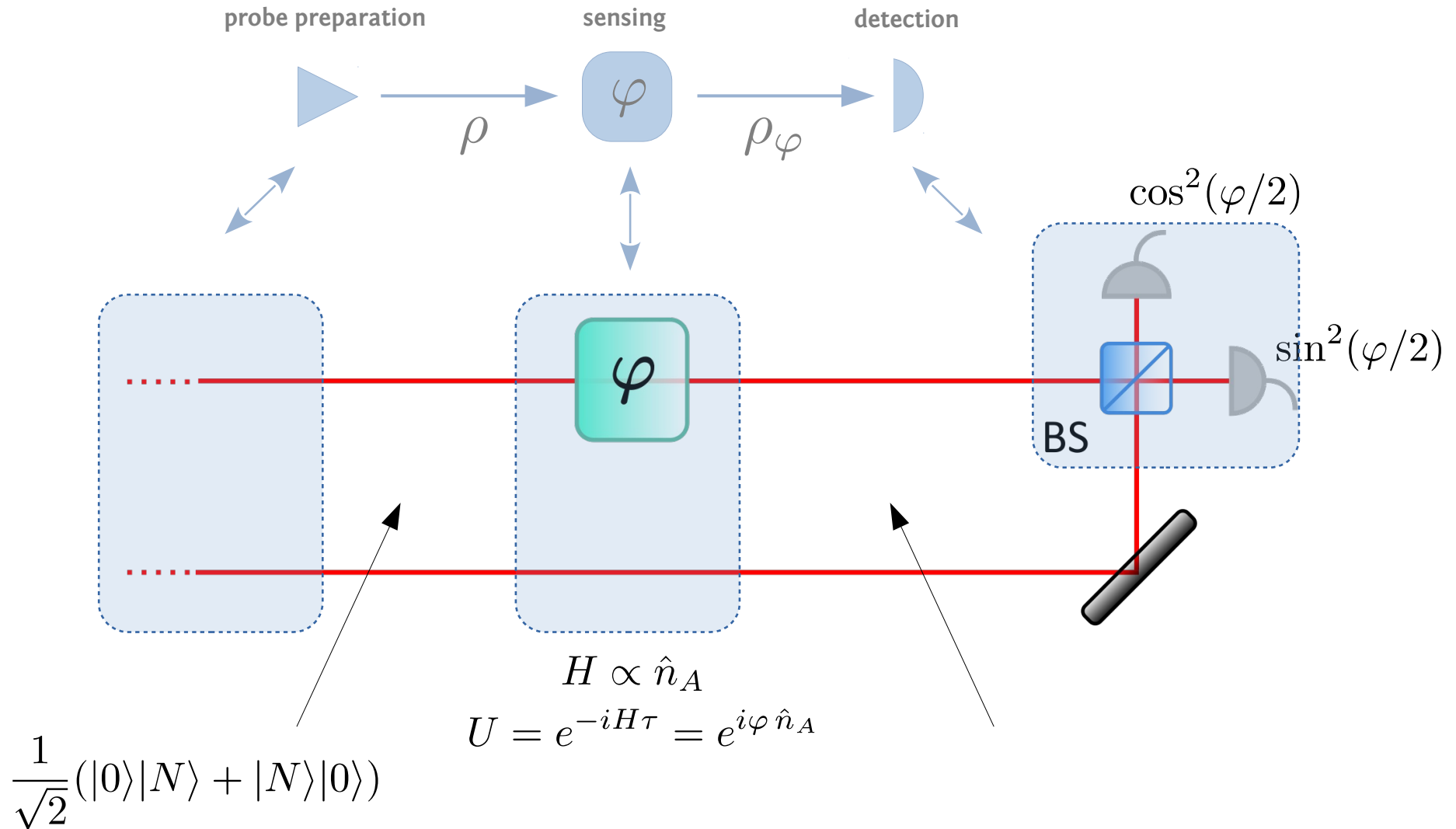
Achieving the Heisenberg limit – optical phase estimation



Classical strategy

$$\mathcal{F}_\varphi = 1 \quad \text{repeat } \nu N \text{ times} \quad \longrightarrow \quad \Delta^2 \varphi = \frac{1}{\nu N}$$

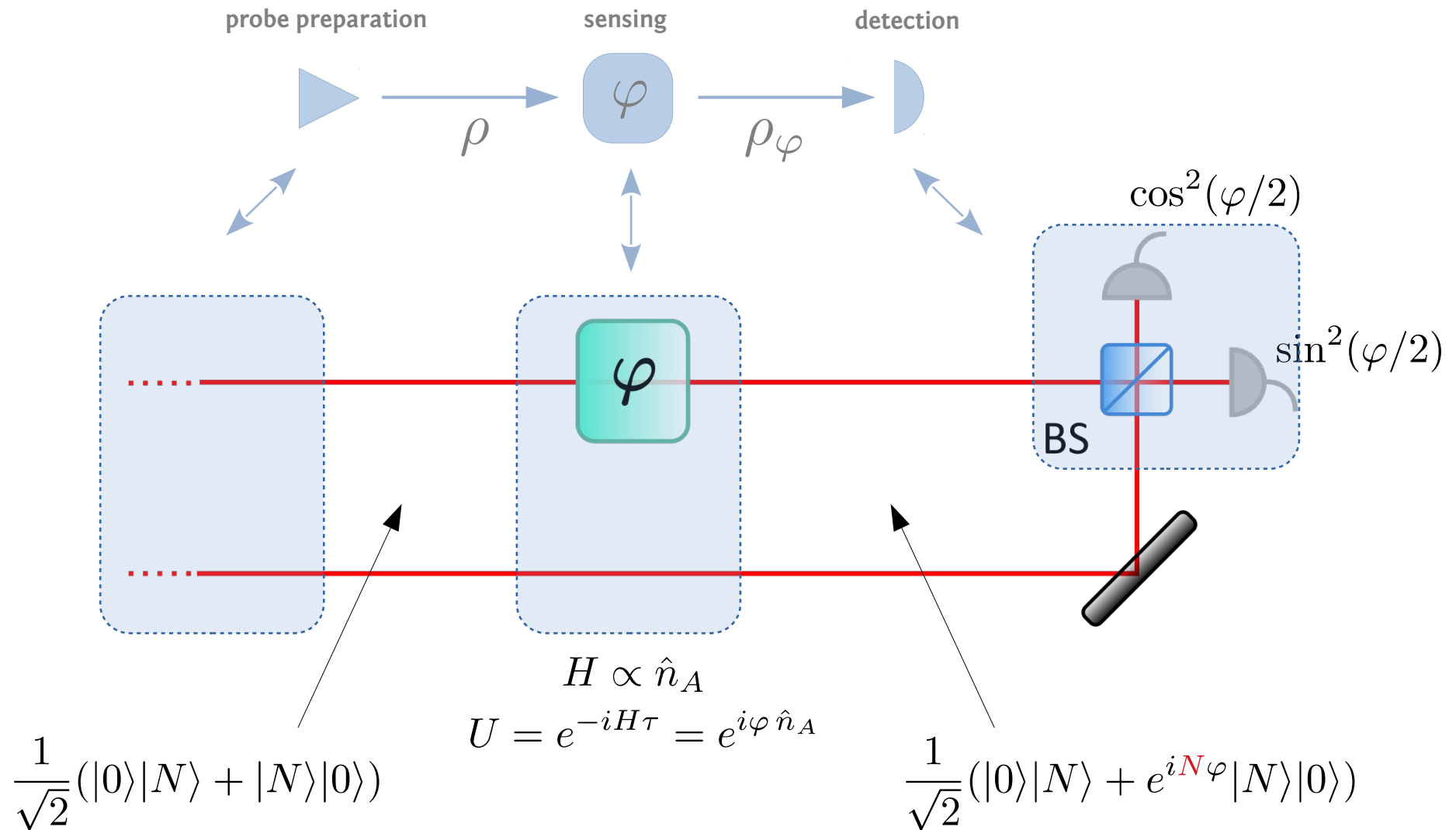
Achieving the Heisenberg limit – optical phase estimation



Classical strategy

$$\begin{array}{c} \mathcal{F}_\varphi = 1 \\ \text{repeat } \nu N \text{ times} \end{array} \quad \longrightarrow \quad \Delta^2 \varphi = \frac{1}{\nu N}$$

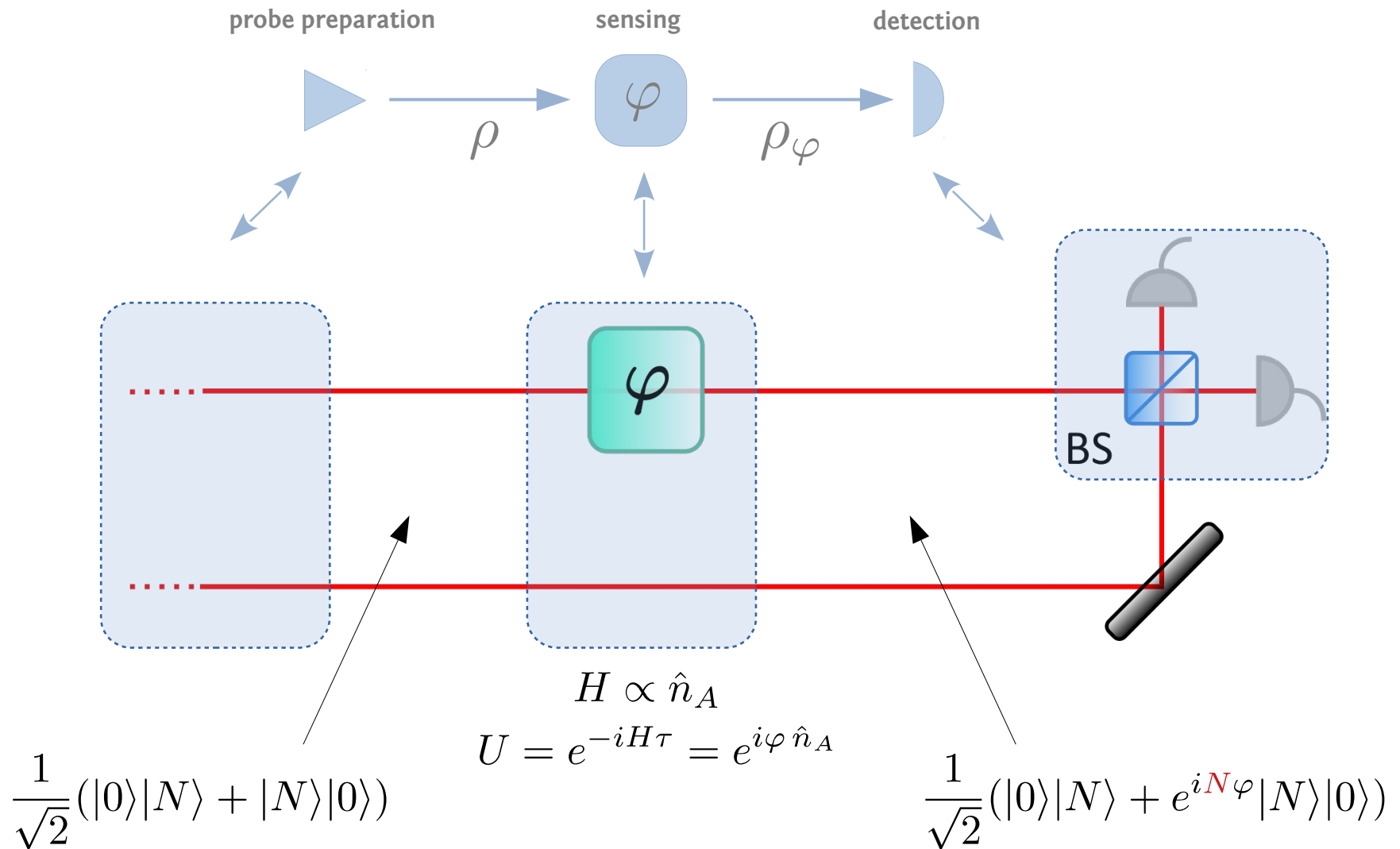
Achieving the Heisenberg limit – optical phase estimation



Classical strategy

$$\mathcal{F}_\varphi = 1 \quad \text{repeat } \nu N \text{ times} \quad \longrightarrow \quad \Delta^2 \varphi = \frac{1}{\nu N}$$

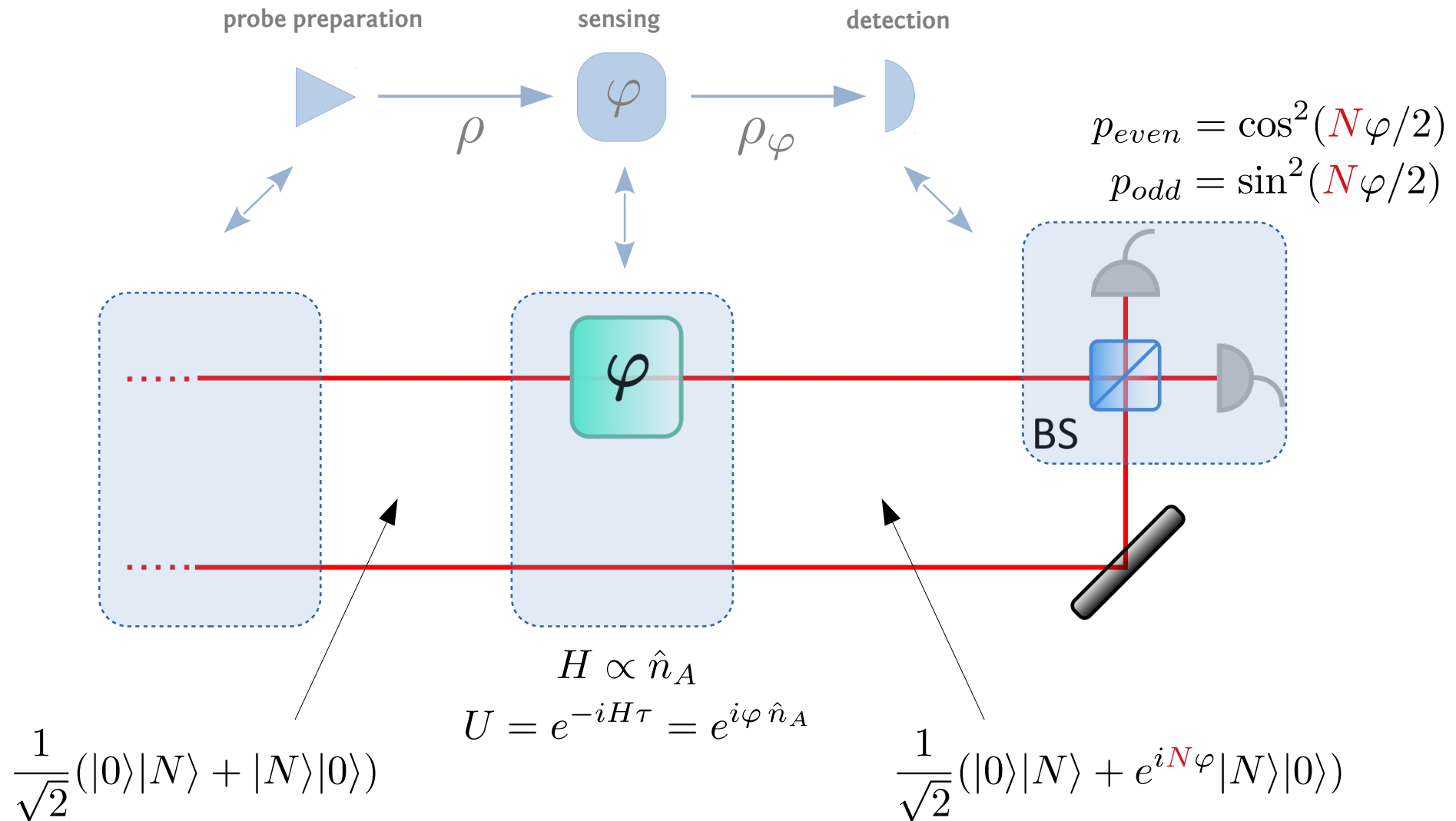
Achieving the Heisenberg limit – optical phase estimation



Classical strategy

$$\begin{array}{ccc} \mathcal{F}_\varphi = 1 & \longrightarrow & \Delta^2 \varphi = \frac{1}{\nu N} \\ \text{repeat } \nu N \text{ times} & & \end{array}$$

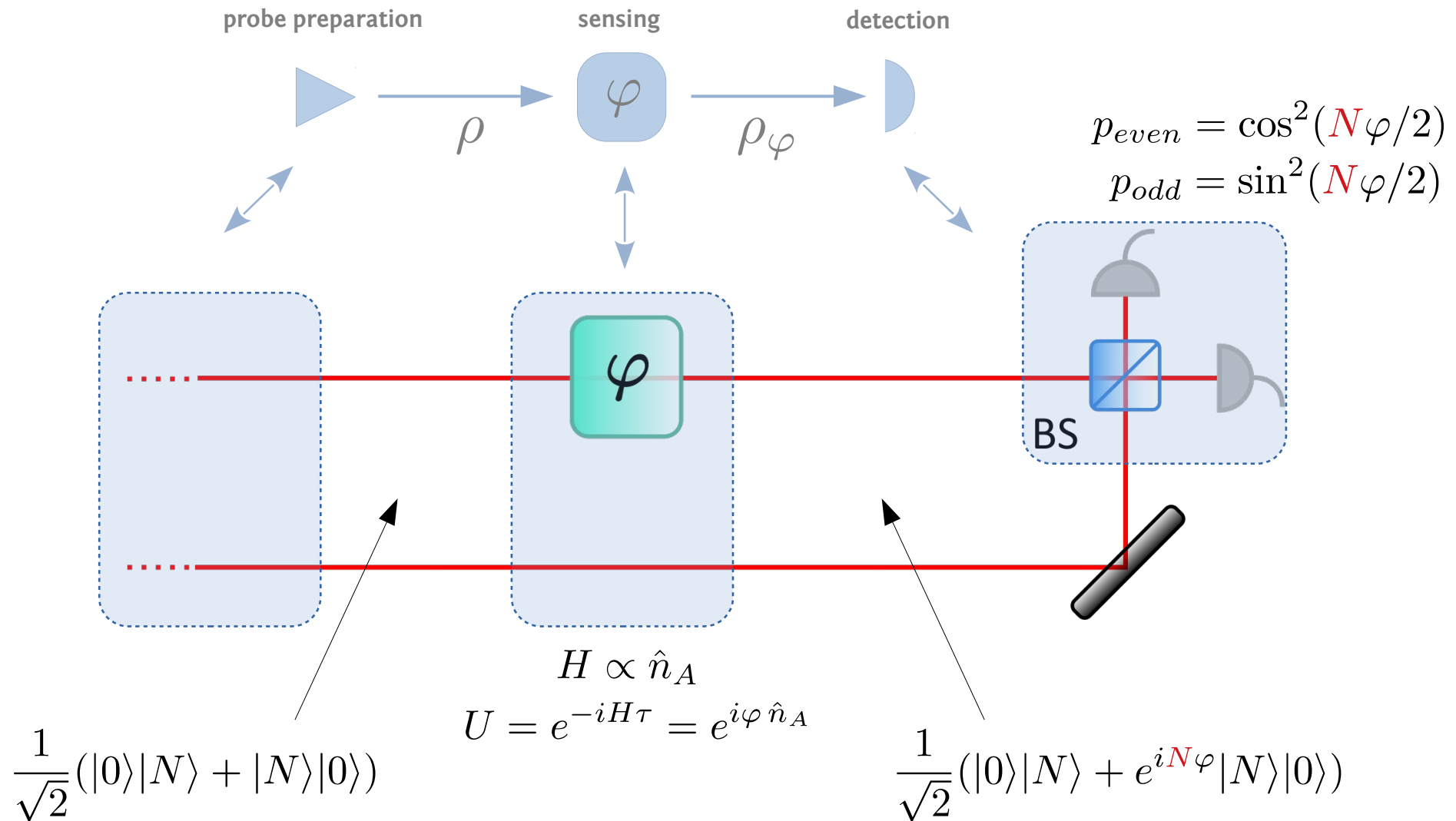
Achieving the Heisenberg limit – optical phase estimation



Classical strategy

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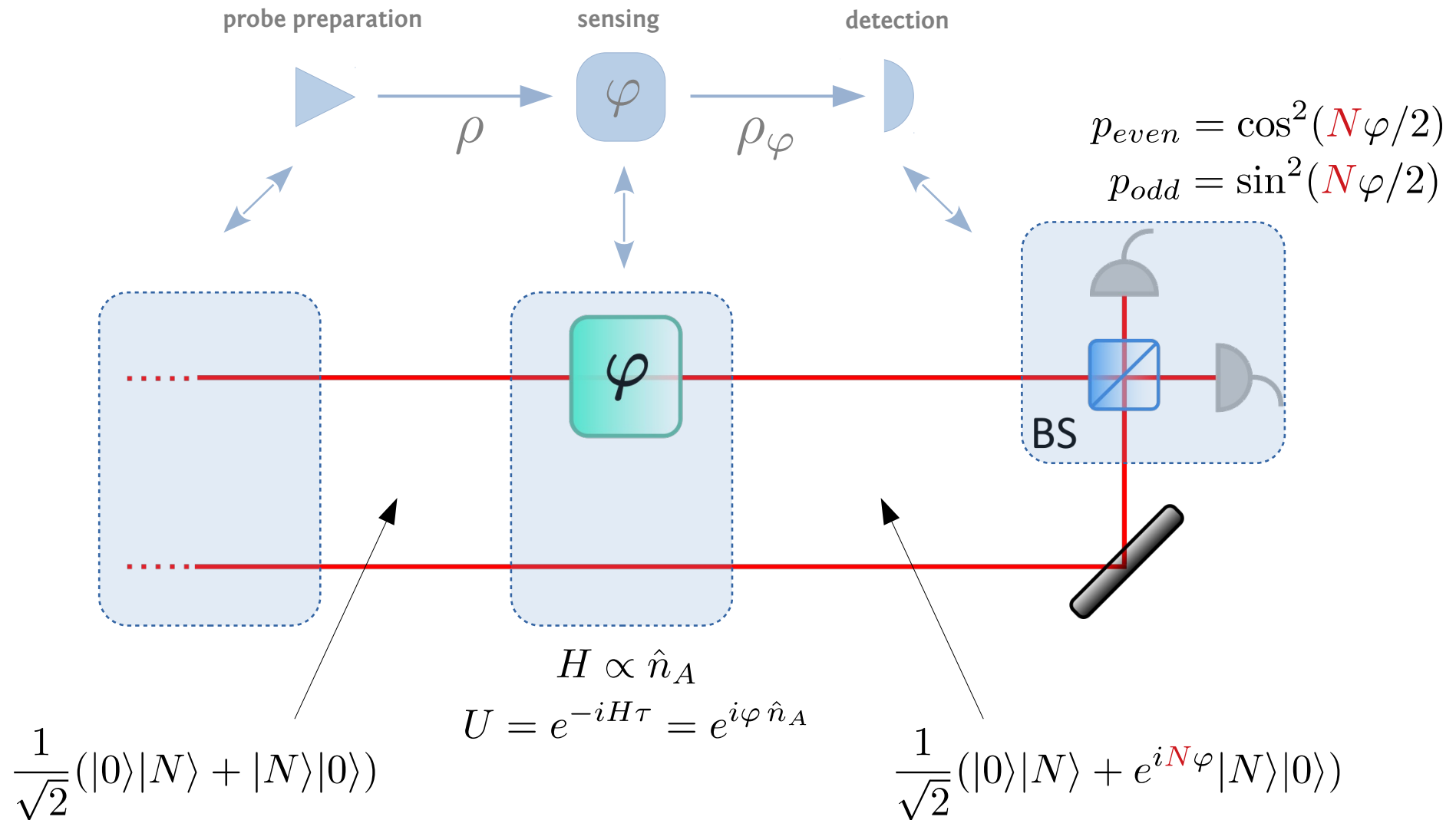
Achieving the Heisenberg limit – optical phase estimation



Classical strategy

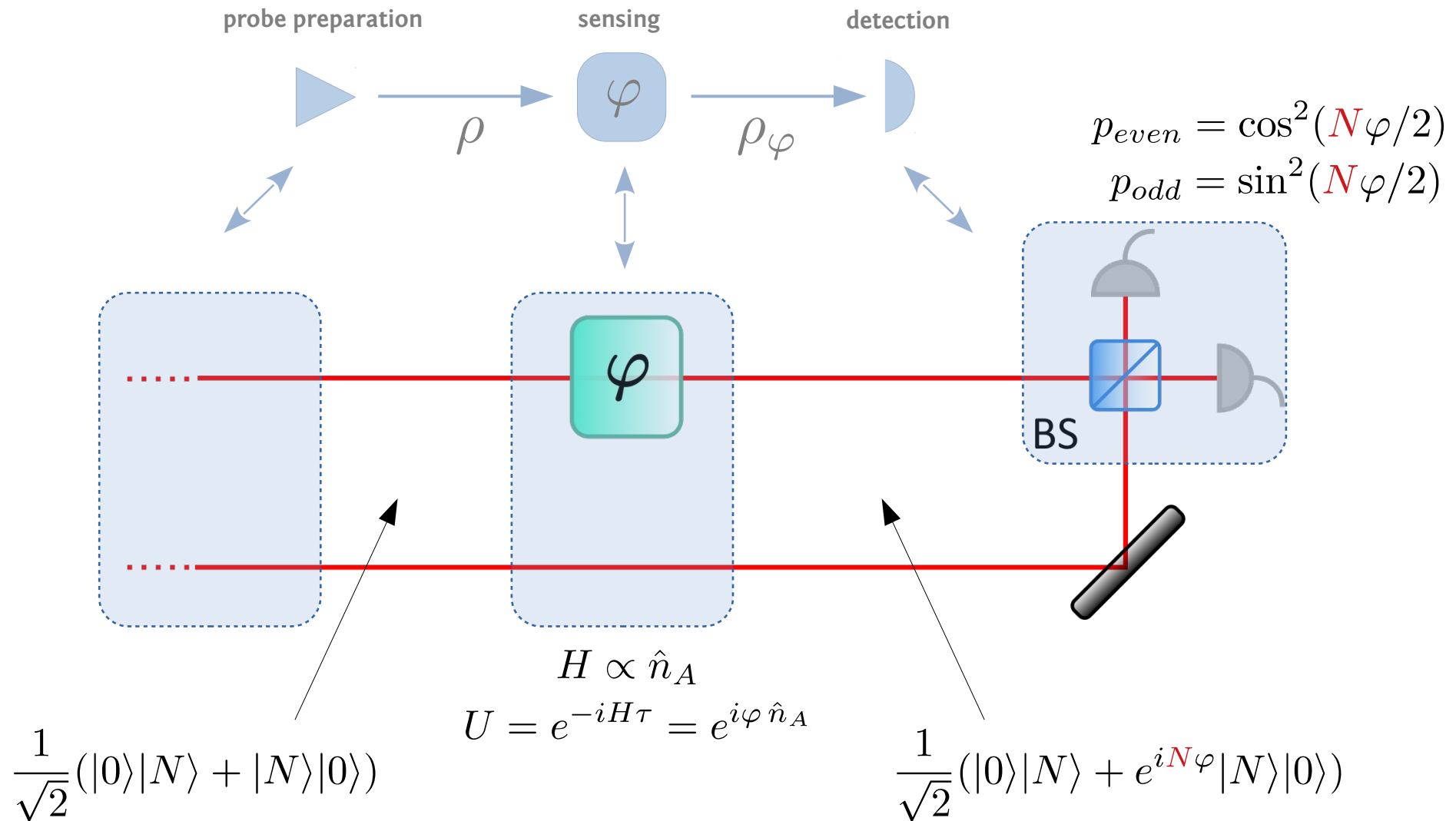
$$\mathcal{F}_\varphi = 1 \quad \text{repeat } \nu N \text{ times} \quad \longrightarrow \quad \Delta^2 \varphi = \frac{1}{\nu N}$$

Achieving the Heisenberg limit – optical phase estimation



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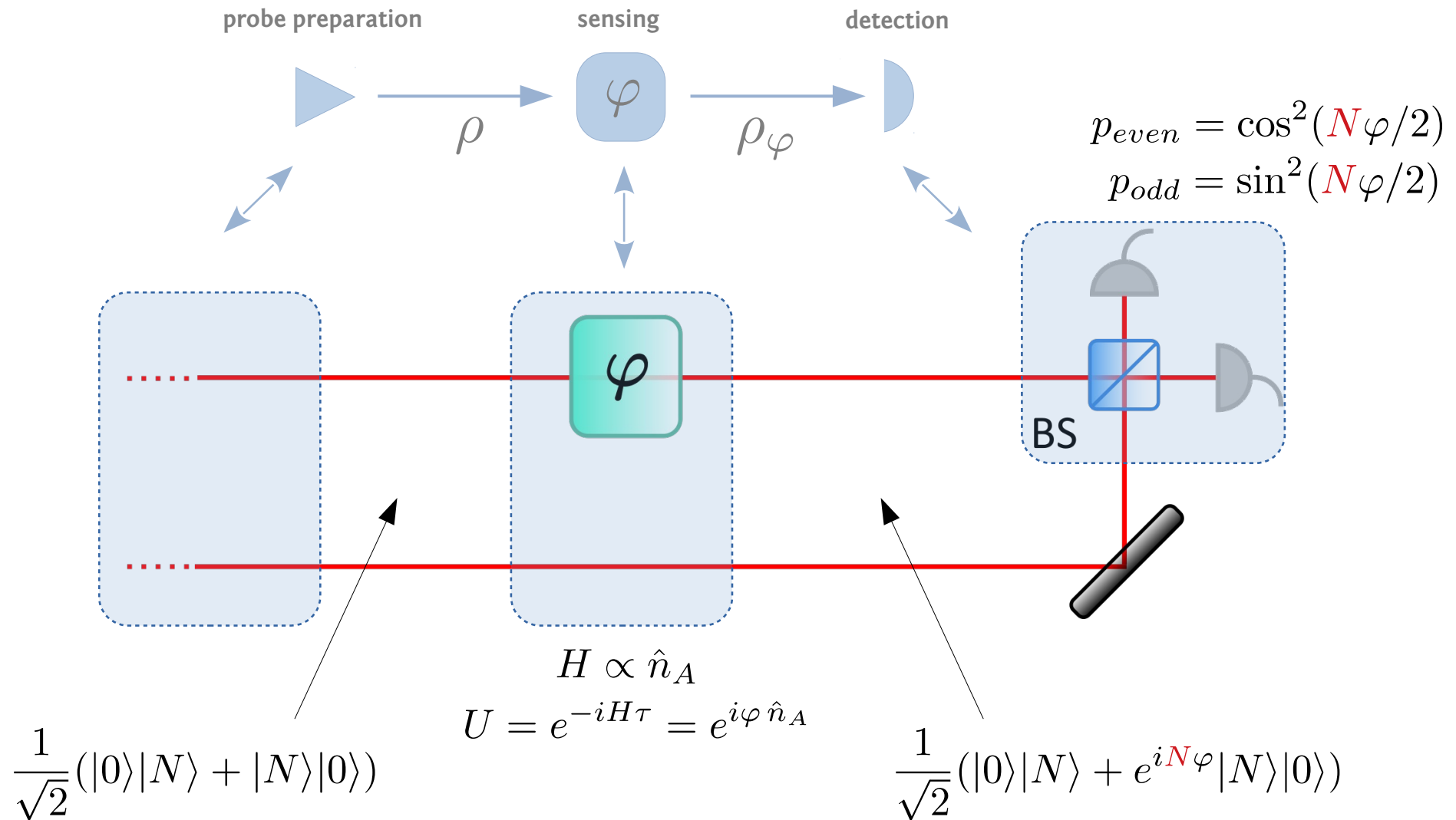
Achieving the Heisenberg limit – optical phase estimation



Quantum strategy

$$\mathcal{F}_\varphi = 1 \quad \text{repeat } \nu N \text{ times} \quad \longrightarrow \quad \Delta^2 \varphi = \frac{1}{\nu N}$$

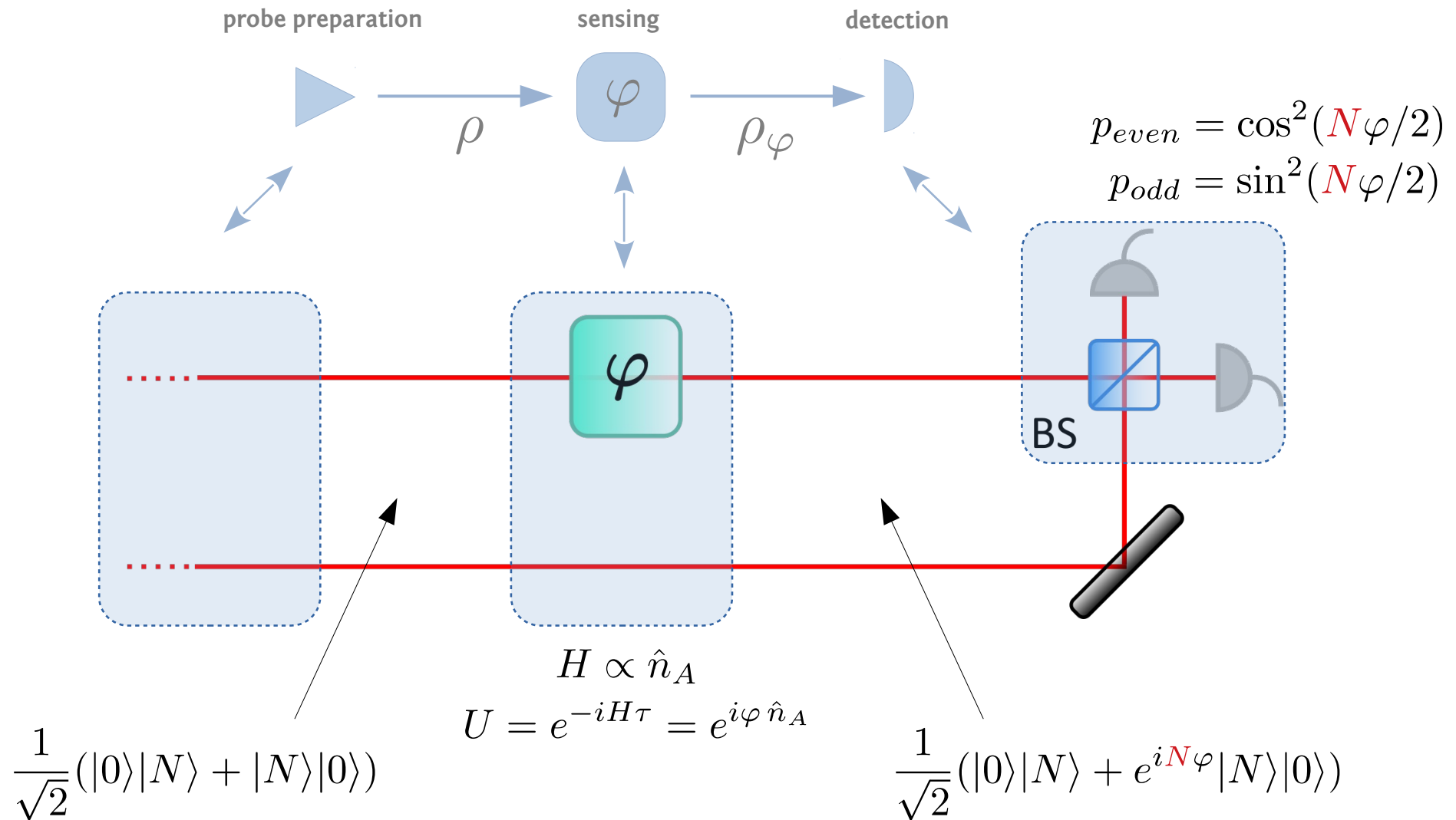
Achieving the Heisenberg limit – optical phase estimation



Quantum strategy

$\Delta^2\varphi = \frac{1}{\nu N}$

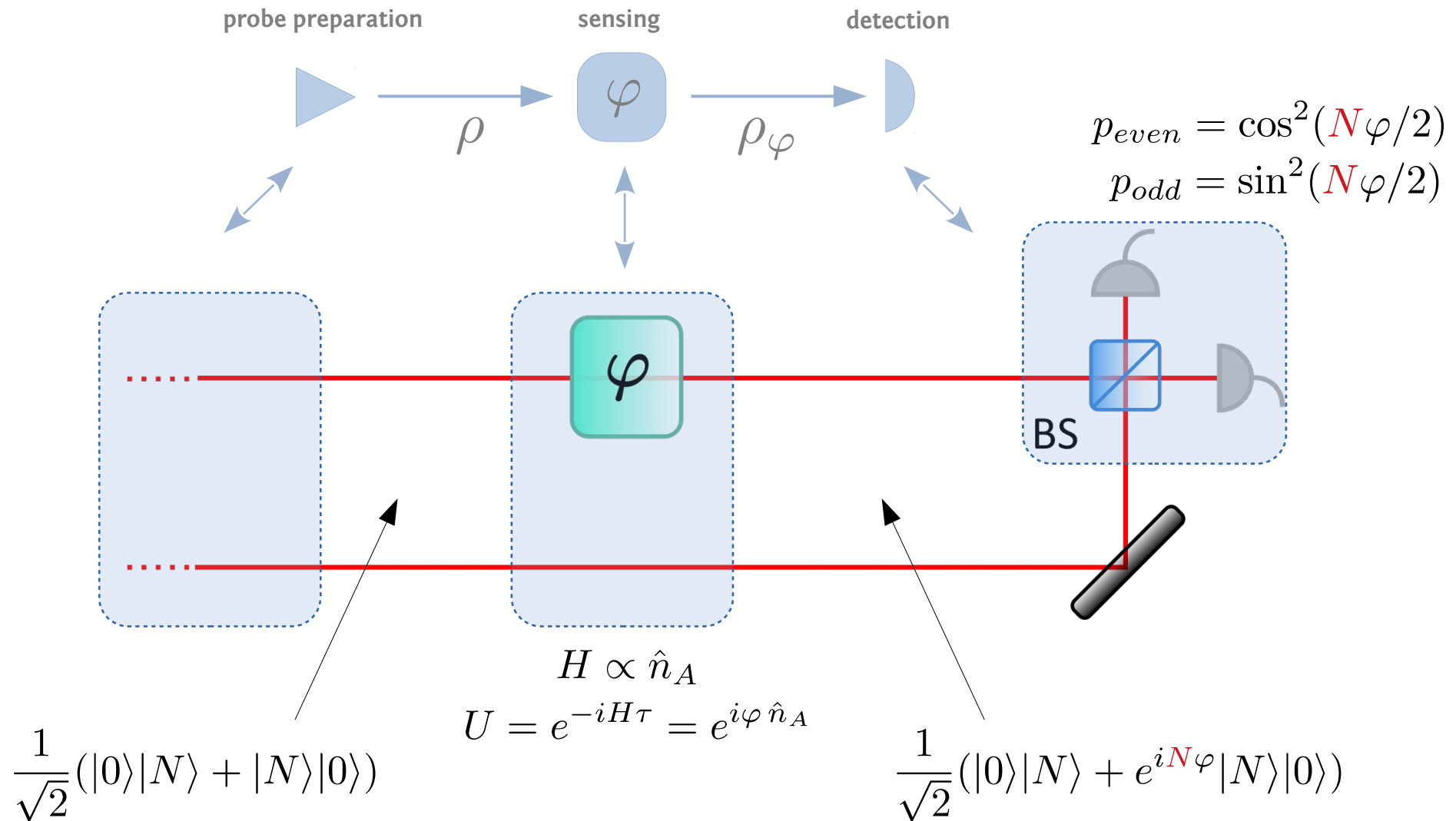
Achieving the Heisenberg limit – optical phase estimation



Quantum strategy

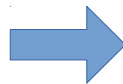
$$\mathcal{F}_\varphi = N^2 \quad \text{repeat } \nu \text{ times} \quad \longrightarrow \quad \Delta^2 \varphi = \frac{1}{\nu N}$$

Achieving the Heisenberg limit – optical phase estimation

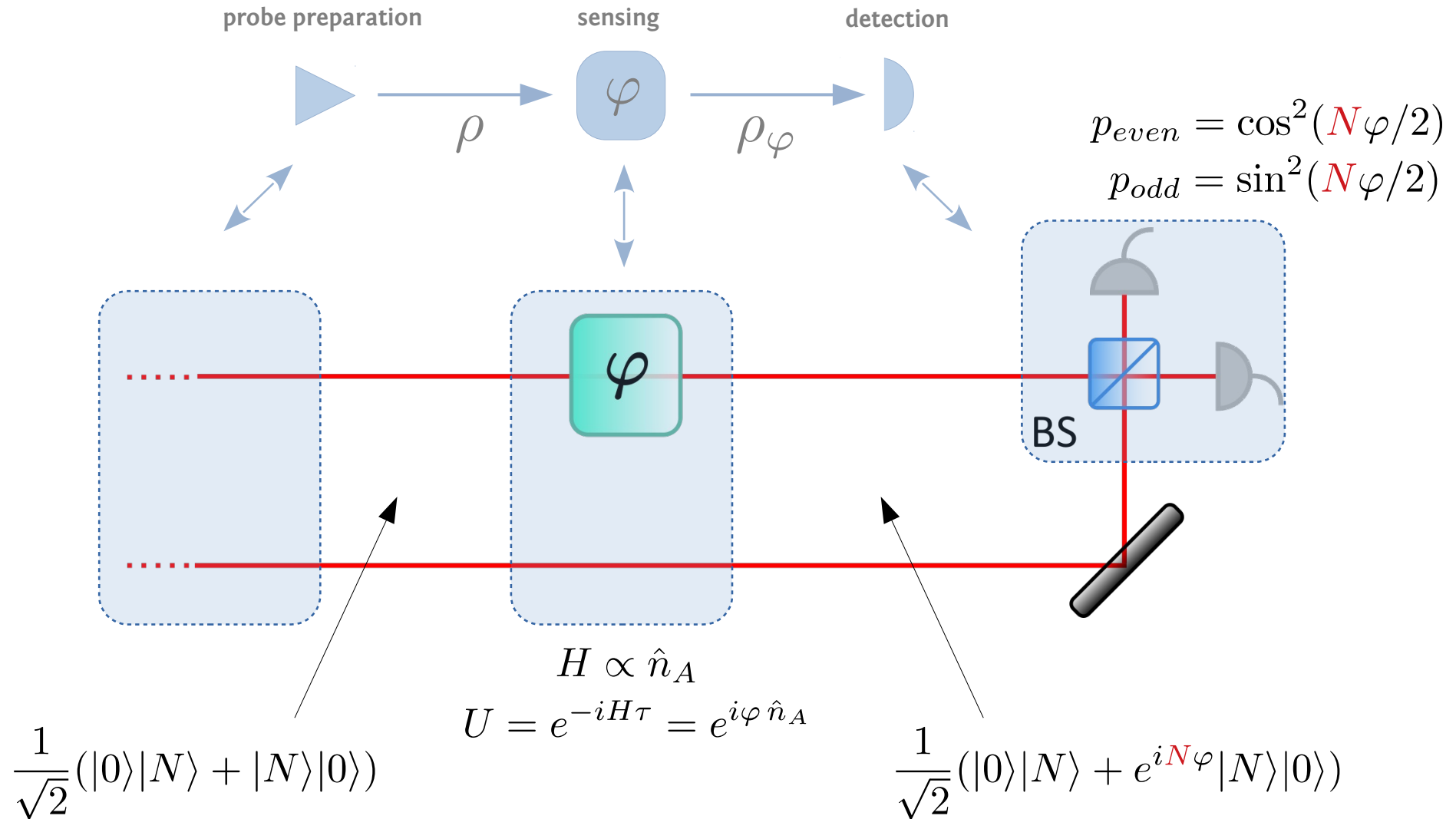


Quantum strategy

$\mathcal{F}_\varphi = N^2$
repeat ν times



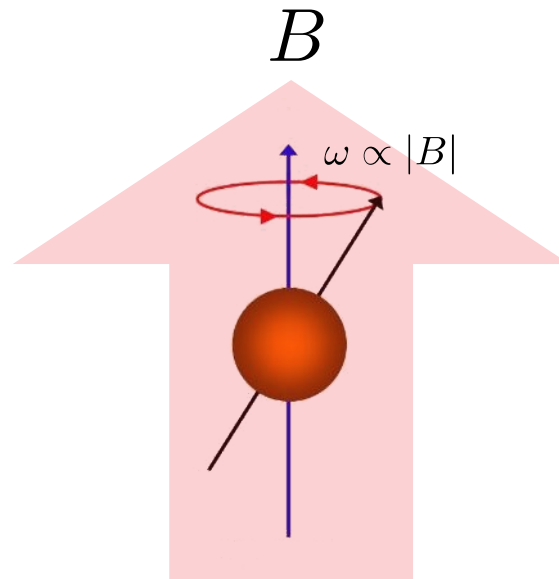
Achieving the Heisenberg limit – optical phase estimation



Quantum strategy

$$\mathcal{F}_\varphi = N^2 \quad \text{repeat } \nu \text{ times} \quad \longrightarrow \quad \Delta^2 \varphi = \frac{1}{\nu N^2}$$

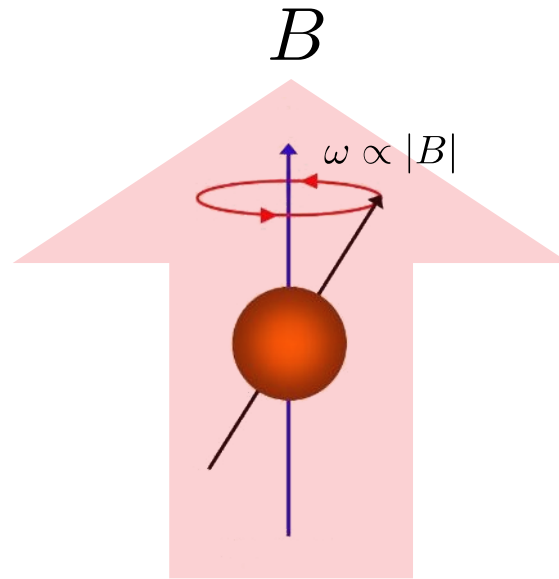
Achieving the Heisenberg limit - magnetometry



$$H = \frac{1}{2}\omega\hat{\sigma}_z$$

Achieving the Heisenberg limit - magnetometry

$$\frac{1}{\sqrt{2}}(|\uparrow\rangle + |\downarrow\rangle)$$



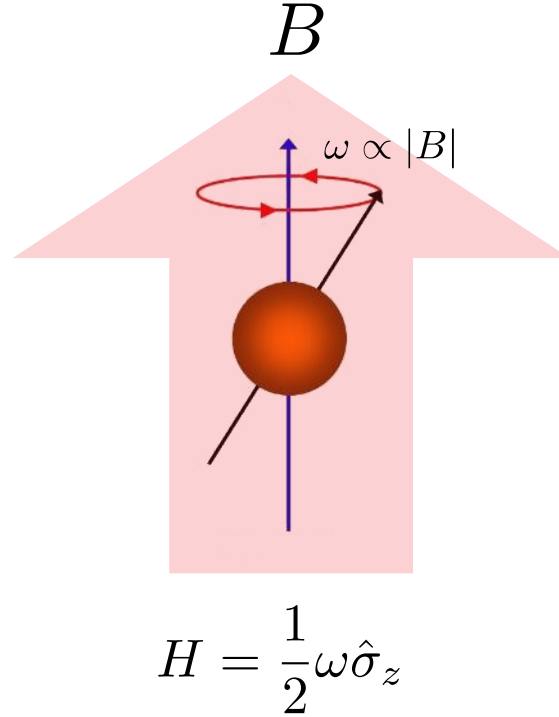
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Achieving the Heisenberg limit - magnetometry

$$\frac{1}{\sqrt{2}}(|\uparrow\rangle + |\downarrow\rangle)$$



$$\frac{1}{\sqrt{2}}(|\uparrow\rangle + e^{-i\omega\tau}|\downarrow\rangle)$$

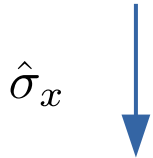


Achieving the Heisenberg limit - magnetometry

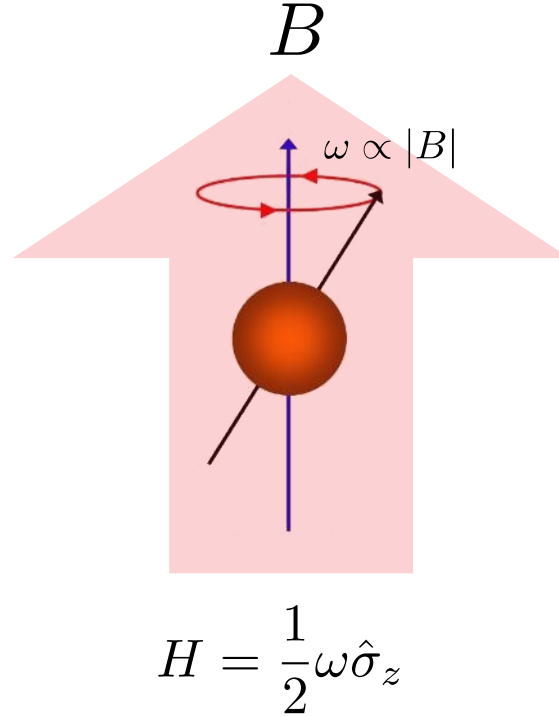
$$\frac{1}{\sqrt{2}}(|\uparrow\rangle + |\downarrow\rangle)$$



$$\frac{1}{\sqrt{2}}(|\uparrow\rangle + e^{-i\omega\tau}|\downarrow\rangle)$$



$$P_+ = \cos^2(\omega\tau/2)$$

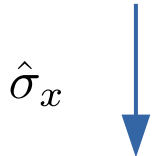


Achieving the Heisenberg limit - magnetometry

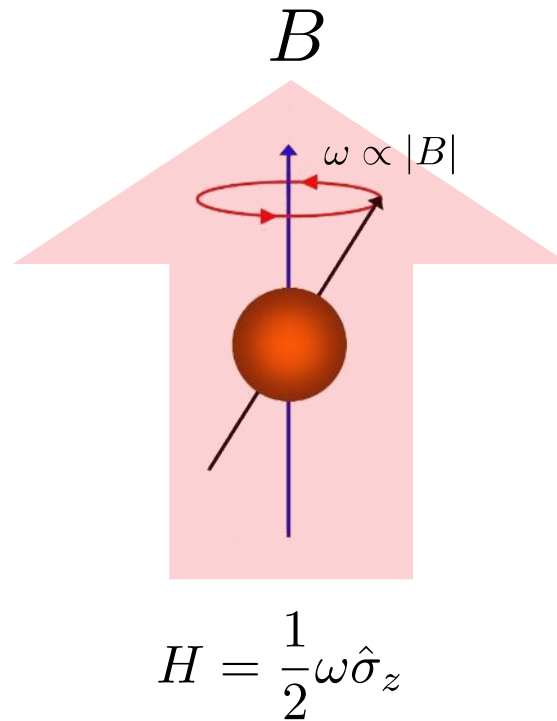
$$\frac{1}{\sqrt{2}}(|\uparrow\rangle + |\downarrow\rangle)$$



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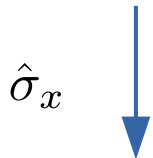
fix total measurement time T
use N particles $N(T/\tau)$ times

Achieving the Heisenberg limit - magnetometry

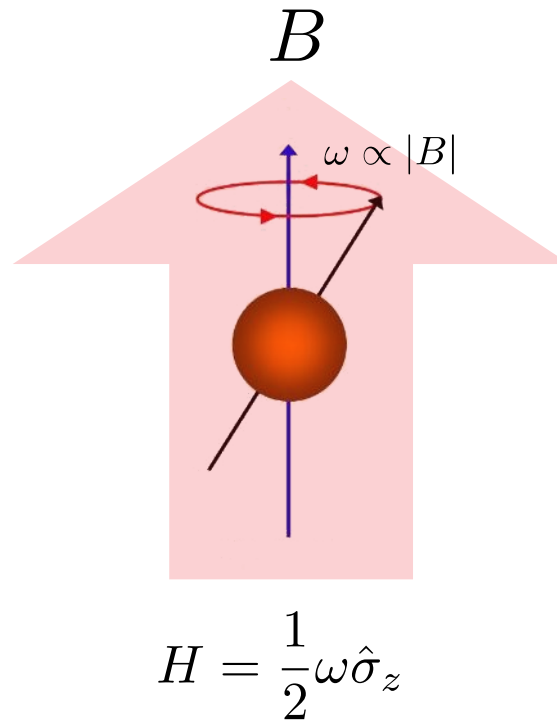
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fix total measurement time T
use N particles $N(T/\tau)$ times

Classical strategy

$$\mathcal{F}_\omega = \tau^2$$

$$\Delta^2\omega = \frac{1}{N\tau^2(T/\tau)} = \frac{1}{N\tau T}$$

Achieving the Heisenberg limit - magnetometry

$$\frac{1}{\sqrt{2}}(|\uparrow\rangle + |\downarrow\rangle)$$

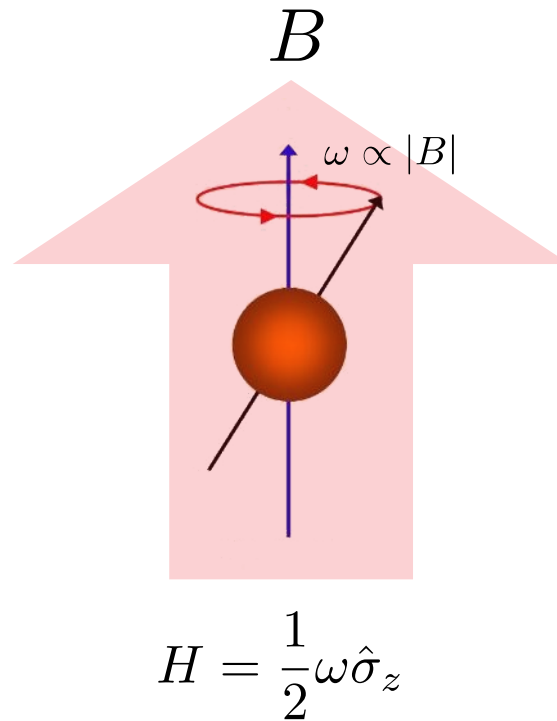


$$\frac{1}{\sqrt{2}}(|\uparrow\rangle + e^{-i\omega\tau}|\downarrow\rangle)$$



$$\hat{\sigma}_x$$

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$$\mathcal{F}_\omega = \tau^2$$

$$\Delta^2\omega = \frac{1}{N\tau^2(T/\tau)} = \frac{1}{N\tau T}$$

Optimal to maximise interaction time

Achieving the Heisenberg limit - magnetometry

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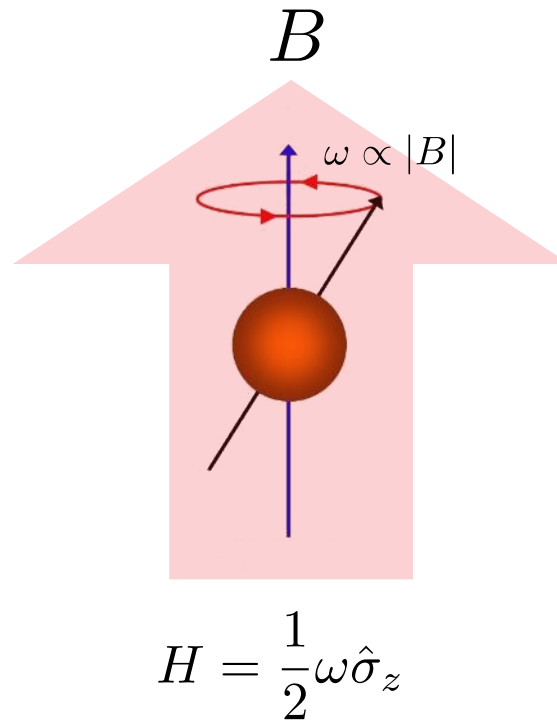


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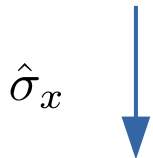
Optimal to maximise interaction time
- not true with noise.

Achieving the Heisenberg limit - magnetometry

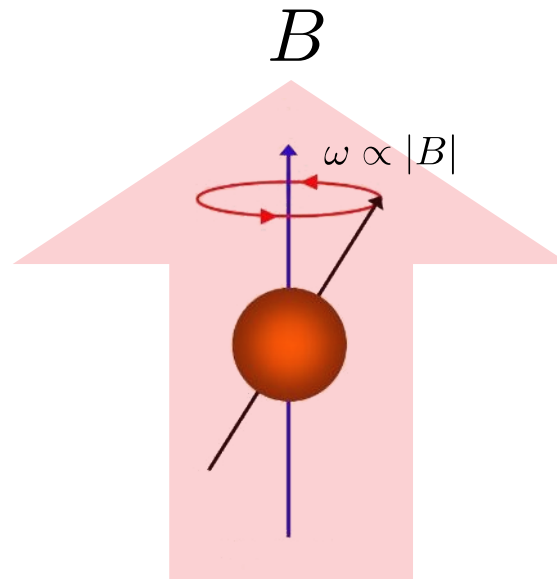
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Achieving the Heisenberg limit - magnetometry

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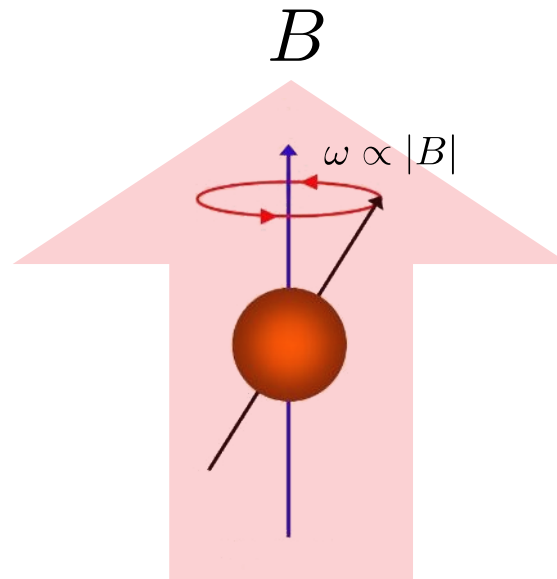


$$\frac{1}{\sqrt{2}}(|\uparrow\rangle + e^{-i\omega\tau}|\downarrow\rangle)$$



$$\hat{\sigma}_x$$

$$P_+ = \cos^2(\omega\tau/2)$$



$$H = \frac{1}{2}\omega \sum_{k=1}^N \hat{\sigma}_z^{(k)}$$

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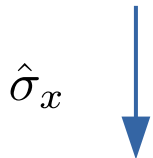
Achieving the Heisenberg limit - magnetometry

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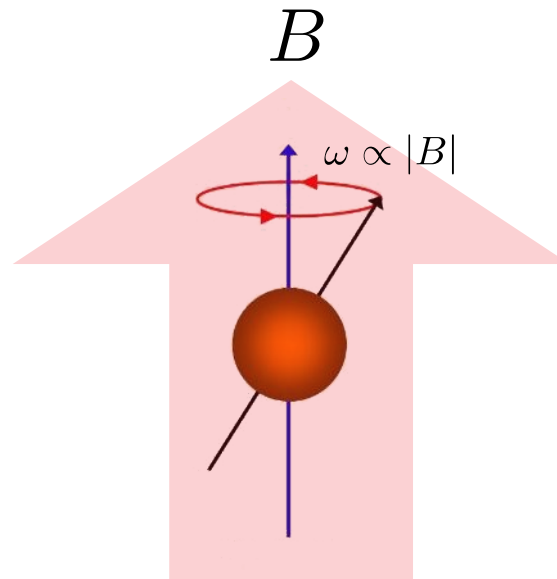
$$\frac{1}{\sqrt{2}}(|\uparrow, \dots, \uparrow\rangle + |\downarrow, \dots, \downarrow\rangle)$$



$$\frac{1}{\sqrt{2}}(|\uparrow\rangle + e^{-i\omega\tau}|\downarrow\rangle)$$



$$P_+ = \cos^2(\omega\tau/2)$$



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Achieving the Heisenberg limit - magnetometry

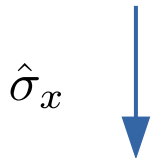
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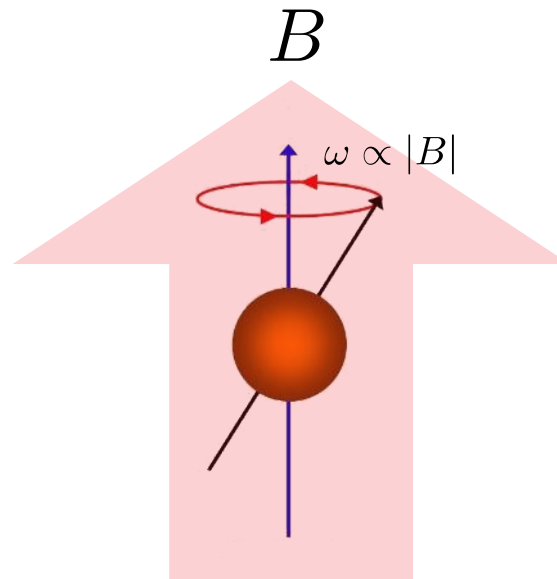


$$\frac{1}{\sqrt{2}}(|\uparrow\rangle + e^{-i\omega\tau}|\downarrow\rangle)$$

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$$H = \frac{1}{2}\omega \sum_{k=1}^N \hat{\sigma}_z^{(k)}$$

fix total measurement time T
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Optimal to maximise interaction time
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Achieving the Heisenberg limit - magnetometry

$$\frac{1}{\sqrt{2}}(|\uparrow\rangle + |\downarrow\rangle)$$

$$\frac{1}{\sqrt{2}}(|\uparrow, \dots, \uparrow\rangle + |\downarrow, \dots, \downarrow\rangle)$$

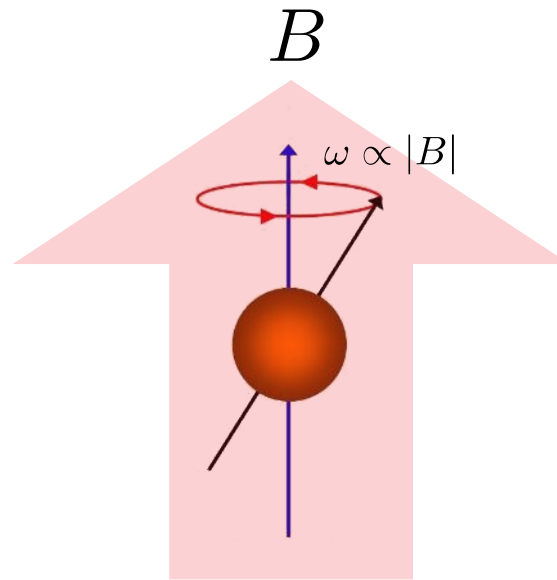


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$$\frac{1}{\sqrt{2}}(|\uparrow, \dots, \uparrow\rangle + e^{-i\mathbf{N}\omega\tau}|\downarrow, \dots, \downarrow\rangle)$$

$$\hat{\sigma}_x \quad \sigma_x^{\otimes N}$$

$$P_+ = \cos^2(\omega\tau/2)$$



$$H = \frac{1}{2}\omega \sum_{k=1}^N \hat{\sigma}_z^{(k)}$$

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Optimal to maximise interaction time
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Achieving the Heisenberg limit - magnetometry

fix total measurement time T
use N particles $N(T/\tau)$ times

$$\frac{1}{\sqrt{2}}(|\uparrow\rangle + |\downarrow\rangle)$$

$$\frac{1}{\sqrt{2}}(|\uparrow, \dots, \uparrow\rangle + |\downarrow, \dots, \downarrow\rangle)$$



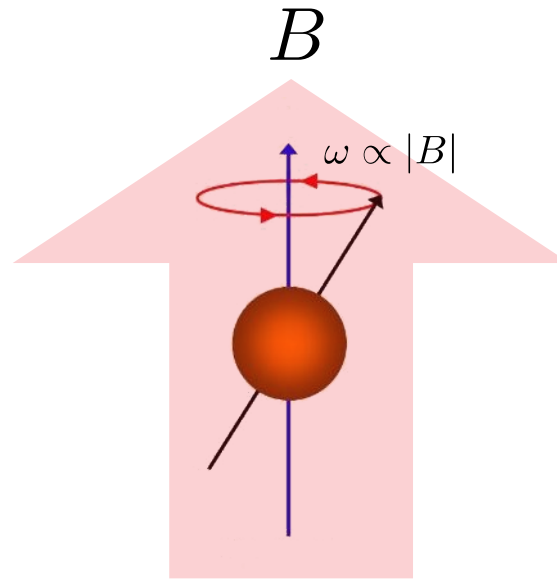
$$\frac{1}{\sqrt{2}}(|\uparrow\rangle + e^{-i\omega\tau}|\downarrow\rangle)$$

$$\frac{1}{\sqrt{2}}(|\uparrow, \dots, \uparrow\rangle + e^{-i\textcolor{red}{N}\omega\tau}|\downarrow, \dots, \downarrow\rangle)$$

$$\hat{\sigma}_x \quad \downarrow \quad \sigma_x^{\otimes N}$$

$$P_+ = \cos^2(\omega\tau/2)$$

$$P_+ = \cos^2(\textcolor{red}{N}\omega\tau/2)$$



$$H = \frac{1}{2}\omega \sum_{k=1}^N \hat{\sigma}_z^{(k)}$$

Classical strategy

$$\mathcal{F}_\omega = \tau^2$$

$$\Delta^2\omega = \frac{1}{N\tau^2(T/\tau)} = \frac{1}{N\tau T}$$

Optimal to maximise interaction time
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Achieving the Heisenberg limit - magnetometry

fix total measurement time T
use N particles $N(T/\tau)$ times

$$\frac{1}{\sqrt{2}}(|\uparrow\rangle + |\downarrow\rangle)$$

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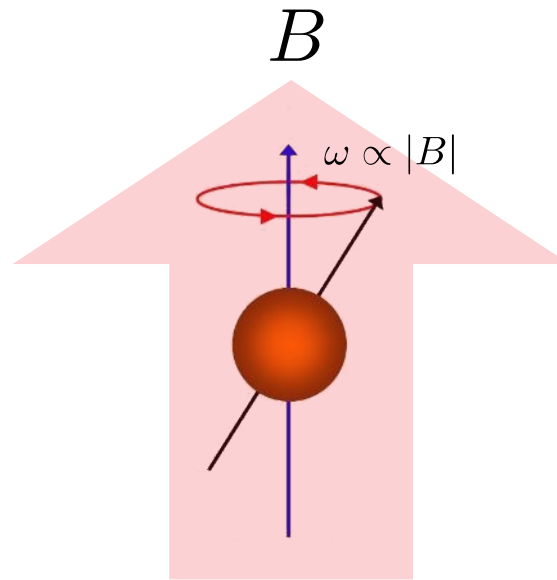
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$$\frac{1}{\sqrt{2}}(|\uparrow, \dots, \uparrow\rangle + e^{-iN\omega\tau}|\downarrow, \dots, \downarrow\rangle)$$

$$\hat{\sigma}_x \quad \downarrow \quad \sigma_x^{\otimes N}$$

$$P_+ = \cos^2(\omega\tau/2)$$

$$P_+ = \cos^2(N\omega\tau/2)$$



$$H = \frac{1}{2}\omega \sum_{k=1}^N \hat{\sigma}_z^{(k)}$$

Classical strategy

$$\mathcal{F}_\omega = \tau^2$$

$$\Delta^2\omega = \frac{1}{N\tau^2(T/\tau)} = \frac{1}{N\tau T}$$

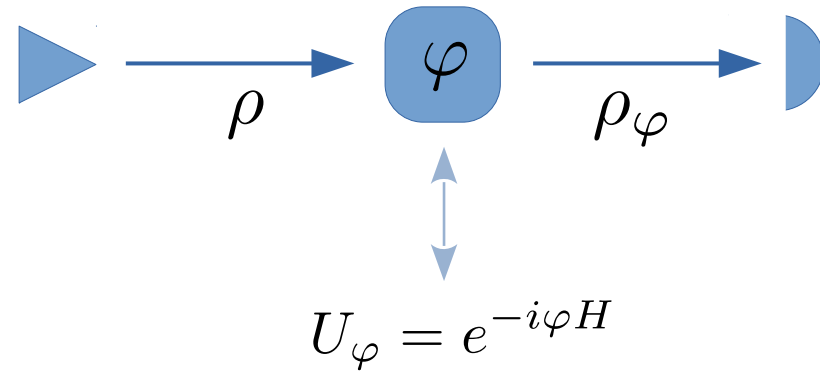
Optimal to maximise interaction time
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Quantum strategy

$$\mathcal{F}_\omega = N^2\tau^2$$

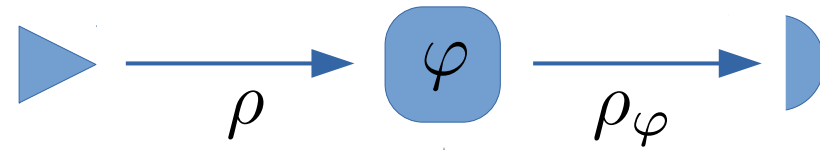
$$\Delta^2\omega = \frac{1}{N^2\tau^2(T/\tau)} = \frac{1}{N^2\tau T}$$

Achieving the Heisenberg Limit – general noiseless case



Unitary encoding of parameter.

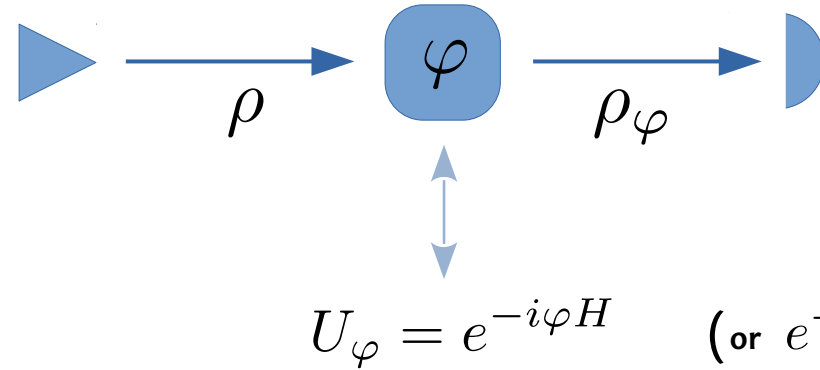
Achieving the Heisenberg Limit – general noiseless case



Unitary encoding of parameter.

$$U_\varphi = e^{-i\varphi H} \quad (\text{or } e^{-i\omega t H})$$

Achieving the Heisenberg Limit – general noiseless case

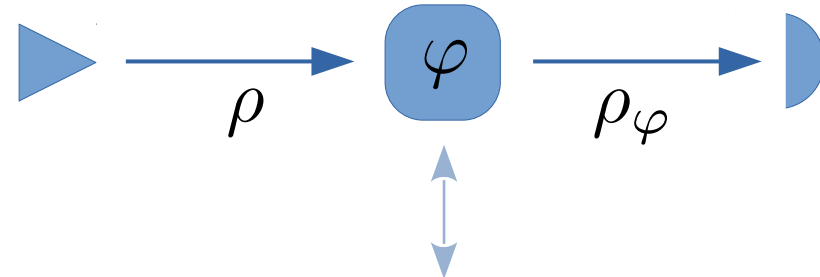


Unitary encoding of parameter.

Pure states are optimal.

QFI convex $\rightarrow$ maximised on pure states.

Achieving the Heisenberg Limit – general noiseless case



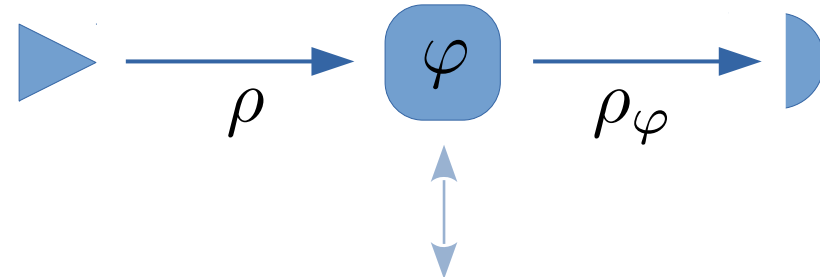
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Pure states are optimal.

QFI convex $\rightarrow$ maximised on pure states.

Achieving the Heisenberg Limit – general noiseless case



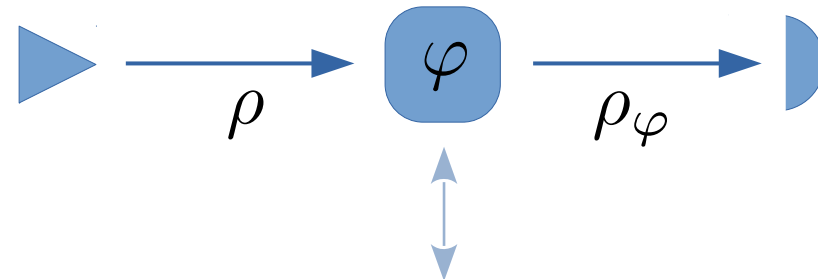
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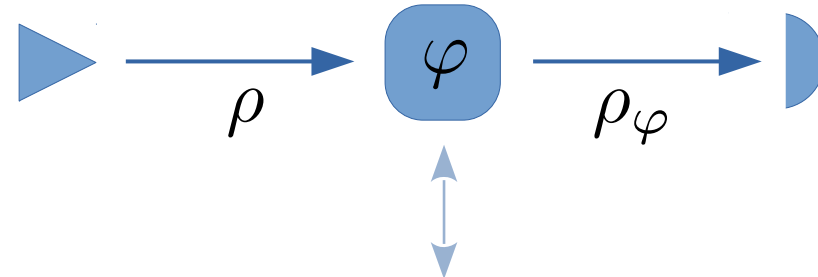
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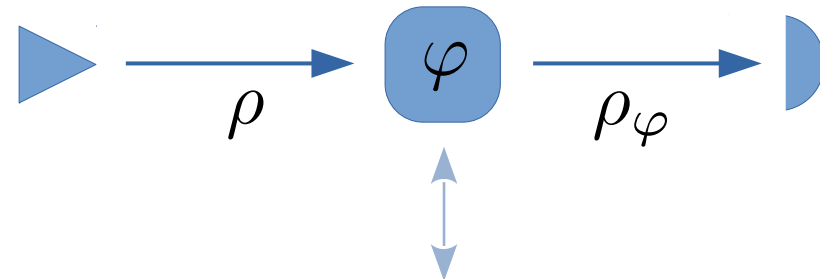
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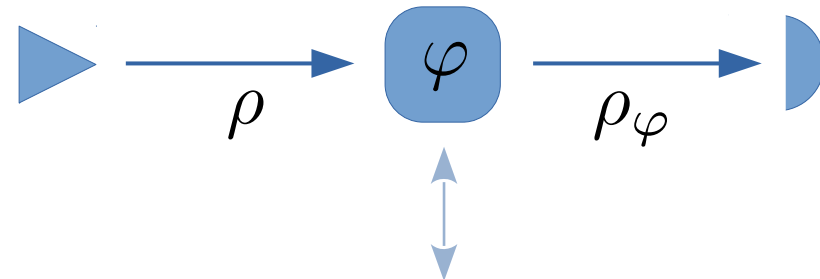
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$$|\psi\rangle = \frac{1}{\sqrt{2}}(|\lambda_{min}\rangle + |\lambda_{max}\rangle) \quad \rightarrow \quad \Delta^2 \varphi = \frac{1}{\nu N^2 (\lambda_{max} - \lambda_{min})}$$

10 MIN BREAK



Precision scaling with noise

Noise might be...

- Fluctuating background fields.
- Interferometer instability.
- Laser noise (power, phase,...)
- Imperfect state preparation.
- Imperfect detectors.
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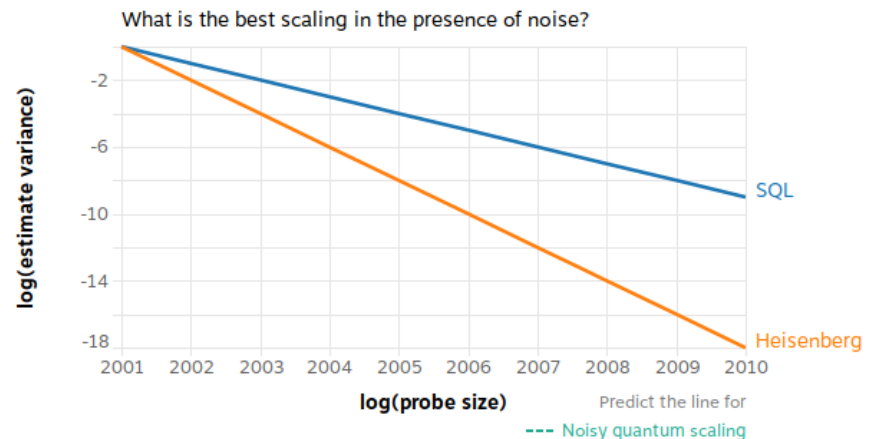
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What do you think the best scaling with noise is?

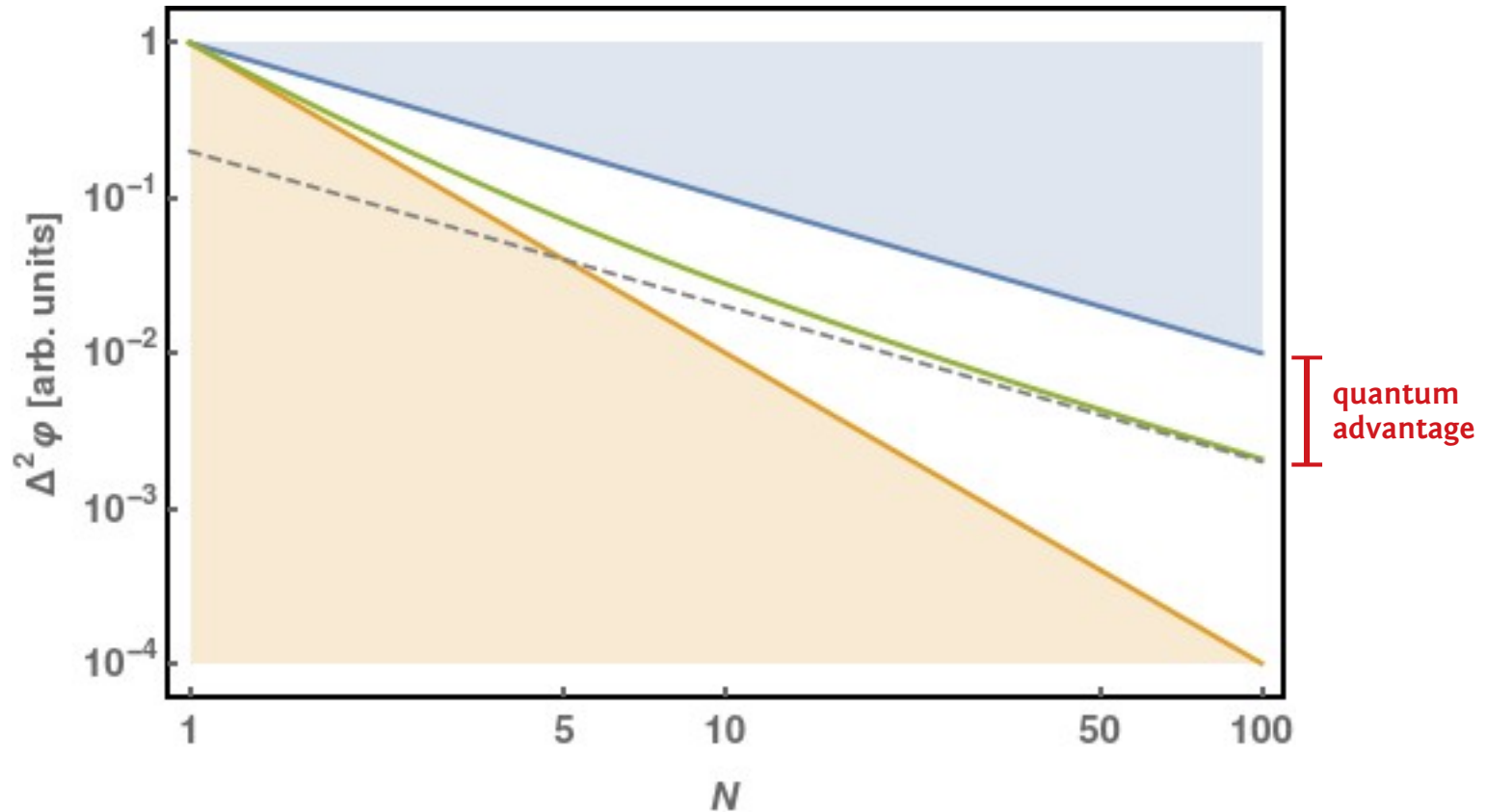
jonatanbohrbrask.dk/scaling2



No-go result for scaling advantage

At most a constant factor for any full-rank channel

'full rank' ~ no subspace is completely free of decoherence



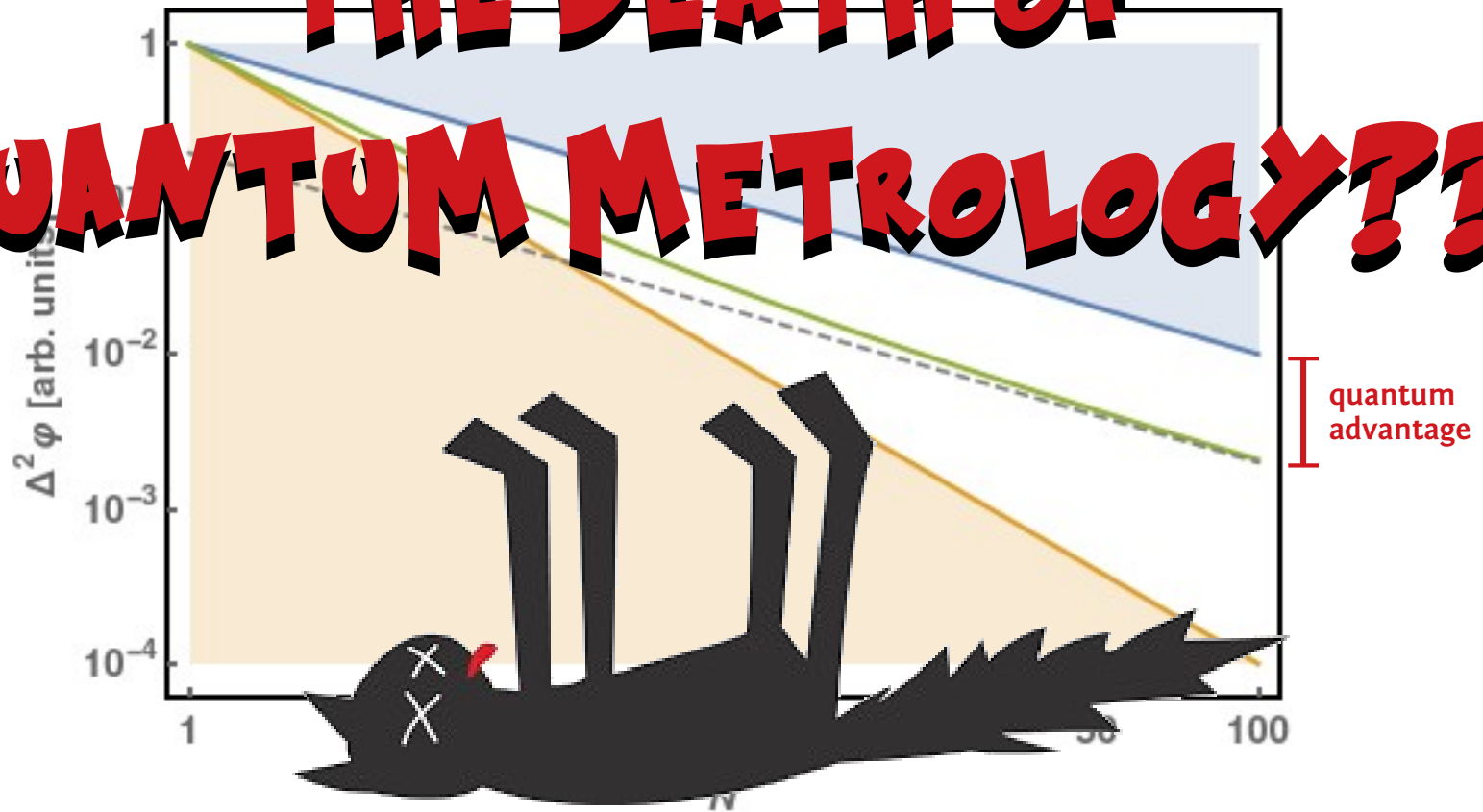
$$(\Delta^2 \phi)_{\text{quantum}} = \kappa (\Delta^2 \phi)_{\text{SQL}} \quad \text{asymptotically}$$

No-go result for scaling advantage

At most a constant factor for any full-rank channel

'full rank' ~ no subspace is completely killed here

THE DEATH OF QUANTUM METROLOGY??



$$(\Delta^2 \phi)_{\text{quantum}} = \kappa (\Delta^2 \phi)_{\text{SQL}} \quad \text{asymptotically}$$

No...



No...



The finite-size regime is important.

No...



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Scaling advantages are possible for directional noise.

No...



The finite-size regime is important.

Scaling advantages are possible for directional noise.

The Heisenberg limit is attainable for certain noise using error correcting codes

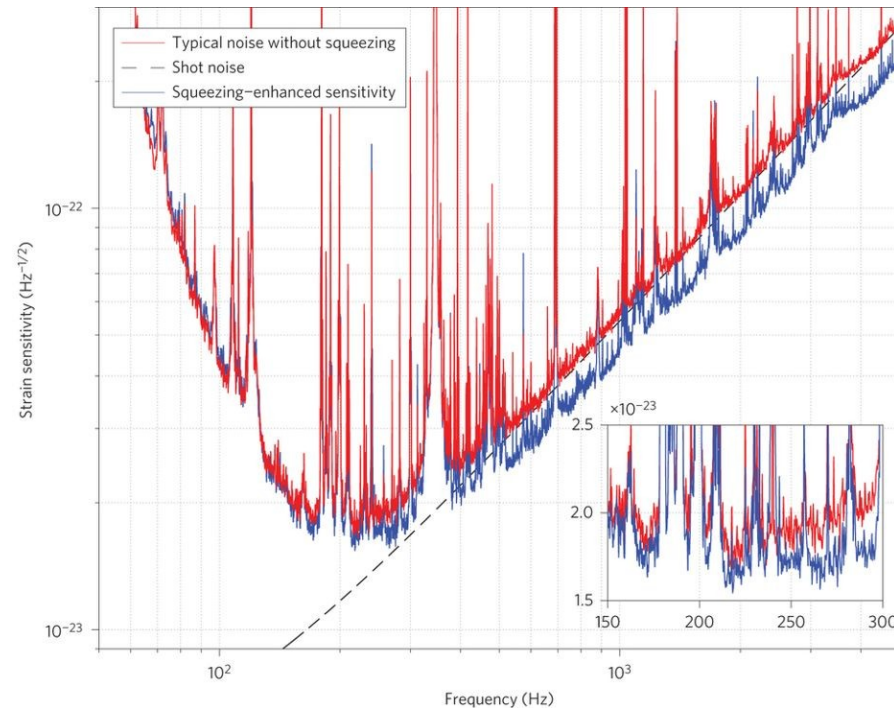
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LIGO: Cannot increase light power much further.

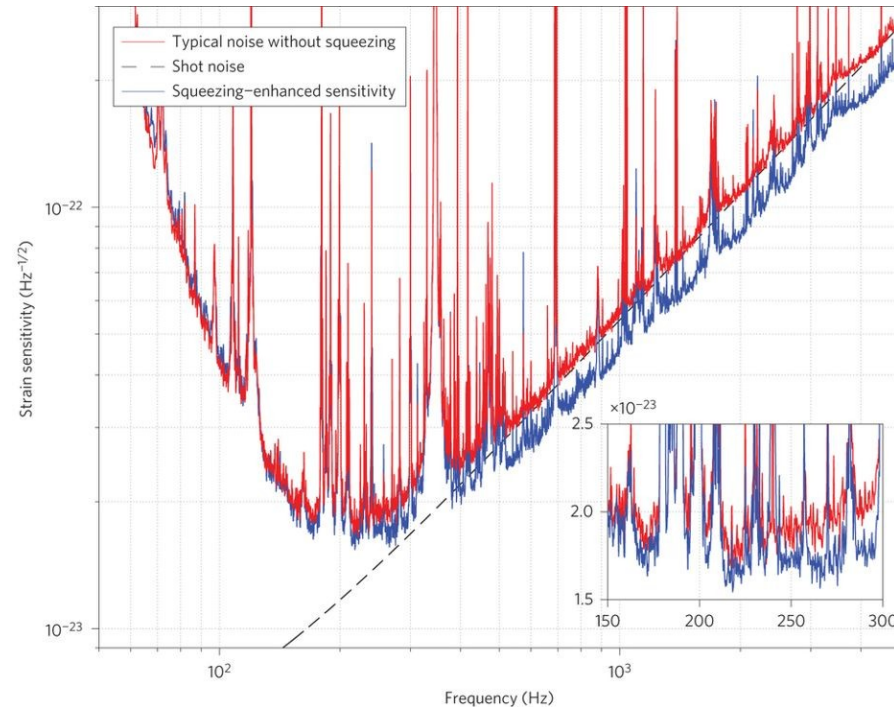


Aasi et al., *Nature Photonics* 7, 613 (2013)

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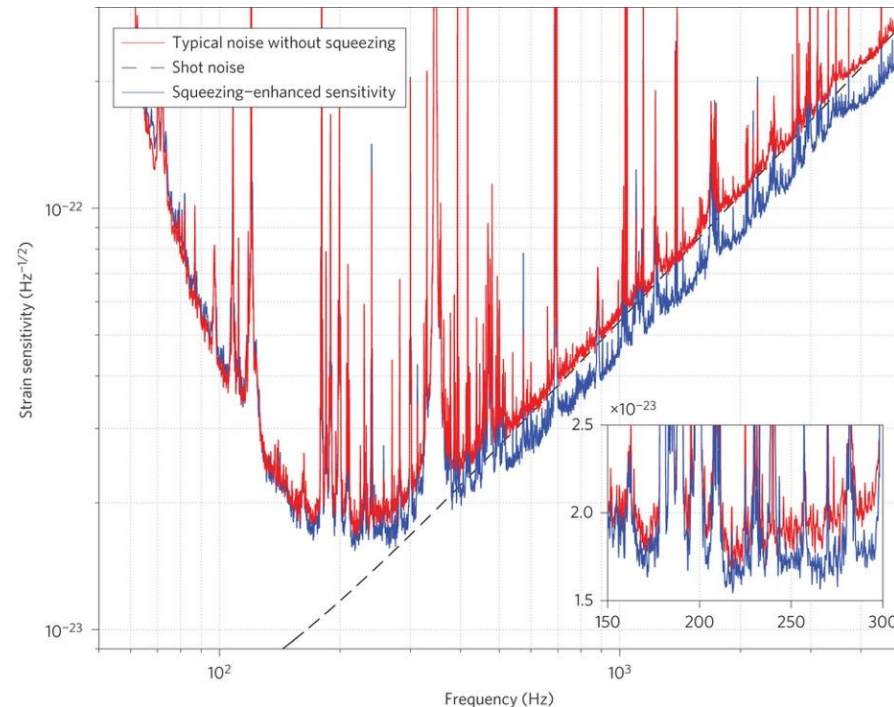
Even moderately better precision may mean significant advances

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Can enable smaller / faster sensors – better precision without increasing probe size

Important e.g. in biological applications.

Quantum scaling advantage for directional noise

Markovian noise described by master equation

$$\dot{\rho} = -i\omega[H, \rho] - \sum_m \gamma_m \left(L_m \rho L_m^\dagger - \frac{1}{2} \{L_m^\dagger L_m, \rho\} \right)$$

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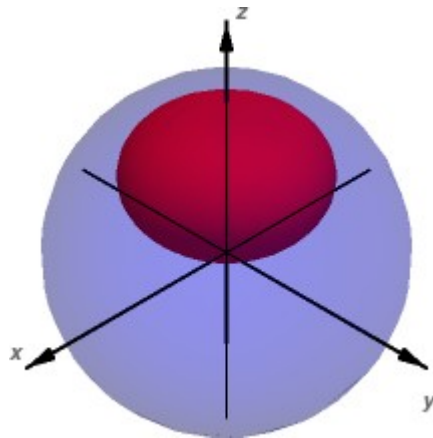
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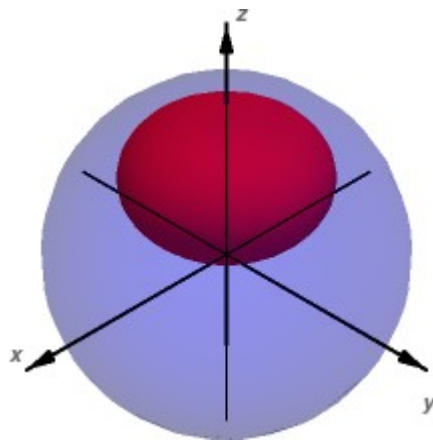
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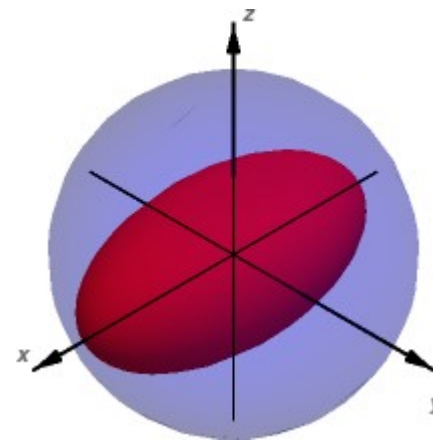
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general



directional

$$\gamma_x \neq 0 \\ \gamma_z, \gamma_y = 0$$

Optimise interaction time for each N

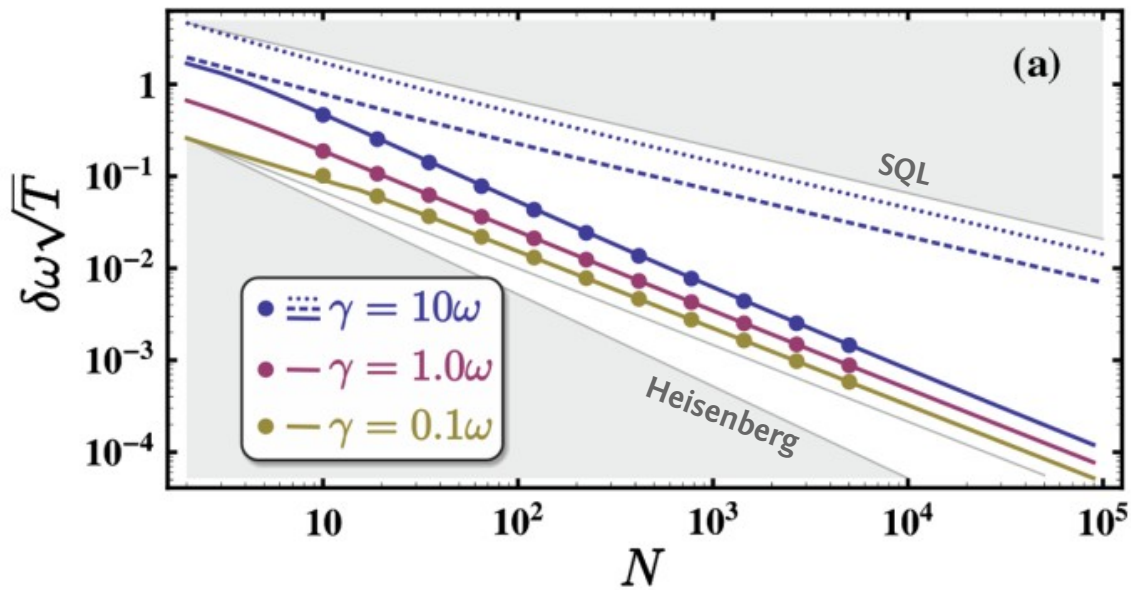
Frequency estimation under perpendicular noise

$$\tau_{opt} \sim N^{-1/3} \quad \rightarrow \quad \Delta^2 \omega \sim \frac{1}{N^{5/3}}$$

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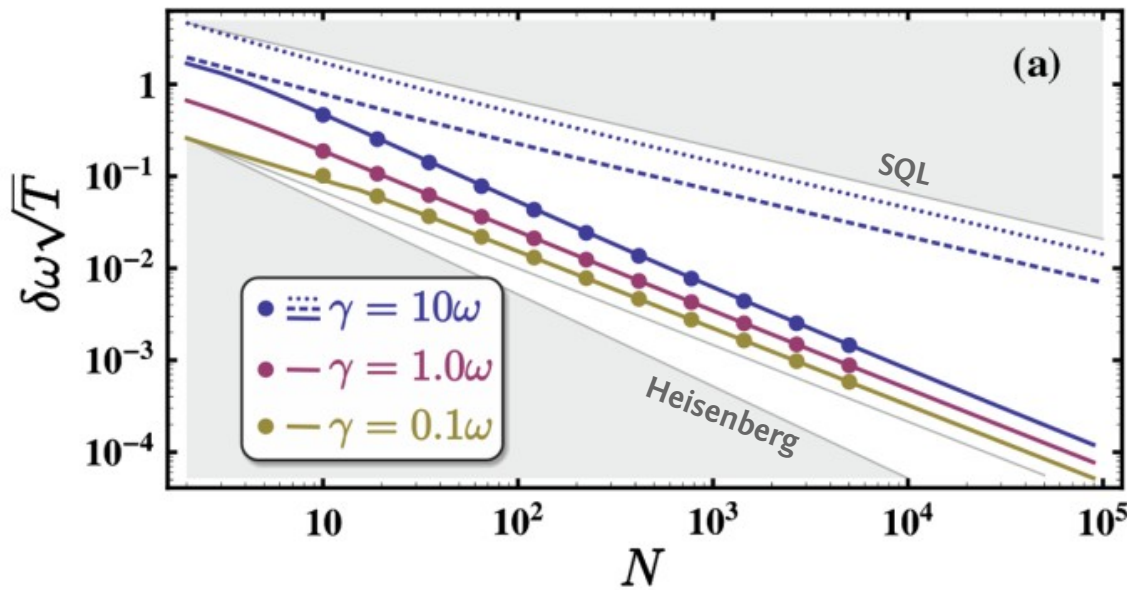
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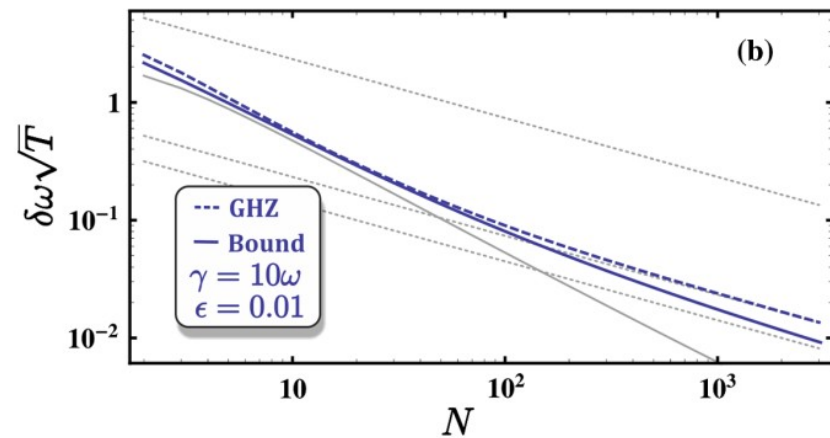
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noise not perfectly parallel / interaction time lower bounded
→ constant factor asymptotic scaling

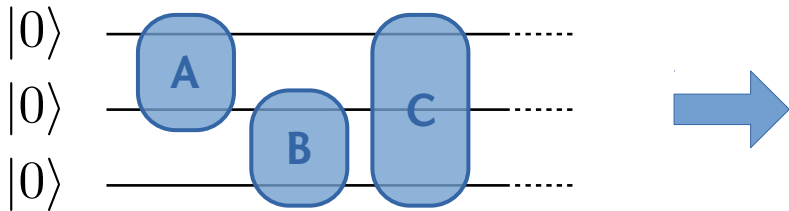
...but cross over can be at large N .



Heisenberg scaling with error correction

Error correcting codes

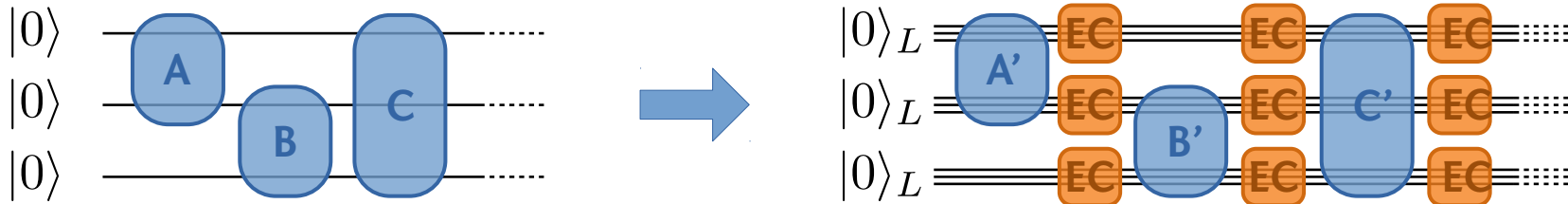
Encode logical qubits in multiple noisy physical qubits



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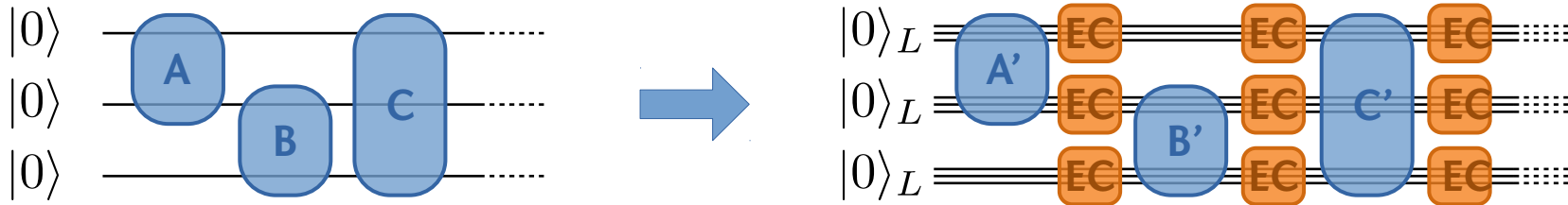
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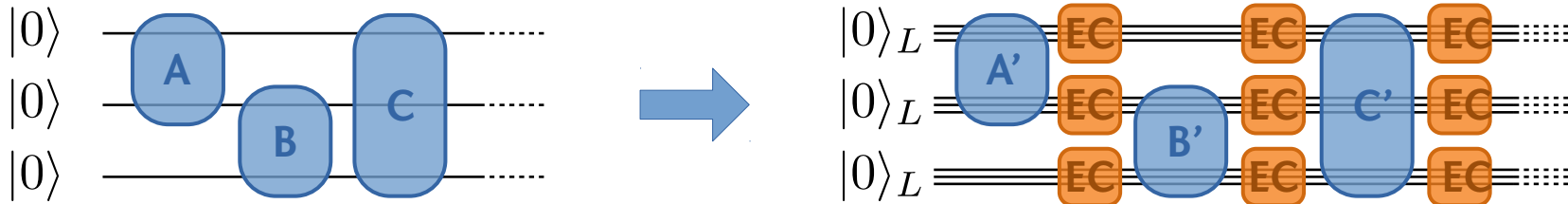
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$$|0\rangle_L = |0\rangle|0\rangle|0\rangle$$

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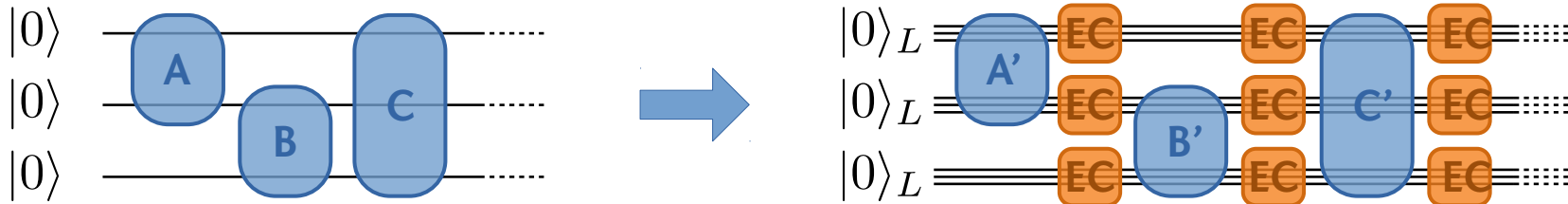
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$$|0\rangle_L = |0\rangle|0\rangle|0\rangle \xrightarrow{\hat{\sigma}_x^{(2)}} |0\rangle|1\rangle|0\rangle$$

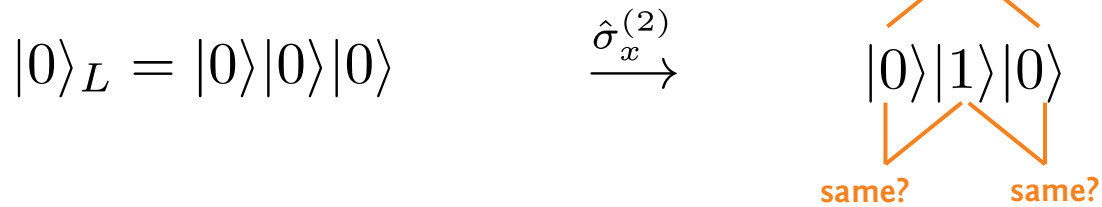
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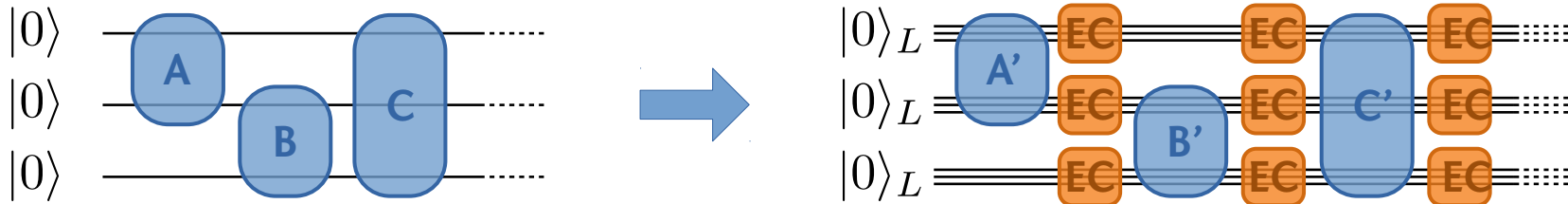
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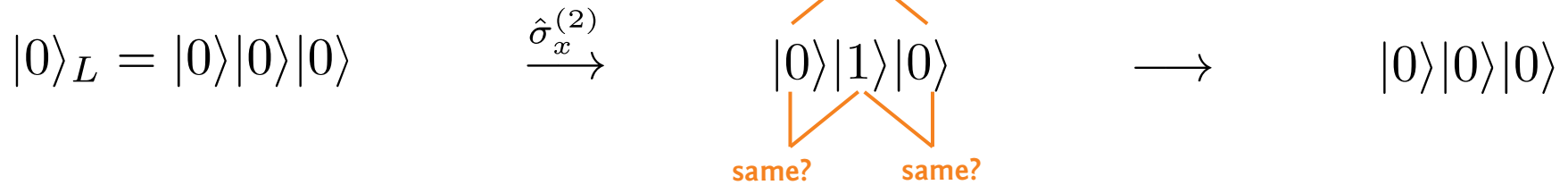
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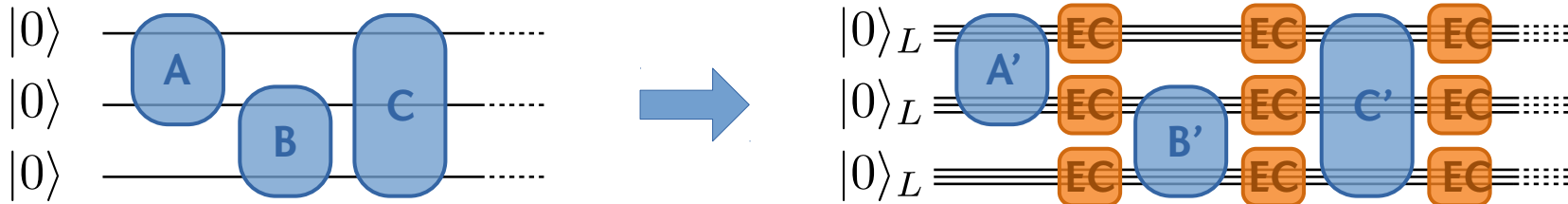
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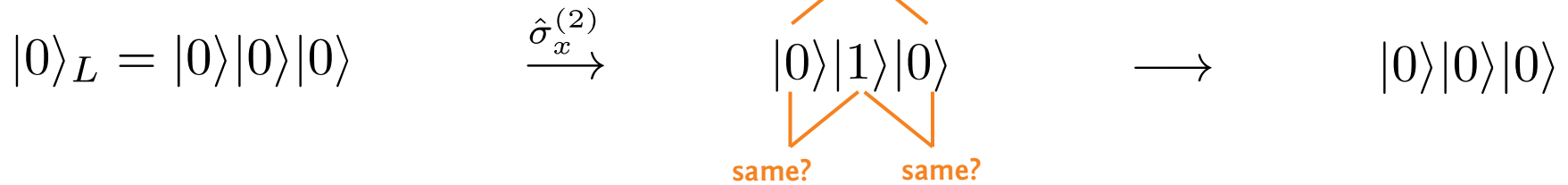
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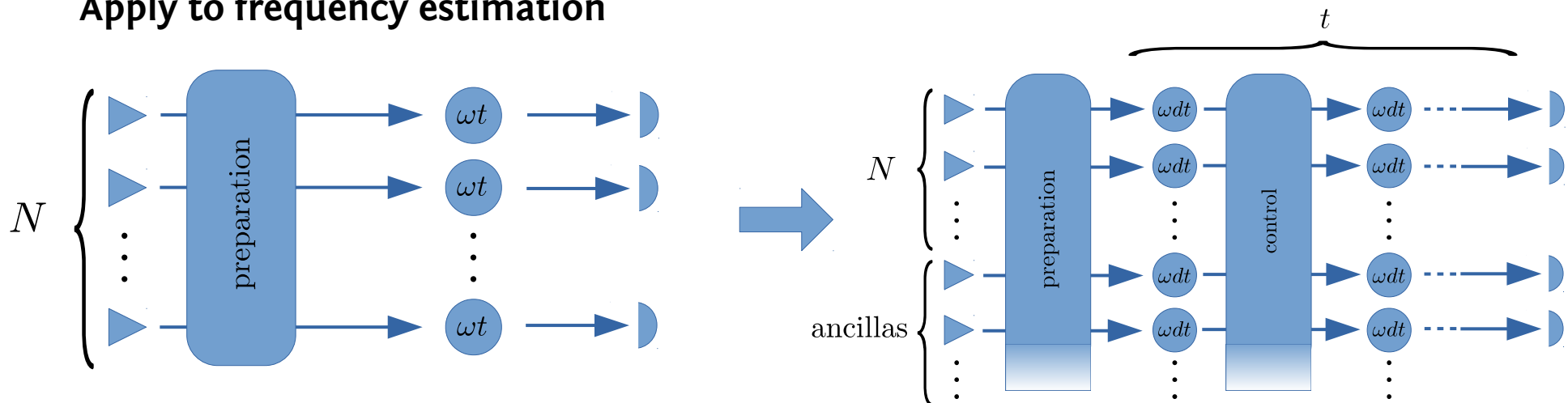
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Apply to frequency estimation



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Hamiltonian not in Lindblad span**

$$H \notin \text{span}\{L_k\}$$

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For qubits: a single Lindblad operator not parallel to Hamiltonian

$$H = \frac{1}{2} \hat{\sigma}_z$$

$$L_x = \hat{\sigma}_x$$

Some literature

Quantum metrology

- <https://science.sciencemag.org/content/306/5700/1330.abstract>
- <https://journals.aps.org/prl/abstract/10.1103/PhysRevLett.96.010401>
- <https://www.nature.com/articles/nphoton.2011.35>
- <https://arxiv.org/abs/1807.11882>

No-go results

- <https://iopscience.iop.org/article/10.1088/1751-8113/41/25/255304>
- <https://www.nature.com/articles/nphys1958>
- <https://www.nature.com/articles/ncomms2067>
- <https://iopscience.iop.org/article/10.1088/1367-2630/15/7/073043>

Directional noise

- <https://journals.aps.org/prl/abstract/10.1103/PhysRevLett.111.120401>
- <https://journals.aps.org/prx/abstract/10.1103/PhysRevX.5.031010>

Error correction

- <https://www.nature.com/articles/s41467-017-02510-3>
- <https://journals.aps.org/prl/abstract/10.1103/PhysRevLett.112.150802>
- <https://journals.aps.org/prl/abstract/10.1103/PhysRevLett.112.150801>
- <https://journals.aps.org/prl/abstract/10.1103/PhysRevLett.112.080801>
- <https://quantum-journal.org/papers/q-2017-09-06-27/>